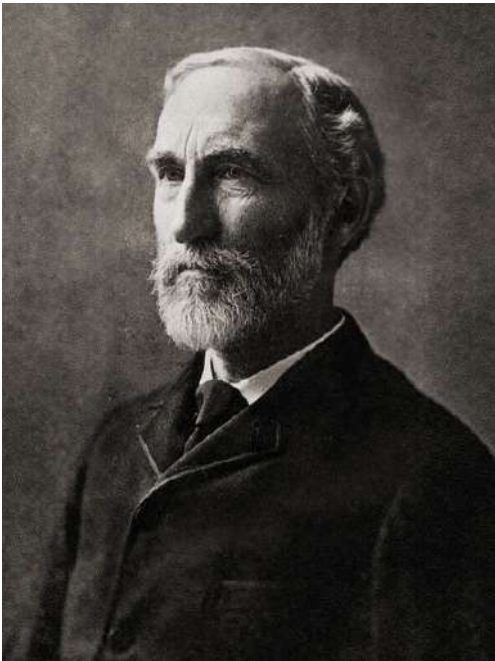


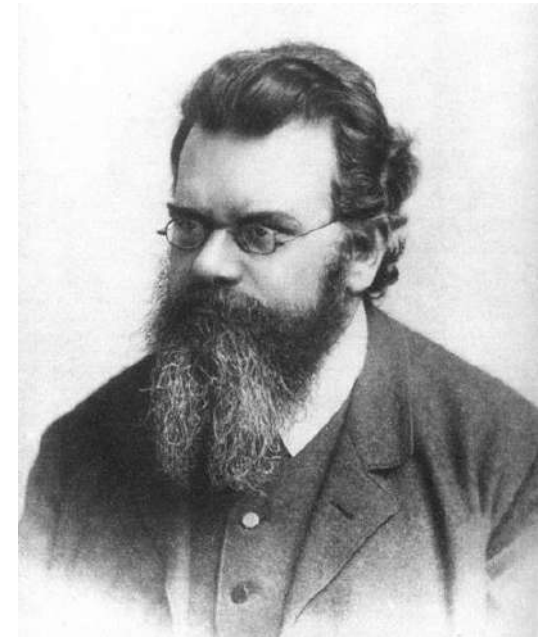
CH 576: Statistical Mechanics



Josiah Willard Gibbs
(1839-1903)



James Clerk Maxwell
(1831-1879)



Ludwig Edward Boltzmann
(1844-1906)

Statistical Mechanics

Classical Mechanics
Quantum Mechanics } Not enough !!

$\{q, p\} \rightsquigarrow$ $\left\{ \begin{array}{l} \text{Newtonian Mechanics} \rightarrow \\ \text{Lagrangian} \rightarrow \\ \text{Hamiltonian} \rightarrow \end{array} \right.$
 \rightarrow they are equivalent

Wavefunction $\psi(q, t) \rightsquigarrow$ All the information of the system

Probabilistic



$\psi^* \psi \rightarrow$ prob. density

$$i\hbar \frac{\partial \psi(q, t)}{\partial t} = \hat{H} \psi(q, t) \rightarrow$$

\downarrow if it has no explicit time dependence

$$\psi(q, t) = e^{-\frac{i\hat{H}t}{\hbar}} \psi(q, 0)$$

$q \equiv$ position co-ordinate

$$\dot{q} = \frac{dq}{dt} = v$$

$$\ddot{q} = \frac{d^2q}{dt^2}$$

$$m\ddot{q} = F$$

\rightarrow Second order differential equation

Initial position $q(0)$
Initial velocity $\dot{q}(0)$ } $q(t)$
 \downarrow

"DETERMINISTIC"

$$F = -V'(q)$$

$$\parallel -\frac{dV}{dq}$$

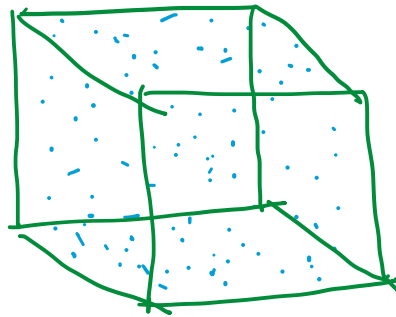
Statistical Mechanics gives a prescription to calculate macroscopic (thermodynamic) quantities, starting from a microscopic description of the system.

Calculate \rightarrow Pressure, Entropy, free energy

Microscopic Description

\downarrow
Classical Mechanics

\rightarrow Postulates and Hypothesis



CM

$(q_k, p_k) \rightarrow k^{\text{th}} \text{ molecule / atom}$

$\{q_k, p_k\}$

Classical Mechanics

- Deterministic -

$$F = m\ddot{q}$$

"Newton"

Example: Simple Harmonic Oscillator (SHO)

$$m\ddot{q} = -k(q - q_0)$$

$$(q - q_0) = \xi$$

$$\dot{q} = \dot{\xi}$$

$$\ddot{q} = \ddot{\xi}$$

$$m \frac{d^2 \xi}{dt^2} = -k \xi$$

$$\frac{d^2 \xi}{dt^2} + \frac{k}{m} \xi = 0$$

$$\xi(t) = A \sin \omega t + B \cos \omega t$$

$$\omega = \left(\frac{k}{m} \right)^{\frac{1}{2}}$$

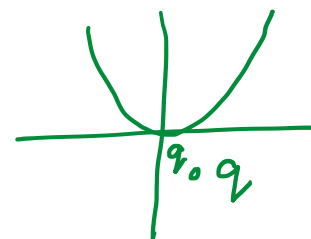
$$2\pi\nu = \left(\frac{k}{m} \right)^{\frac{1}{2}}$$

$$\nu = \frac{1}{2\pi} \sqrt{\frac{k}{m}}$$

$$F = -\frac{\partial V}{\partial q}$$

$$V = \frac{1}{2} k (q - q_0)^2$$

Displacement



Lagrangian Approach: 1 particle

$$K(\dot{x}, \dot{y}, \dot{z}) = \frac{m}{2} (\dot{x}^2 + \dot{y}^2 + \dot{z}^2)$$

↓ ↓
Kinetic energy velocities

It is not
a fⁿ of (x, y, z) but
a fⁿ of $(\dot{x}, \dot{y}, \dot{z})$

only a fⁿ of (x, y, z)

$$L(x, y, z, \dot{x}, \dot{y}, \dot{z}) = K(\dot{x}, \dot{y}, \dot{z}) - V(x, y, z)$$

↓
Lagrangian

$$\frac{\partial L}{\partial \dot{x}} = \frac{\partial K}{\partial \dot{x}} = \frac{m}{2} 2 \dot{x} = m \dot{x}$$

$$\frac{\partial L}{\partial x} = -\frac{\partial V}{\partial x}$$

$$\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{x}} \right) = m \ddot{x}$$

$$m \ddot{x} = F = -\frac{\partial V}{\partial x}$$

L. eqⁿ of
motion

$$\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{x}} \right) = \left(\frac{\partial L}{\partial x} \right)$$

$$\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}_j} \right) = \left(\frac{\partial L}{\partial q_j} \right)$$

$j = 1, 2, 3, \dots$

Hamiltonian Approach:

$$H(p_1, p_2, p_3, q_1, q_2, q_3) = \sum_{j=1}^3 p_j \dot{q}_j - L(q_1, q_2, q_3, \dot{q}_1, \dot{q}_2, \dot{q}_3)$$

↓ "Hamiltonian"

$$\frac{\partial L}{\partial \dot{q}_j} = m \dot{q}_j = p_j = \frac{\partial K}{\partial \dot{q}_j}$$

$$= 2a_j \dot{q}_j$$

$$H = \sum_j p_j \dot{q}_j - L$$

$$= \sum_j 2a_j \dot{q}_j \dot{q}_j - L$$

$$= 2 \sum_j a_j \dot{q}_j^2 - L$$

$$= 2K - L = 2K - (K - V) = K + V$$

$$= \underline{\text{Total energy}}$$

$$K = \sum_{j=1}^{3N} a_j(q_1, q_2, q_3, \dots) \dot{q}_j^2$$

↪ K. E. for N particle system.

$$m \dot{q}_j$$

$$a_j = \frac{m}{2}$$

→ depends on position but not on velocity

$$\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}_j} \right) = \left(\frac{\partial L}{\partial q_j} \right)$$

$$p_j = \frac{\partial L}{\partial \dot{q}_j}$$

$$H = \sum_{j=1}^3 p_j \dot{q}_j - L$$

$$dH = d \left(\sum_{j=1}^3 p_j \dot{q}_j \right) - dL$$

$$= \sum_{j=1}^3 p_j d\dot{q}_j + \sum_{j=1}^3 \dot{q}_j dp_j -$$

$$dH = \sum_j \dot{q}_j dp_j - \sum_j p_j d\dot{q}_j$$

$$H = H(p_j, q_j)$$

$$dH = \sum \frac{\partial H}{\partial q_j} dq_j + \sum \frac{\partial H}{\partial p_j} dp_j$$

Comparing

$$\left. \begin{aligned} \frac{\partial H}{\partial p_j} &= \dot{q}_j ; & \frac{\partial H}{\partial q_j} &= -\dot{p}_j \end{aligned} \right\} \text{Hamilton's eq}^n \text{ of motion}$$

$$L = L(q_j, \dot{q}_j)$$

$$dL = \sum_{j=1}^3 \frac{\partial L}{\partial q_j} dq_j + \sum_{j=1}^3 \frac{\partial L}{\partial \dot{q}_j} d\dot{q}_j$$

$$= \sum_j \underbrace{\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}_j} \right)}_{p_j} d\dot{q}_j + \sum p_j d\dot{q}_j$$

$$= \sum p_j d\dot{q}_j + \sum p_j d\dot{q}_j$$

Simple Harmonic Oscillator

$$m\ddot{x} = -kx$$

\searrow
 $= m\dot{v}$
 $= \dot{p}$

$$H = \frac{p^2}{2m} + \frac{1}{2}kx^2$$

$$\frac{\partial H}{\partial p} = \frac{p}{m} = \dot{x}$$

$$\frac{\partial H}{\partial x} = \frac{1}{2}2kx = -(-kx)$$

\parallel
 $- \dot{p}$

$$H(q_i, p_i)$$

$$\frac{dH}{dt} = \underbrace{\sum \left(\frac{\partial H}{\partial q_i} \frac{\partial q_i}{\partial t} \right) + \sum \left(\frac{\partial H}{\partial p_i} \frac{\partial p_i}{\partial t} \right)}_{= \sum -\dot{p}_i \dot{q}_i + \sum \dot{q}_i \dot{p}_i} + \frac{\partial H}{\partial t}$$

\nearrow Partial derivative
 $\searrow = 0$ (no explicit time dependence)

$= 0$
 \rightsquigarrow Hamiltonian is constant of motion
 \hookrightarrow Total energy of the system is conserved

Newton's equation:

$$m\ddot{q} = F = -\frac{dV}{dq}$$

↳ conservative force \Rightarrow Derivative of a potential

Hamiltonian Approach:

$$H(p, q) = p\dot{q} - L(\dot{q}, q)$$

}

$$\frac{\partial H}{\partial p} = \dot{q} \quad \frac{\partial H}{\partial q} = -\dot{p}$$

Hamilton's equation of motion

Lagrangian approach:

Define $L(\dot{q}, q) \equiv K(\dot{q}) - V(q)$

$$\frac{\partial L}{\partial \dot{q}} = \frac{\partial K}{\partial \dot{q}} = m\dot{q}$$

$$\frac{\partial L}{\partial q} = -\frac{\partial V}{\partial q}$$

Thus, $\boxed{\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}} \right) = \frac{\partial L}{\partial q}}$

$m\ddot{q} = F$

Newton's equation
in terms of Lagrangian

Thermodynamics:

Work done by the system

$$W = \int_A^B P dV$$

Heat absorbed by the system

$$Q = \int_A^B dq$$

$$\underbrace{Q - W}_{\text{path functions}} = \underbrace{\Delta E}_{\text{State fn}} \rightarrow \text{Change in internal energy}$$

First law of thermodynamics

Change in entropy $\Delta S = \int_A^B \frac{dq_{\text{rev}}}{T}$ for a reversible process

Second law

for all other processes $\Delta S > \int_A^B \frac{dq}{T} \Rightarrow \boxed{ds > \frac{dq}{T}}$

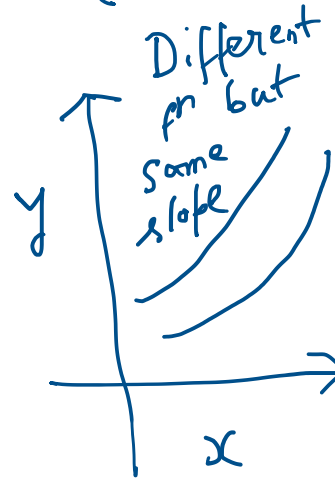
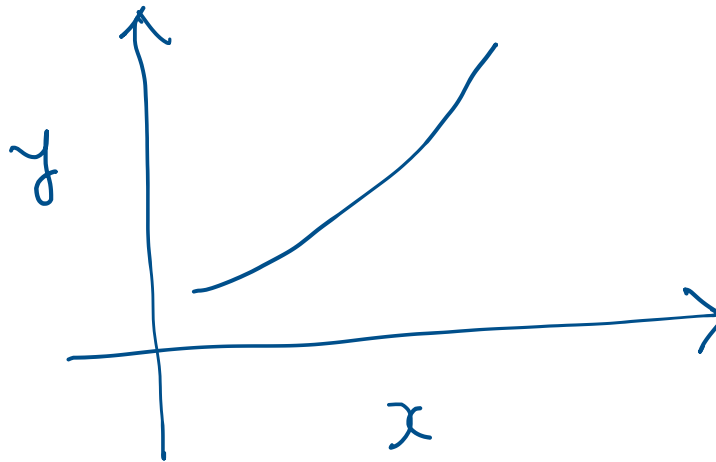
$$S - S_0 = \int_0^T \frac{dq_{rev}}{T} = S \quad \text{as } S_0 = 0 \quad (S \text{ at } T=0 \text{ is zero})$$

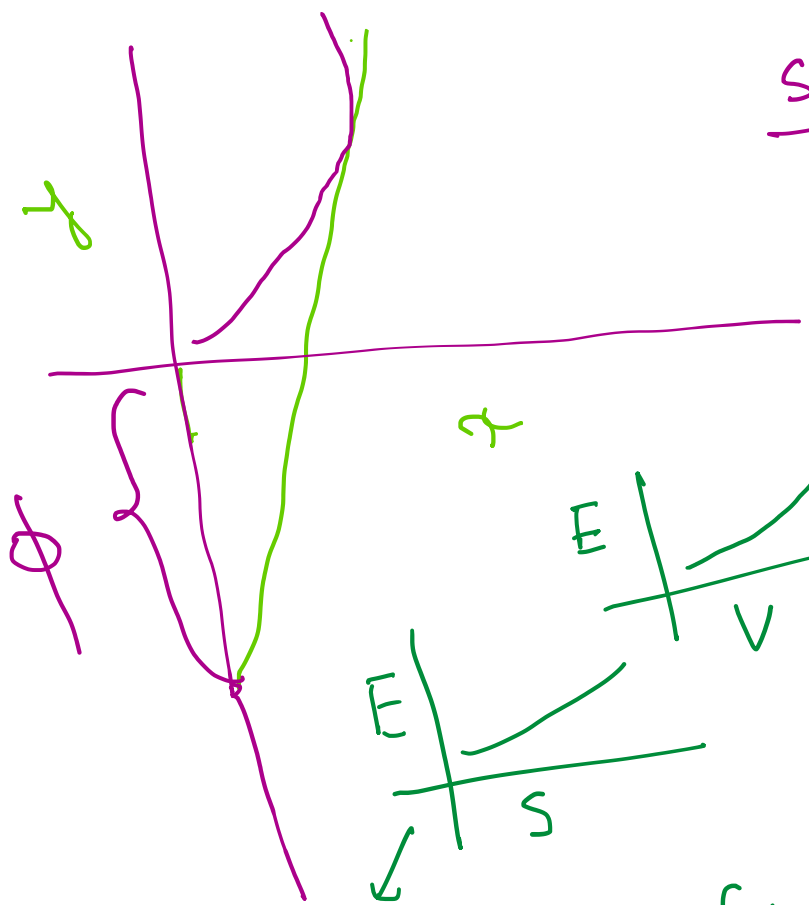
"third law"

$$dE = dq - dw \quad \text{First law}$$

$$dE = T dS - P dV \leadsto (S, V) \text{ are the natural variables for } E \equiv E(S, V)$$

$$\left(\frac{\partial E}{\partial S} \right)_V = T \quad \left(\frac{\partial E}{\partial V} \right)_S = -P$$





Slope:

$$p = \frac{y - \phi}{x - 0} \Rightarrow$$

slope
 (not pressure
 or momentum)

$\phi(p) = y - px$

$$y(x) \Leftrightarrow \phi(p = \frac{dy}{dx})$$

Legendre transformation

$$E(S, V) \Rightarrow f(T, V)$$

$$E(S) \Rightarrow f(T)$$

$$f(T) = E - TS = A(T, V)$$

→ Helmholtz free

$$dE = TdS - P.dV$$

$$\left(\frac{\partial E}{\partial S}\right)_V = T ; \left(\frac{\partial E}{\partial V}\right)_S = -P$$

$$g = E - \left(\frac{\partial E}{\partial V}\right)_S V = E + PV \rightarrow \text{enthalpy}$$

$$H(p, q) = p\dot{q} - L(\dot{q}, q)$$

$$H(p) = p\dot{q} - L(\dot{q})$$

$$f(p) = L(\dot{q}) - \left(\frac{\partial L}{\partial \dot{q}} \right) \dot{q}$$

$$f(p) = L(\dot{q}) - p\dot{q}$$

$$-f(p) = p\dot{q} - L(\dot{q})$$

$$\xrightarrow{\quad} H(p)$$

$$\phi(p) = y - px$$

Not momentum

$$p = \frac{dy}{dx}$$

$$\frac{\partial L}{\partial \dot{q}} = p$$

→ They are related by Legendre transformations

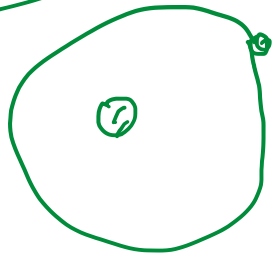
Quantum Mechanics:

Probabilistic

$$\psi(x,t) \quad \psi(q,t)$$
$$\int dx \psi^*(x,t) \psi(x,t) \rightarrow \text{Probability density}$$

↓
Total probability

H-atom

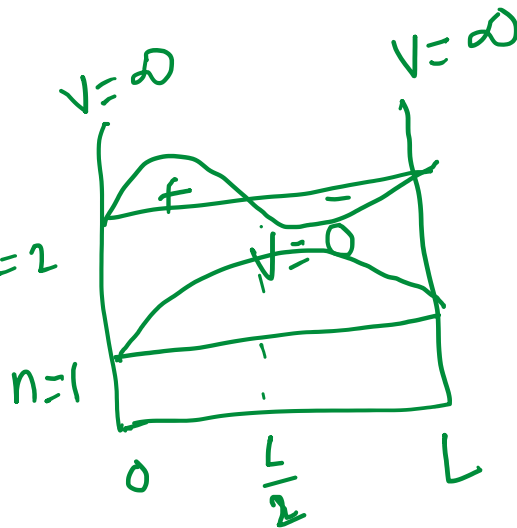


Simple Harmonic Oscillator
(SHO)

Rigid Rotors

Particle in a box: $n=2$

$$E = \frac{n^2 h^2}{8mL^2}$$



Time independent
Schrödinger's equation

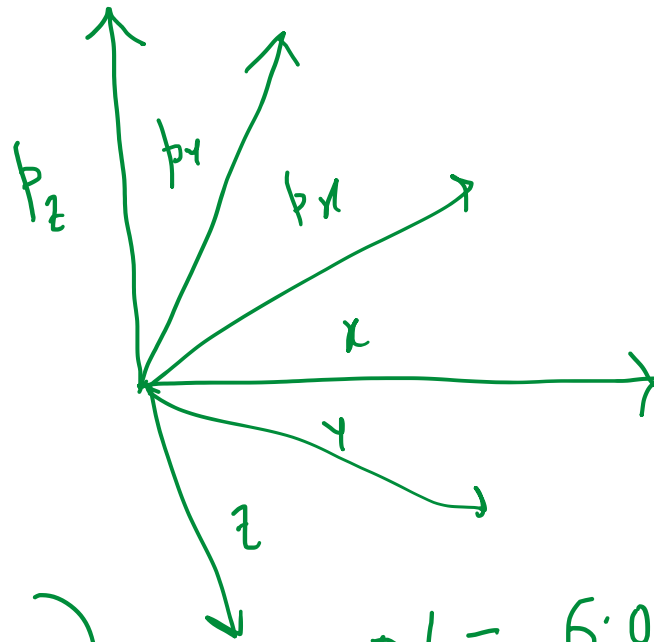
$$-\frac{\hbar^2}{2m} \frac{d^2 \psi}{dx^2} = E \psi$$

$$\hookrightarrow \psi = \sqrt{\frac{2}{L}} \sin\left(\frac{n\pi x}{L}\right)$$

Single particle

Phase Space

(q, p)
↓
 x, y, z
→ p_x, p_y, p_z
6-co-ordinates



Think of N particles

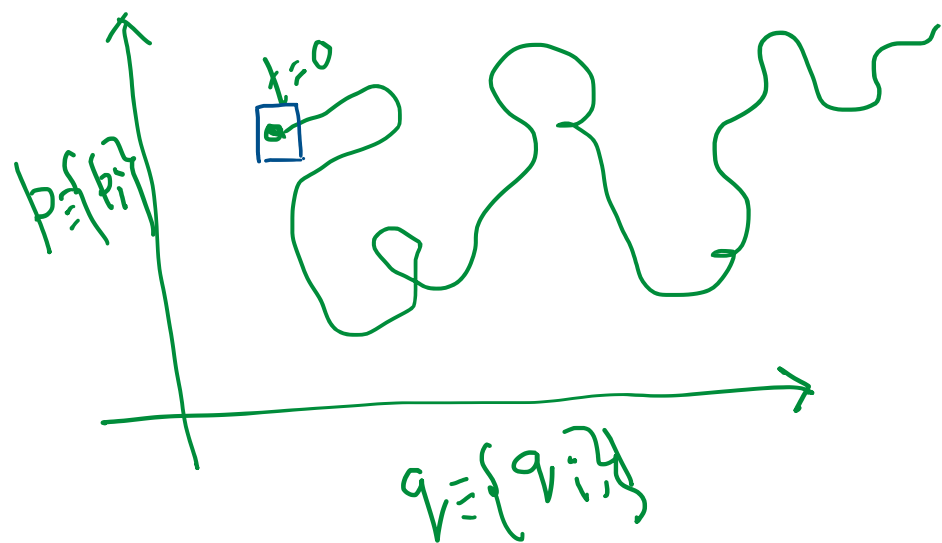
$3N \rightarrow$ position
co-ordinates

$3N \rightarrow$ momentum
co-ordinates

$$N = 6.023 \times 10^{23}$$
$$6N = 36 \dots \times 10^{23}$$

Huge No.
at some given
time

$\{p_i, q_i\}$ or $\{q_i, p_i\} \Rightarrow$ phase space $\rightarrow 6N$ dimensional



Phase Space trajectory

What is rule the system follows in this phase space?

Each of these particles is following Newton's equation

No Computer in this world can store

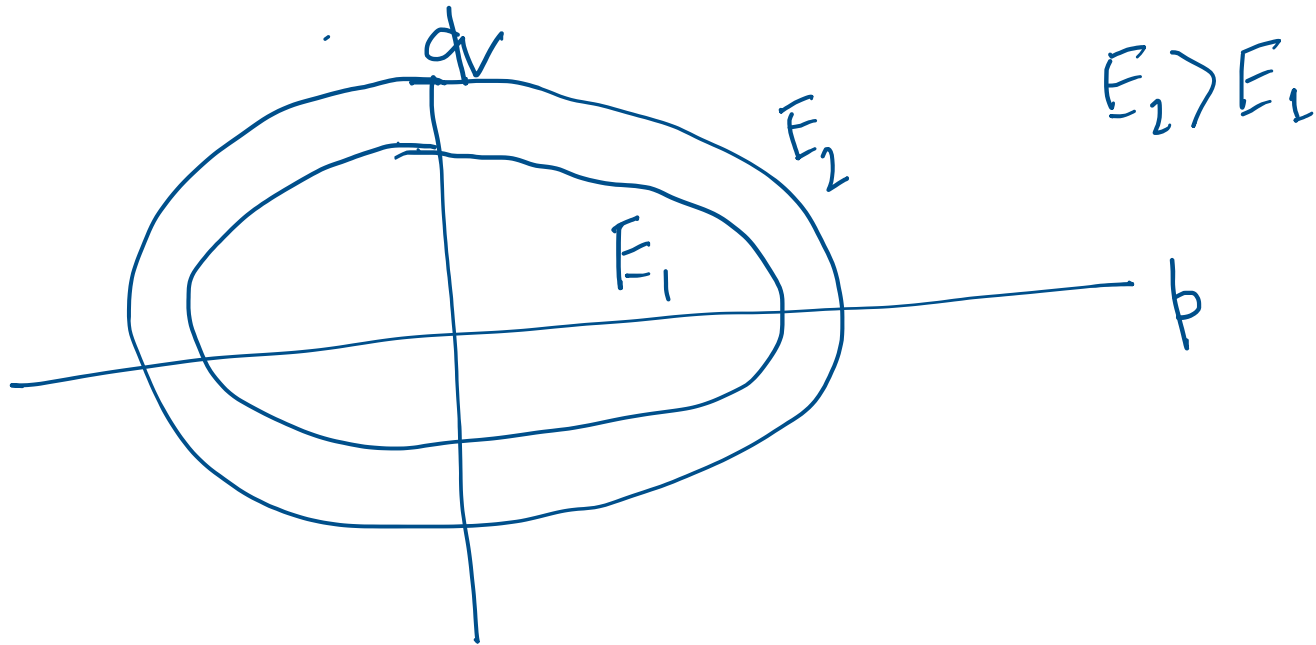
\rightarrow Phase space distribution $f^n \rightarrow f(p_i, q_i, t)$

\rightarrow follows some eqⁿ \rightarrow Liouville's eqⁿ.

Simple Harmonic Oscillator

SHO

$$H = \frac{p^2}{2m} + \frac{1}{2}kq^2 = E_{\text{Total}}$$

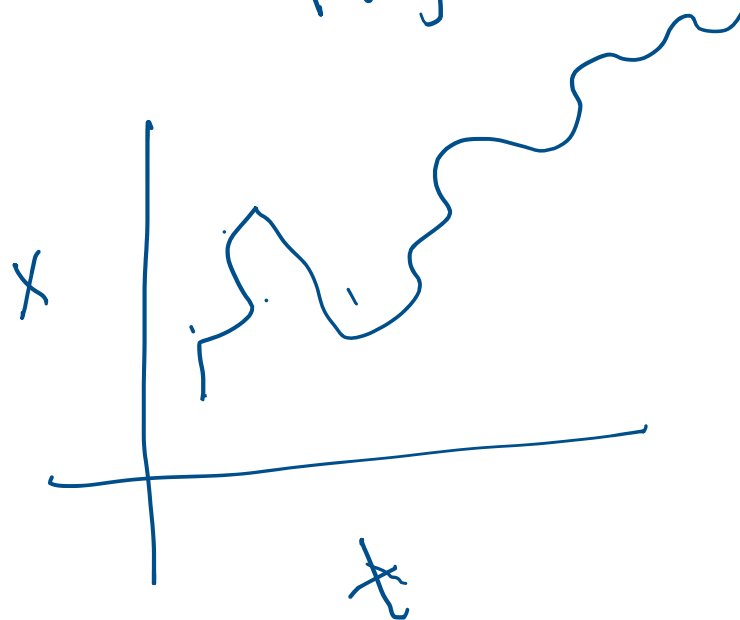


Time Average:

$$\bar{X} = \lim_{\tau \rightarrow \infty} \frac{1}{\tau} \int_{t_0}^{t_0 + \tau} ds X(s)$$

→ time average
over a
long
trajectory

$$\bar{X} = \lim_{t \rightarrow \infty} \frac{1}{t} \int_{t_0}^{t_0 + t} dt' X(t')$$



Ensemble & Ensemble average:

Microstate of a system $\rightarrow (p_i, q_i)$

Macrostate \Rightarrow Many microstates
can correspond
to a single macrostate

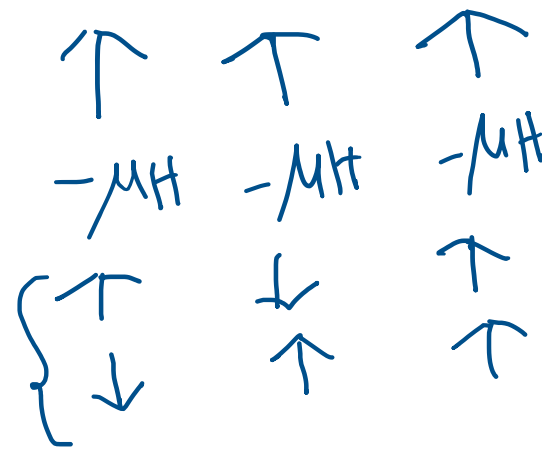


Magnetic system (three tiny magnets)

External magnetic field

$$E = -\vec{\mu} \cdot \vec{H}$$

two microstates \leftarrow



$$E = -3\mu H$$

$$\left. \begin{array}{l} E = -\mu H \\ E = -\mu H \end{array} \right\} \text{Same macrostate}$$

Time Average:

$$\bar{X} = \lim_{\tau \rightarrow \infty} \frac{1}{\tau} \int_{t_0}^{t_0 + \tau} X(s) ds$$

$$\bar{X} = \lim_{\tau \rightarrow \infty} \frac{1}{\tau} \int_0^\tau X(s) ds$$

$X(t)$

First postulate of Statistical Mechanics:

$$\bar{X} = \langle X \rangle$$

Time average = Ensemble average

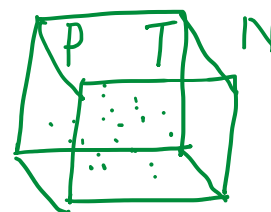
Ensemble Average:

$$\langle X \rangle = \lim_{N \rightarrow \infty} \frac{1}{N} \sum_{i=1}^N p_i X_i$$

Macrostate of a system:

$M(N, P, T)$
 \downarrow
 Macrostate

Variables
State

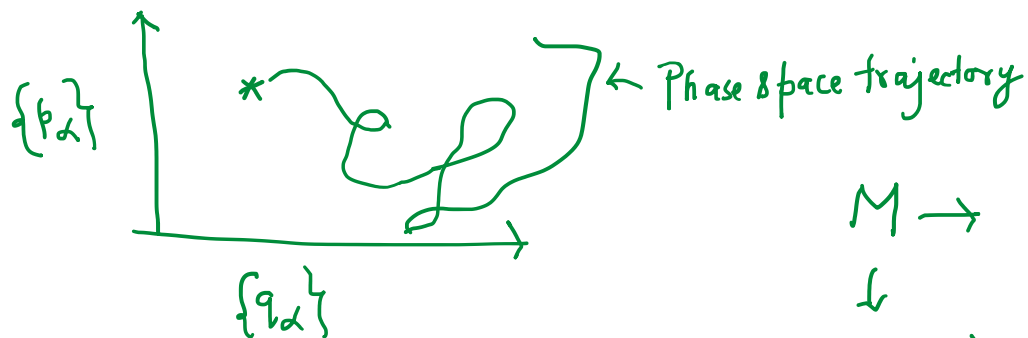
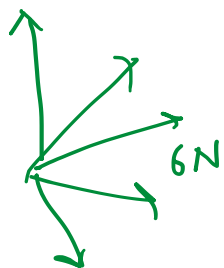


$3N \rightarrow$ position
 $3N \rightarrow$ momenta

 $6N$
 $\sim 10^{24}$

Microstate \rightarrow position and the momentum co-ordinates of all the constituents

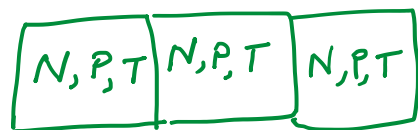
Phase space $\rightarrow 6N$ dimensional space



$M \rightarrow 1$ microstate



another copy of $M \rightarrow 1$ microstate



An ensemble

$\dots dV^0$
Such macro-states

N_M no. of macrostates (mental copies of) $\rightsquigarrow N_M$ no. of microstates

$\int \mathcal{N}^0 \rightarrow$ No. of microstates

$\rightarrow 3N$ momenta co-ordinates

$\mathcal{P}(\underline{p}, \underline{q}) \rightarrow 3N$ position coordinates

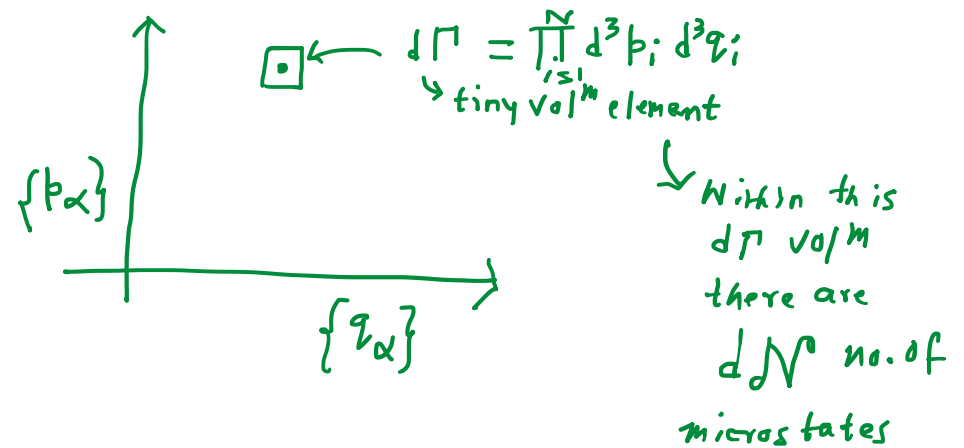
\downarrow
Ensemble density

$\underline{p} \equiv p_1, p_2, \dots$

$\underline{q} \equiv q_1, q_2, \dots$

Gibbs Ensemble \rightsquigarrow

$$P(\underline{p}, \underline{q}) = \lim_{N \rightarrow \infty} \frac{dN}{\int d\Gamma}$$



$$\int P(\underline{p}, \underline{q}) d\Gamma = \int \frac{dN}{N} = \frac{N}{N} = 1$$

↳ probability / density f^n

How this prob. density f^n evolve in time?

ensemble density

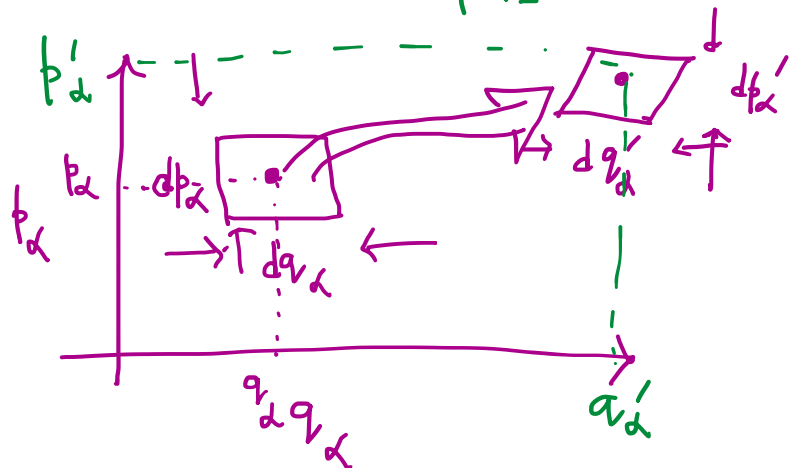
$$\langle x \rangle = \int x P(x) dx$$

$$\langle O(\underline{p}, \underline{q}) \rangle = \int O(\underline{p}, \underline{q}) P(\underline{p}, \underline{q}, t) d\Gamma$$

↓
Ensemble average

↑
in general it depends on time

Time evolution of $p(\underline{p}, \underline{q}, t)$ is given by "Liouville's theorem"



$d\Gamma$ (at time t)

$d\Gamma'$ (at time $t+dt$)

→ Distorted

$$\begin{aligned} q'_\alpha &= q_\alpha + \dot{q}_\alpha dt + O(dt^2) \rightarrow \sim 0 \\ p'_\alpha &= p_\alpha + \dot{p}_\alpha dt + O(dt^2) \rightarrow \sim 0 \end{aligned}$$

$$d\Gamma = dq_\alpha dp_\alpha$$

$$d\Gamma' = dq'_\alpha dp'_\alpha$$

$$d\Gamma' = dq_\alpha dp_\alpha \left[1 + \left(\frac{\partial \dot{q}_\alpha}{\partial q_\alpha} + \frac{\partial \dot{p}_\alpha}{\partial p_\alpha} \right) dt + O(dt^2) \right]$$

$$dq'_\alpha = dq_\alpha + \left(\frac{\partial \dot{q}_\alpha}{\partial q_\alpha} \right) dq_\alpha dt + O(\dots)$$

Derivative of the velocity with respect to the separation multiplied by the separation

$$dp'_\alpha = dp_\alpha + \left(\frac{\partial \dot{p}_\alpha}{\partial p_\alpha} \right) dp_\alpha dt + O(\dots)$$

$$d\Gamma' = \underbrace{dp_\alpha dq_\alpha}_{d\Gamma} \left[1 + \left\{ \frac{\partial \dot{q}_\alpha}{\partial q_\alpha} + \frac{\partial \dot{p}_\alpha}{\partial p_\alpha} \right\} dt + O(dt^2) \right]$$

$$d\Gamma = \prod_i d\vec{p}_i d\vec{q}_i$$

Hamilton's equation of motion

$$\frac{\partial H}{\partial p_\alpha} = \dot{q}_\alpha \quad \frac{\partial H}{\partial q_\alpha} = -\dot{p}_\alpha$$

$$d\Gamma' = d\Gamma \left[1 + \left(\frac{\partial^2 H}{\partial q_\alpha \partial p_\alpha} - \frac{\partial^2 H}{\partial p_\alpha \partial q_\alpha} \right) dt \right] \Rightarrow \text{Phase Space volume is}$$

Conserved Under Hamiltonian Dynamics

$$\prod_\alpha dp_\alpha dq_\alpha = \prod_\alpha dp'_\alpha dq'_\alpha$$

\leadsto Density does not change \leadsto Incompressible fluid

$$\vec{\nabla} \cdot \vec{v} = 0$$

\hookrightarrow Divergence of velocity field is zero.

$$\rho(\underline{p}', \underline{q}', t + dt) = \rho(\underline{p}, \underline{q}, t)$$

$$\begin{aligned} \rightarrow \rho(\underline{p}', \underline{q}', t + dt) &= \rho(\underline{p} + \dot{\underline{p}} \cdot dt, \underline{q} + \dot{\underline{q}} \cdot dt, t + dt) \\ &= \rho(\underline{p}, \underline{q}, t) + \left[\underbrace{\sum_{\alpha} \left(\dot{p}_{\alpha} \frac{\partial \rho}{\partial p_{\alpha}} + \dot{q}_{\alpha} \frac{\partial \rho}{\partial q_{\alpha}} \right) + \frac{\partial \rho}{\partial t}}_{=0} \right] dt \end{aligned}$$

Recall

$$\frac{df(\underline{p}, \underline{q}, t)}{dt} = \frac{\partial f}{\partial t} + \sum_{\alpha} \left(\frac{\partial f}{\partial p_{\alpha}} \cdot \dot{p}_{\alpha} + \frac{\partial f}{\partial q_{\alpha}} \cdot \dot{q}_{\alpha} \right)$$

Total derivative
of the fⁿ f

Streamline
derivative

Partial
derivative

$$\frac{d\rho}{dt} = 0 \rightarrow \text{Like incompressible fluid}$$

$$\boxed{\frac{\partial \rho}{\partial t} + \left[\sum_{\alpha} \left(\dot{p}_{\alpha} \frac{\partial \rho}{\partial p_{\alpha}} + \dot{q}_{\alpha} \frac{\partial \rho}{\partial q_{\alpha}} \right) \right] = 0}$$

$$\frac{\partial \rho}{\partial t} + \sum_{\alpha} \left\{ \left(\frac{\partial \rho}{\partial p_{\alpha}} \right) \dot{p}_{\alpha} + \left(\frac{\partial \rho}{\partial q_{\alpha}} \right) \dot{q}_{\alpha} \right\} = 0$$

$$\{A, B\} = \sum_{\alpha} \left(\frac{\partial A}{\partial q_{\alpha}} \frac{\partial B}{\partial p_{\alpha}} - \frac{\partial B}{\partial q_{\alpha}} \frac{\partial A}{\partial p_{\alpha}} \right)$$

↳ Poisson Bracket

$$\frac{\partial \rho}{\partial t} + \underbrace{\sum_{\alpha} \left\{ \frac{\partial \rho}{\partial p_{\alpha}} \left(-\frac{\partial H}{\partial q_{\alpha}} \right) + \left(\frac{\partial \rho}{\partial q_{\alpha}} \right) \left(\frac{\partial H}{\partial p_{\alpha}} \right) \right\}}_{\{ \rho, H \}}$$

$$\dot{p}_{\alpha} = -\frac{\partial H}{\partial q_{\alpha}}$$

$$\dot{q}_{\alpha} = \frac{\partial H}{\partial p_{\alpha}}$$

$$\frac{\partial \rho}{\partial t} + \{ \rho, H \} = 0$$

$$\text{or } \frac{\partial \rho}{\partial t} = -\{ \rho, H \} = \{ H, \rho \}$$

$$\boxed{\frac{\partial \rho}{\partial t} = \{ H, \rho \}}$$

↳ Liouville's theorem
or equation.

What is equilibrium in this perspective?

$$\frac{\partial \rho}{\partial t} = \{H, \rho\}$$

$\rho = \rho(H)$ → If the phase space density is a function of the Hamiltonian

$$\frac{\partial \rho}{\partial t} = 0$$

↳ ρ has no explicit time dependence

$$\{\rho, H\} = \{\rho(H), H\}$$

$$= \sum_{\alpha} \left(\underbrace{\left(\frac{\partial \rho}{\partial H} \right)}_{\rho'} \frac{\partial H}{\partial q_{\alpha}} \frac{\partial H}{\partial p_{\alpha}} - \underbrace{\left(\frac{\partial \rho}{\partial H} \right)}_{\rho'} \frac{\partial H}{\partial p_{\alpha}} \frac{\partial H}{\partial q_{\alpha}} \right)$$

$$= 0$$

$$= \rho' \{H, H\}$$

$$\langle O(\bar{p}, \bar{q}) \rangle = \int d\Gamma \underbrace{\rho(\underline{p}, \underline{q}, t)}_{\text{time-independent if}} \hat{O}(\bar{p}, \bar{q})$$

→ Definition of equilibrium

time-independent if $\rho(\underline{p}, \underline{q}, t) = \rho(\underline{p}, \underline{q}) = \rho(H) \leadsto$ is a fⁿ of H

$$P_{eq} = P(H)$$

$$L(\underline{p}, \underline{q})$$

Say for a system \rightarrow Angular momentum
is also conserved

$$\{L, H\} = 0$$

$$\frac{dL(\underline{p}, \underline{q})}{dt} = 0$$

$$= \sum_{\alpha} \left(\frac{\partial L}{\partial \dot{p}_{\alpha}} \dot{p}_{\alpha} + \frac{\partial L}{\partial \dot{q}_{\alpha}} \dot{q}_{\alpha} \right)$$

$$= \sum_{\alpha} \frac{\partial L}{\partial p_{\alpha}} \left(-\frac{\partial H}{\partial q_{\alpha}} \right) + \frac{\partial L}{\partial q_{\alpha}} \frac{\partial H}{\partial p_{\alpha}} = 0$$

$$P_{eq}(H, L)$$

$$P_{eq}(H(\underline{p}, \underline{q}), L(\underline{p}, \underline{q}))$$

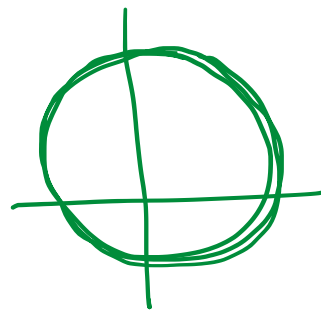
$$\{P_{eq}, H\} = P_1' \{H, H\} + P_2' \{L, H\} = 0$$

Isolate system
Energy is conserved
 E

} Microcanonical ensemble

$$\rho(H) \sim \delta(H - E)$$

↙ ↘
 \mathcal{E}_q^m distribution $f(H)$



Canonical
closed system

→ Boltzmann distribution

$$p_{eq}(H) \sim e^{-\beta H} \quad \beta = \frac{1}{k_B T}$$

Phase Space volume is conserved:

$$\frac{dP(\underline{p}, \underline{q}, t)}{dt} \rightarrow \text{Total time derivative of phase space density is zero}$$

What is equilibrium?

$$\langle O(\underline{p}, \underline{q}) \rangle = \text{Time independent}$$

↓
Ensemble average

$$= \int \underbrace{P(\underline{p}, \underline{q}, t)}_{P_{eq}(\underline{p}, \underline{q})} O(\underline{p}, \underline{q}) d\Gamma$$

$$P_{eq} \equiv P_{eq}(H) \rightarrow \text{Hamiltonian of the system}$$

$$H \rightarrow \text{Constant of motion} \quad \frac{dH}{dt} = 0$$

$$E = H \Rightarrow P_{eq} \sim \delta(H - E)$$

Canonical ensemble, $P_{eq} \sim e^{-\beta H}$ $\beta = \frac{1}{k_B T}$ → Boltzmann distribution

Microcanonical ensemble

Isolated System

(N, V, E)

First Postulate of Statistical Mechanics:

Time average = Ensemble average

↘ Distribution fn $P(\underline{p}, \underline{q}) \rightarrow$ System following classical mechanics
But if the system is quantum mechanical
↘ eigenfunctions and energy eigen values

$$\langle X \rangle = \frac{1}{N} \sum_{i=1}^N p_i x_i$$

$$\langle x \rangle = \frac{\sum p_i x_i}{\sum p_i} \quad \sum p_i = N$$

Postulate of equal a priori

Microcanonical Ensemble:

Systems of the ensemble are distributed uniformly, that is with equal probability over all possible microscopic states of the system. In the language of time trajectory, each state is visited on equal number of times if waited enough.

$$E = -\mu H$$

$H \rightarrow$ Magnetic field
 $\mu \rightarrow$ Magnetic moment

$$\begin{array}{ccccccc}
 \uparrow & \uparrow & \downarrow & \uparrow & \downarrow & \uparrow & \downarrow & \uparrow & \uparrow & \uparrow & \uparrow H \\
 -\mu H & -\mu H & +\mu H & & & -\mu H & & & -\mu H & & \\
 \hline
 & = -\mu H & & & & & & & & & \\
 \text{All are equally probable}
 \end{array}$$

Ergodic Hypothesis:

During its trajectory, in phase space, a system is free to explore all the microscopic states and given a sufficiently long period of time, spends time in a state that is proportional to the volume of the phase space.

↓
state in

$$\begin{array}{l}
 dp dq \sim h \rightarrow \text{Planck's Const} \\
 \downarrow \\
 d^{3N}p d^{3N}q \sim h^{3N} \\
 \rightarrow \text{smallest vol}^M \text{ element}
 \end{array}$$



Huge barrier

Energy landscape of glass
(Rugged energy landscape)

"Non-ergodic"

^ Supercooled liquid ^

n possible outcomes, each with probability p_j , where $j=1, 2, \dots, n$

→ Discrete prob. distribution

If the experiment is repeated indefinitely,

$$p_j = \lim_{N \rightarrow \infty} \frac{N_j}{N} \quad j=1, 2, \dots, n$$

$N_j \rightarrow$ is the # of times of outcome j

$N \rightarrow$ total # of repetitions of the expt.

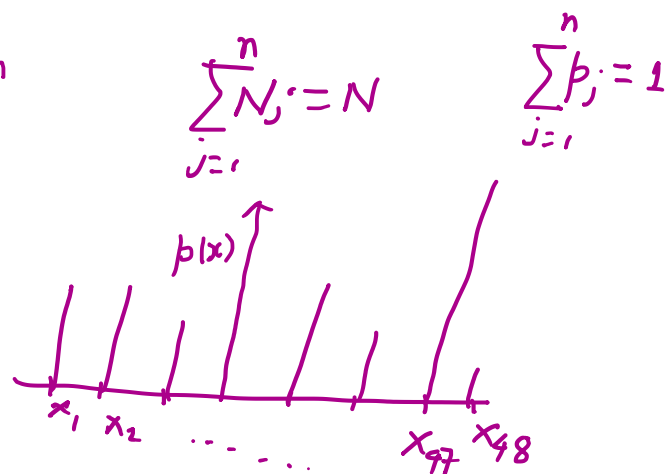
Average x

$$\langle x \rangle = \sum_{j=1}^n x_j p_j$$

↓
First moment of this distribution

$$\langle x^2 \rangle = \sum_{j=1}^n x_j^2 p_j$$

↳ second moment



Second central moment or Variance:

$$\sigma_x^2 = \langle (x - \langle x \rangle)^2 \rangle = \sum_{j=1}^n (x_j - \langle x \rangle)^2 p_j \rightarrow \text{positive definition}$$

↳ measure of the spread of the distribution

$$\sigma_x^2 = \sum_{j=1}^n (x_j^2 - 2x_j \langle x \rangle + \langle x \rangle^2) p_j = \underbrace{\sum_{j=1}^n x_j^2 p_j}_{\langle x^2 \rangle} - 2 \underbrace{\left(\sum_{j=1}^n x_j p_j \right)}_{\langle x \rangle} \langle x \rangle + \langle x \rangle^2 \underbrace{\sum_{j=1}^n p_j}_1$$

$$= \langle x^2 \rangle - 2\langle x \rangle^2 + \langle x \rangle^2 = \langle x^2 \rangle - \langle x \rangle^2 \geq 0$$

$$\langle x^2 \rangle \geq \langle x \rangle^2$$

Continuous distribution

$$\text{Prob}(x, x+dx) = p(x) dx$$
$$-\infty \leq x \leq \infty$$

$$\int_{-\infty}^{+\infty} p(x) dx = 1 \quad (\text{normalized / total prob} = 1)$$
$$\int_{-\infty}^{+\infty} x p(x) dx = \langle x \rangle$$
$$\int_{-\infty}^{+\infty} x^2 p(x) dx = \langle x^2 \rangle$$
$$\int_{-\infty}^{+\infty} x^n p(x) dx = \langle x^n \rangle \rightarrow n^{\text{th}} \text{ moment}$$

Example:

$$p(x) = A$$

$$a \leq x \leq b$$

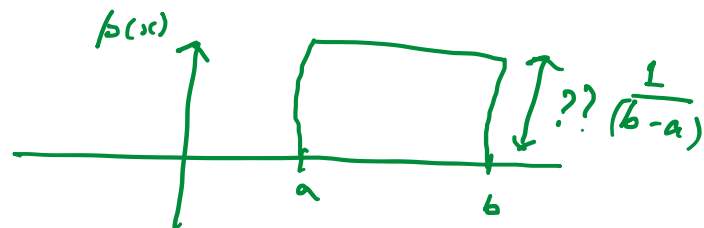
= 0

otherwise

$$\int_a^b p(x) dx = A \int_a^b dx = 1 \Rightarrow A(b-a) = 1$$

or $A = \frac{1}{(b-a)}$

$$\langle x \rangle = \frac{1}{(b-a)} \int_a^b x dx = \frac{1}{(b-a)} \left[\frac{x^2}{2} \right]_a^b = \frac{b^2 - a^2}{2(b-a)} = \frac{1}{2} (b+a)$$



$$\langle x^2 \rangle = \frac{1}{b-a} \int_a^b x^2 dx = \frac{1}{(b-a)} \frac{1}{3} (b^3 - a^3) = \frac{(b-a)(b^2 + ab + a^2)}{3(b-a)} = \frac{1}{3} (b^2 + ab + a^2)$$

Variance

$$\sigma_x^2 = \langle x^2 \rangle - \langle x \rangle^2 = \frac{1}{3} (b^2 + ab + a^2) - \frac{1}{4} (b + 2ab + a^2)$$

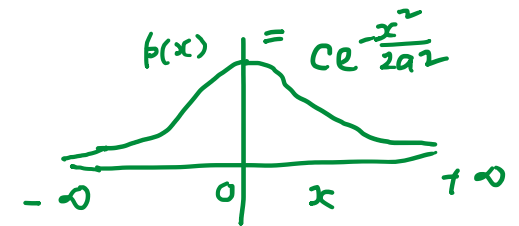
$$= \frac{(b-a)^2}{12}$$

$$\sqrt{\sigma_x^2} = \text{Standard deviation}$$

$$= \frac{1}{\sqrt{12}} (b-a)$$

* Most common continuous prob. distribution \rightarrow Gaussian or a normal distribution

$$p(x) = C e^{-x^2/2a^2} \quad -\infty < x < \infty$$



$$\int_{-\infty}^{+\infty} p(x) dx = 1 = C \int_{-\infty}^{+\infty} e^{-\frac{x^2}{2a^2}} dx$$

$$\alpha \equiv \frac{1}{2a^2}$$

Standard Integral

$$\int_{-\infty}^{+\infty} e^{-\alpha x^2} dx = (\pi/\alpha)^{\frac{1}{2}}$$

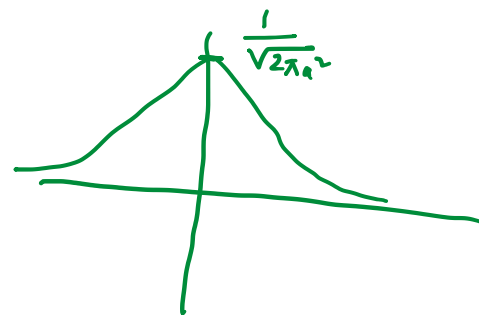
$$1 = C \sqrt{\pi} \frac{1}{\sqrt{\frac{1}{2a^2}}}$$

$$1 = C \sqrt{\pi} \sqrt{2} a$$

$$\Rightarrow C = \frac{1}{\sqrt{2\pi} a^2}$$

$$p(x) = \frac{1}{\sqrt{2\pi} a^2} e^{-\frac{x^2}{2a^2}}$$

$$\text{at } x=0 \quad p(0) = \frac{1}{\sqrt{2\pi} a^2}$$



$\langle x \rangle = ?$ $\langle x^3 \rangle = 0$ $\langle x^n \rangle = 0$ if n is odd (Integration of an odd integrand within symmetric limits)

$$\langle x^2 \rangle = \int_{-\infty}^{+\infty} x^2 \frac{1}{\sqrt{2\pi a^2}} e^{-\frac{x^2}{2a^2}} dx = \frac{1}{\sqrt{2\pi a^2}} \int_{-\infty}^{+\infty} x^2 e^{-\frac{x^2}{2a^2}} dx$$

$$I = \int_{-\infty}^{+\infty} e^{-\alpha x^2} = \left(\pi/\alpha\right)^{\frac{1}{2}}$$

$$\frac{dI}{d\alpha} = \int_{-\infty}^{+\infty} -x^2 e^{-\alpha x^2} \Rightarrow \int_{-\infty}^{+\infty} x^2 e^{-\alpha x^2} = -\frac{dI}{d\alpha} = +\frac{1}{2} \sqrt{\pi} \alpha^{-3/2} = \frac{1}{2\alpha} \left(\pi/\alpha\right)^{\frac{1}{2}}$$

$$\alpha \equiv \frac{1}{2a^2}$$

$$\langle x^2 \rangle = \frac{1}{\sqrt{2\pi a^2}} \frac{1}{2 \cdot \frac{1}{2a^2}} \left(\pi/\frac{1}{2a^2}\right)^{\frac{1}{2}} = \frac{1}{\sqrt{2\pi a^2}} a^2 \sqrt{2\pi a^2} = a^2$$

$$\sigma_x^2 = \langle x^2 \rangle - \langle x \rangle^2 = a^2 - 0 = a^2 \Rightarrow p(x) = \frac{1}{\sqrt{2\pi\sigma^2}} e^{-\frac{x^2}{2\sigma^2}} \quad \sigma^2 \equiv \text{variance}$$

$$p(x) = \frac{1}{\sqrt{2\pi\sigma^2}} e^{-\frac{x^2}{2\sigma^2}}$$

More general $\langle x \rangle = \mu$ (non-zero)

$$\int_{-\infty}^{+\infty} e^{-x^2} dx = \sqrt{\pi}$$

$$p(x) = \frac{1}{\sqrt{2\pi\sigma_x^2}} e^{-\frac{(x-\mu)^2}{2\sigma_x^2}} \quad \text{is this normalized?}$$

$$\frac{x-\mu}{\sqrt{2}\sigma_x} = y$$

$$dx = dy \sqrt{2}\sigma_x$$

$$\sigma_x^2 = \langle x^2 \rangle - \langle x \rangle^2$$

$$\int_{-\infty}^{+\infty} p(x) dx = \int_{-\infty}^{+\infty} \frac{1}{\sqrt{2\pi\sigma_x^2}} e^{-y^2} \sqrt{2}\sigma_x dy = \frac{\sqrt{2}\sigma_x}{\sqrt{2}\sigma_x\sqrt{\pi}} \int_{-\infty}^{+\infty} e^{-y^2} dy$$

$$\langle x \rangle = \mu \quad \langle x^2 \rangle = \sigma_x^2 + \mu^2 \quad \text{check}$$

$$x - \mu = z$$

$$x = z + \mu$$

$$dx = dz$$

$$\langle x \rangle = \int_{-\infty}^{+\infty} x p(x) = \int_{-\infty}^{+\infty} dz (z + \mu) e^{-\frac{z^2}{2\sigma_x^2}} \left(\frac{1}{\sqrt{2\pi\sigma_x^2}} \right) = \frac{1}{\sqrt{2\pi\sigma_x^2}} \int_{-\infty}^{+\infty} z e^{-\frac{z^2}{2\sigma_x^2}} dz + \frac{1}{\sqrt{2\pi\sigma_x^2}} \mu \int_{-\infty}^{+\infty} dz e^{-\frac{z^2}{2\sigma_x^2}} = \mu$$

odd integrand

$= \frac{1}{\sqrt{\pi}} \sqrt{\pi} = 1$ (normalized)

$$\langle x^2 \rangle = \int_{-\infty}^{+\infty} dx x^2 \frac{1}{\sqrt{2\pi\sigma_x^2}} e^{-\frac{(x-\mu)^2}{2\sigma_x^2}} = \int_{-\infty}^{+\infty} (z+\mu)^2 \frac{dz}{\sqrt{2\pi\sigma_x^2}} e^{-\frac{z^2}{2\sigma_x^2}} = \int_{-\infty}^{+\infty} z^2 \frac{1}{\sqrt{2\pi\sigma_x^2}} e^{-\frac{z^2}{2\sigma_x^2}} dz + \mu^2 \int_{-\infty}^{+\infty} \frac{dz}{\sqrt{2\pi\sigma_x^2}} e^{-\frac{z^2}{2\sigma_x^2}} \quad \downarrow = 1$$

$$\langle x^2 \rangle = \mu^2 + \frac{1}{\sqrt{2\pi\sigma_x^2}} \int_{-\infty}^{+\infty} z^2 e^{-\frac{z^2}{2\sigma_x^2}} dz$$

$$\text{Let, } \frac{z}{\sqrt{2}\sigma_x} = \tilde{z}$$

$$dz = \sqrt{2}\sigma_x d\tilde{z}$$

$$\frac{1}{\sqrt{2\pi\sigma_x^2}} \int_{-\infty}^{+\infty} \tilde{z}^2 (2\sigma_x^2) e^{-\tilde{z}^2} d\tilde{z} (\sqrt{2}\sigma_x)$$

$$= \frac{1}{\sqrt{2\pi\sigma_x^2}} 2\sigma_x^2 \sqrt{2}\sigma_x \int_{-\infty}^{+\infty} \tilde{z}^2 e^{-\tilde{z}^2} d\tilde{z}$$

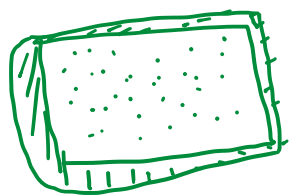
$$= \frac{2}{\sqrt{\pi}} \sigma_x^2 \frac{1}{2} \sqrt{\pi} = \sigma_x^2$$

$$\langle x^2 \rangle = \mu^2 + \sigma_x^2 \quad (\text{checked})$$

$$+ 2\mu \int_{-\infty}^{+\infty} \frac{dz}{\sqrt{2\pi\sigma_x^2}} z e^{-\frac{z^2}{2\sigma_x^2}} \quad \downarrow 0$$

$$\int_{-\infty}^{+\infty} x^2 e^{-\alpha x^2} dx = \frac{1}{2\alpha} \left(\frac{\pi}{\alpha} \right)^{\frac{1}{2}}$$

Microcanonical Ensemble:



$$\begin{aligned} dQ &= 0 & dW &= 0 \\ dE &= 0 & & \text{(Isolated System)} \end{aligned}$$

$M \equiv (N, V, E) \rightarrow \text{fixed}$
 $\rightarrow \text{Macrostate}$

"Postulate of equal a priori"

$$\Omega(N, V, E) \rightarrow \frac{1}{\Omega(N, V, E)} \rightarrow \text{prob. of finding the system in } i\text{th microstate}$$

\downarrow No. of microstates

Gibbs $\rightarrow S = -k_B \sum_i p_i \ln p_i = -k_B \langle \ln p \rangle$

\nwarrow Boltzmann constant

\downarrow Shannon's entropy

Third law:

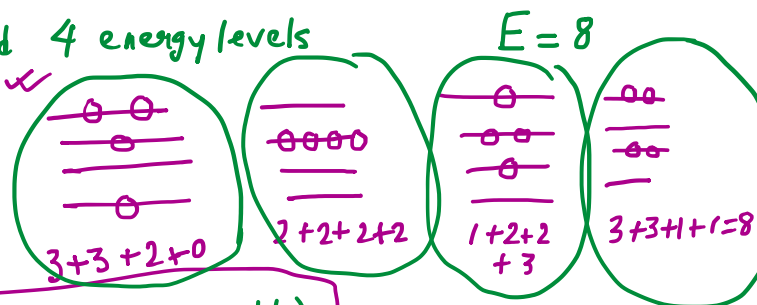
$$\begin{aligned} \Omega &= 1 \text{ (only one microstate)} \\ S &= 0 \end{aligned}$$

4 particles and 4 energy levels

3 —○—
 2 —○—
 1 —○—
 0 —○—

$$3+2+1+0 = 6 \neq E \text{ (Not acceptable)}$$

"4 microstate"



Each of these arrangements represent a microstate

$$S = -k_B \sum_i \frac{1}{\Omega} \ln\left(\frac{1}{\Omega}\right)$$

$$S = -k_B \frac{\Omega}{\Omega} (\ln 1 - \ln \Omega)$$

$$S = k_B \ln \Omega \rightarrow \text{"Boltzmann entropy"}$$

Cyclic Relation

$$x = f(y, z) \quad y = f(x, z)$$

$$\left(\frac{\partial x}{\partial y}\right)_z \left(\frac{\partial y}{\partial z}\right)_x \left(\frac{\partial z}{\partial x}\right)_y = -1$$

$$dx = \left(\frac{\partial x}{\partial y}\right)_z dy + \left(\frac{\partial x}{\partial z}\right)_y dz \quad \dots \dots \textcircled{1}$$

$$dy = \left(\frac{\partial y}{\partial x}\right)_z dx + \left(\frac{\partial y}{\partial z}\right)_x dz \quad \dots \dots \textcircled{2}$$

$$dx = \left(\frac{\partial x}{\partial z}\right)_y dz + \underbrace{\left(\frac{\partial x}{\partial y}\right)_z \left(\frac{\partial y}{\partial z}\right)_x}_{1} dz + \left(\frac{\partial x}{\partial y}\right)_z \left(\frac{\partial y}{\partial z}\right)_x dz$$

$$dx = \left(\frac{\partial x}{\partial z}\right)_y dz + \left(\frac{\partial x}{\partial y}\right)_z \left(\frac{\partial y}{\partial z}\right)_x dz + dx \Rightarrow \left(\frac{\partial x}{\partial y}\right)_z \left(\frac{\partial y}{\partial z}\right)_x \left(\frac{\partial z}{\partial x}\right)_y = -1$$

$$\left(\frac{\partial x}{\partial y}\right)_z \left(\frac{\partial y}{\partial z}\right)_x \left(\frac{\partial z}{\partial x}\right)_y = -1$$

$$S, V, E \quad \left(\frac{\partial S}{\partial V}\right)_{N, E} \left(\frac{\partial V}{\partial E}\right)_{N, S} \left(\frac{\partial E}{\partial S}\right)_{N, V} = -1$$

$$\left(\frac{\partial S}{\partial V}\right)_{N, E} = \frac{-1}{\left(\frac{\partial V}{\partial E}\right)_{N, S} \left(\frac{\partial E}{\partial S}\right)_{N, V}}$$

$$\left(\frac{\partial S}{\partial V}\right)_{N, E} = \frac{-1}{-\frac{1}{P} T} = \frac{P}{T}$$

$$S, N, E \quad \left(\frac{\partial S}{\partial N}\right)_{E, V} \left(\frac{\partial N}{\partial E}\right)_{S, V} \left(\frac{\partial E}{\partial S}\right)_{N, V} = -1$$

$$\Rightarrow \left(\frac{\partial S}{\partial N}\right)_{E, V} = -\frac{1}{\left(\frac{\partial N}{\partial E}\right)_{S, V} \left(\frac{\partial E}{\partial S}\right)_{N, V}}$$

$$\left(\frac{\partial S}{\partial N}\right)_{E, V} = \frac{-1}{\left(\frac{1}{\mu}\right) T} = -\frac{\mu}{T}$$

$$E = TS - PV + \mu N$$

$$\left(\frac{\partial E}{\partial S}\right)_{V, N} = T \quad \left(\frac{\partial E}{\partial V}\right)_{S, N} = -P$$

$$\left(\frac{\partial E}{\partial N}\right)_{S, V} = \mu \text{ (chemical potential)}$$

$$S(N, V, E)$$

$$dS = \left(\frac{\partial S}{\partial N} \right)_{V, E} dN + \left(\frac{\partial S}{\partial V} \right)_{E, N} dV + \left(\frac{\partial S}{\partial E} \right)_{N, V} dE$$

$$dS = - \sum_i \frac{\mu_i dN_i}{T} + \frac{PdV}{T} + \frac{dE}{T}$$

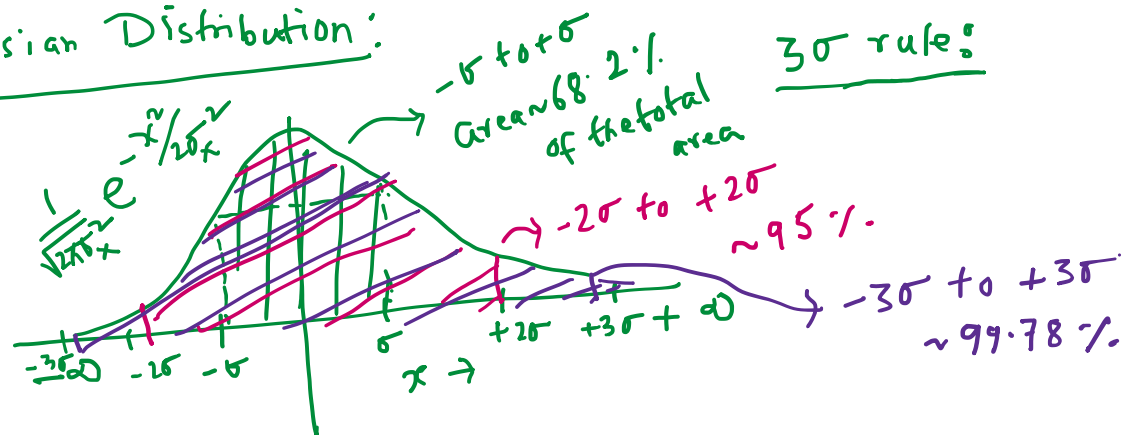
$$\frac{1}{T} = \left(\frac{\partial S}{\partial E} \right)_{V, N} = \frac{\partial}{\partial E} (k_B \ln \Omega)_{V, N} = k_B \left(\frac{\partial \ln \Omega}{\partial E} \right)_{V, N}$$

↓
Definition of
temperature

$$\frac{P}{T} = \left(\frac{\partial S}{\partial V} \right)_{N, E} = k_B \left(\frac{\partial \ln \Omega}{\partial V} \right)_{E, N}$$

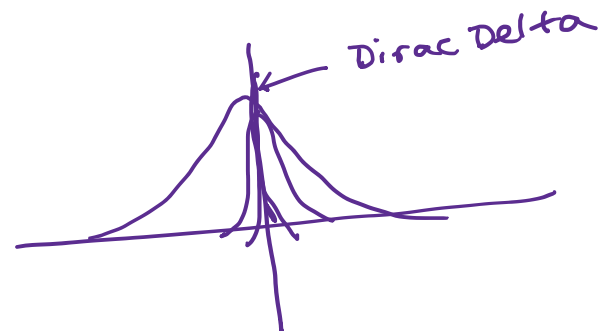
$$-\frac{\mu_i}{T} = \left(\frac{\partial S}{\partial N_i} \right)_{E, V, N_{\alpha \neq i}} = k_B \left(\frac{\partial \ln \Omega}{\partial N_i} \right)_{E, V, N_{\alpha \neq i}}$$

Gaussian Distribution:



$$\sigma_x \rightarrow 0 \quad p(x) \sim \delta(x)$$

↳ Dirac-Delta function



$$f(x) \sim \infty \quad x=0$$

$$= 0 \quad \text{Otherwise}$$

↳ $\delta(x) \rightarrow$ Dirac-Delta function.

$\delta(x-a)$ peaked at $x=a$

$$\int_{-\infty}^{+\infty} \delta(x) dx = 1 \quad (\text{normalized})$$

Dimension of Dirac-Delta function is $\frac{1}{\text{argument}}$

$$\delta(x) \sim \frac{1}{x}$$

$$\delta(t) \sim \frac{1}{t}$$

$$\int_{-\infty}^{+\infty} dx f(x) \delta(x) = f(0) \quad \int_{-\infty}^{+\infty} dx f(x) \delta(x-a) = f(a)$$

$$\delta(-x) = \delta(x) \quad \text{even function}$$

Revisiting $\Omega(N, V, E)$

$$H(\underline{p}, \underline{q}) = E \quad (\text{constraint})$$

\hookrightarrow is a fⁿ of
6N-coordinates

Recall Liouville's theorem

Equilibrium probability

distribution in microcanonical ensemble

$$P_{eq}(H(\underline{p}, \underline{q})) = N_c \delta(H(\underline{p}, \underline{q}) - E)$$

(6N-1) dimensional hypersurface

$$\Omega(N, V, E)$$

\hookrightarrow Number of allowed microstates \equiv measure of the amount of phase space available

$$\Omega(N, V, E) = \tilde{N} \int d^3p_N \int d^3q_N \delta(H(\underline{p}, \underline{q}) - E) \rightarrow \text{integration over entire phase space}$$

Quantum mechanics:

$\Delta p \Delta q \sim h$ (Planck's constant)
- Heisenberg uncertainty principle -

We divide the entire phase space into hypervolumes, each with vol^M $\Delta \Gamma = \Delta p \Delta q = h^{3N}$, such that, each of these hypervolumes contain ONLY one(1) microstate.

$$\Omega(N, V, E) = \sum_{\substack{\text{hypervol}^M \\ E < H(\underline{p}, \underline{q}) < E + E_0}} \frac{\Delta \Gamma}{h^{3N}}$$

Mark E Tuckerman
Statistical Mechanics

$\Delta \Gamma$ is small compared to the entire phase space vol^M

$$\Delta \Gamma = \int \dots \int d^3p_N d^3q_N$$

$$\Omega(N, V, E) = \frac{1}{h^{3N}} \int_{E < H < E + E_0} d\Gamma \rightsquigarrow \text{Approximate} = \frac{E_0}{h^{3N}} \int \dots \int d^3p_N d^3q_N \delta(H(\underline{p}, \underline{q}) - E)$$

$$\Omega(N, V, E) = \frac{E_0}{h^{3N} N!} \int \dots \int d^3p_N d^3q_N \delta(H(p, q) - E)$$

\tilde{M}

Since the particles are indistinguishable

Particles are distinguishable

$$\text{prob.} = \frac{p_i}{\sum p_i}$$

$$P_{eq} = \frac{1}{\Omega(N, V, E)} \times \tilde{M} \delta(H(p, q) - E)$$

Microcanonical ensemble

$$P_{eq} = \frac{\tilde{M} \delta(H(p, q) - E)}{\tilde{M} \int \dots \int d^3p_N d^3q_N \delta(H(p, q) - E)} = \frac{\delta(H(p, q) - E)}{\int \dots \int d^3p_N d^3q_N \delta(H(p, q) - E)}$$

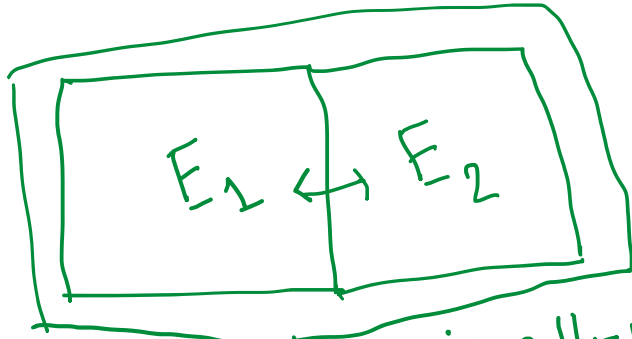
$$\langle A(\underline{p}, \underline{q}) \rangle = \frac{\int A(\underline{p}, \underline{q}) \delta(H(\underline{p}, \underline{q}) - E) d\Gamma}{\int d\Gamma \delta(H(\underline{p}, \underline{q}) - E)}$$

$\Omega(N, V, E) \leadsto$ Microcanonical partition function

$\left\{ \begin{array}{c} 3.2.22 \\ \text{in} \\ \text{MT.} \\ \text{SM} \end{array} \right\}$ Page 80

$$\Omega(N, V, E) \equiv \Omega(E, X) \equiv \Omega(E)$$

↪ generalized
displacement
co-ordinate



Heat exchange is allowed -

$$S_1 = k_B \ln \Omega(E_1)$$

$$S_2 = k_B \ln \Omega(E_2)$$

$$E_1 + E_2 = E \text{ (constant)}$$

$\Omega(E) \rightarrow$ Accessible no. of
microstates for
the combined system



$$\Omega(E) = \int dE_1 \Omega(E_1) \Omega(E-E_1) = \int dE_1 \exp\left[\frac{S_1(E_1)}{k_B} + \frac{S_2(E_2)}{k_B}\right] \simeq \underbrace{\exp\left[\frac{S_1(E_1^*)}{k_B} + \frac{S_2(E_2^*)}{k_B}\right]}_{\text{Approximate}}$$

$$\underbrace{k_B \ln \Omega(E)}_{S(E)} = k_B \ln \Omega(E_1^*) + k_B \ln \Omega(E_2^*)$$

- extensivity of entropy -

$$S(E) = S_1(E_1^*) + S_2(E_2^*)$$

Extremise the exponent with respect to E_1

$$\frac{\partial}{\partial E_1} \left[\frac{S_1(E_1)}{k_B} + \frac{S_2(E_2)}{k_B} \right] = 0$$

$$\left. \frac{\partial S_1(E_1)}{\partial E_1} \right|_{x_1} + \left. \frac{\partial S_2(E_2)}{\partial E_1} \right|_{x_2} = 0$$

Condition for (thermal) \leftarrow
eqm

$$\rightarrow \left. \frac{\partial S_1(E_1)}{\partial E_1} \right|_{x_1} - \left. \frac{\partial S_2(E_2)}{\partial E_2} \right|_{x_2} = 0$$

$$\left. \frac{\partial S_1(E_1)}{\partial E_1} \right|_{x_1} = \left. \frac{\partial S_2(E_2)}{\partial E_2} \right|_{x_2} = \frac{1}{T}$$

(emfirical temperature)

Zerth
law of Thermo-
dynamics

$$E = E_1 + E_2$$

$$E_1 = E - E_2$$

$$\frac{\partial}{\partial E_1} = - \frac{\partial}{\partial E_2}$$

First law of thermodynamics:

changing the co-ordinate reversibly by δX
 $X \rightarrow X + \delta X$

$$S(E, X)$$

\hookrightarrow

$$S(E + J \cdot \delta X, X + \delta X)$$

$$\delta W = J \cdot \delta X \quad (\text{work done})$$

\hookrightarrow generalized force

$$\delta S = S(E + J \cdot \delta X, X + \delta X) - S(E, X)$$

\downarrow

change
in entropy

$$= S(E, X) + \left. \frac{\partial S}{\partial E} \right|_X J \delta X + \left. \frac{\partial S}{\partial X} \right|_E \delta X - S(E, X)$$

$$\delta S = \left(\left. \frac{\partial S}{\partial E} \right|_X J + \left. \frac{\partial S}{\partial X} \right|_E \right) \delta X$$

$\delta S = 0$ (system is already in eq^m)

$$\Rightarrow \left(\left. \frac{\partial S}{\partial E} \right|_X J + \left. \frac{\partial S}{\partial X} \right|_E \right) \delta X = 0$$
$$\Rightarrow \left. \frac{\partial S}{\partial E} \right|_X J = - \left. \frac{\partial S}{\partial X} \right|_E$$
$$\Rightarrow \frac{1}{T} = - \left. \frac{\partial S}{\partial X} \right|_E$$

$$dS(E, X) = \left. \frac{\partial S}{\partial E} \right|_X dE + \left. \frac{\partial S}{\partial X} \right|_E dX = \frac{dE}{T} - \frac{J dX}{T}$$

$$\Rightarrow \boxed{dE = T dS + J dX}$$

\hookrightarrow 1st Law of thermodynamics

Second law of thermodynamics:

$$\Omega_1(E_1^*, x) \Omega_2(E_2^*, x) \geq \Omega_1(E_1, x) \Omega_2(E_2, x)$$

E_{eq} has a large
no. of accessible
states than any
(other state) starting point

$$\Delta S = S_1(E_1^*) + S_2(E_2^*) - S_1(E_1) - S_2(E_2) \geq 0$$

$$\Delta S = \left(\left. \frac{\partial S_1}{\partial E_1} \right|_{x_1} - \left. \frac{\partial S_2}{\partial E_2} \right|_{x_2} \right) \Delta E_1 = \left(\frac{1}{T_1} - \frac{1}{T_2} \right) \Delta E_1 \geq 0$$

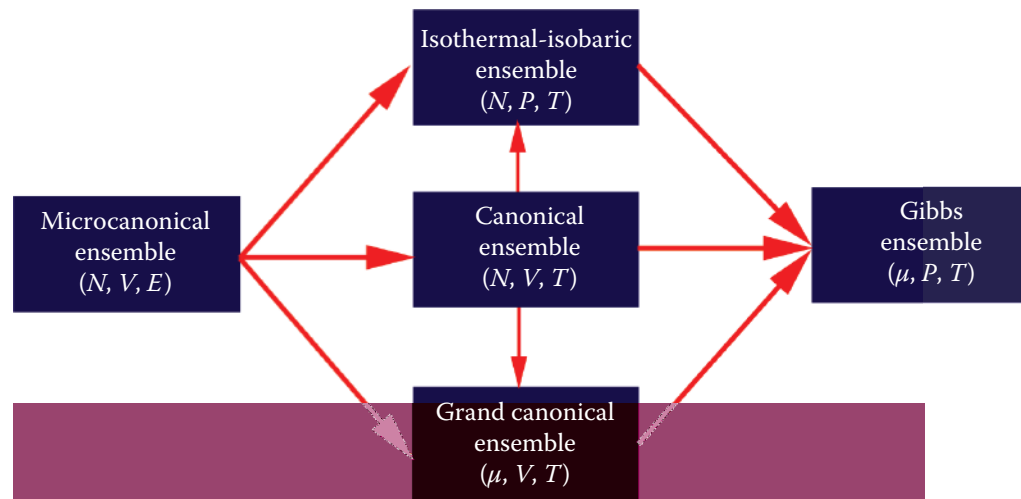
heat flows from hot to cold body \rightarrow 2nd law of thermodynamics

Ensemble

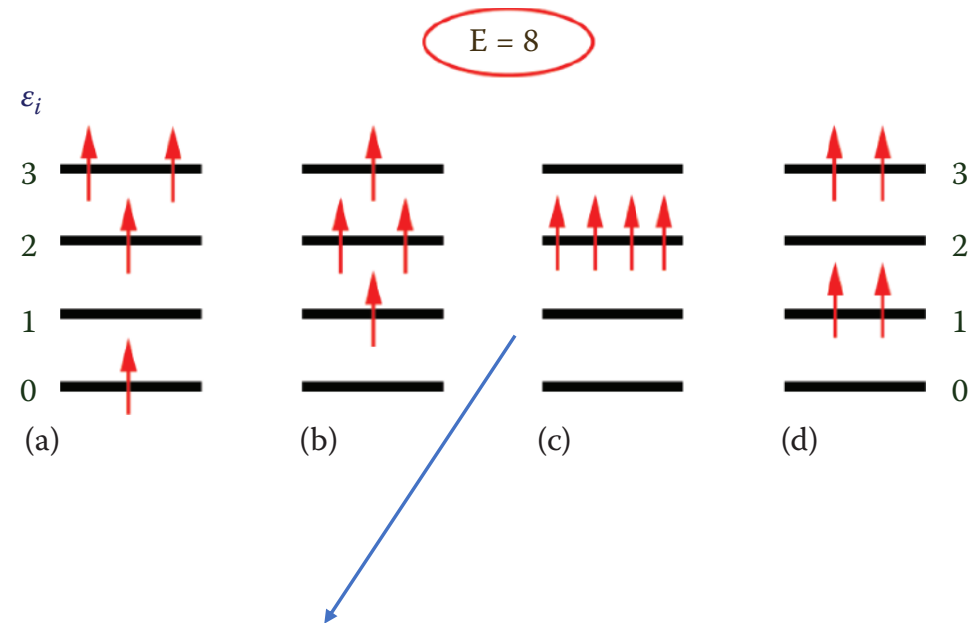
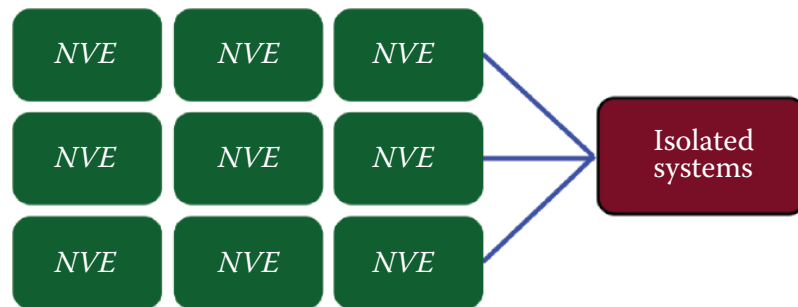
The concept of an ensemble is a **brilliant mental construct**

System must have a large number of microscopic states (positions and momenta), and natural motion of system at non-zero temperature takes the system through a finite fraction of these states in a time comparable to time of measurement of the macroscopic properties.

Different types of Ensembles

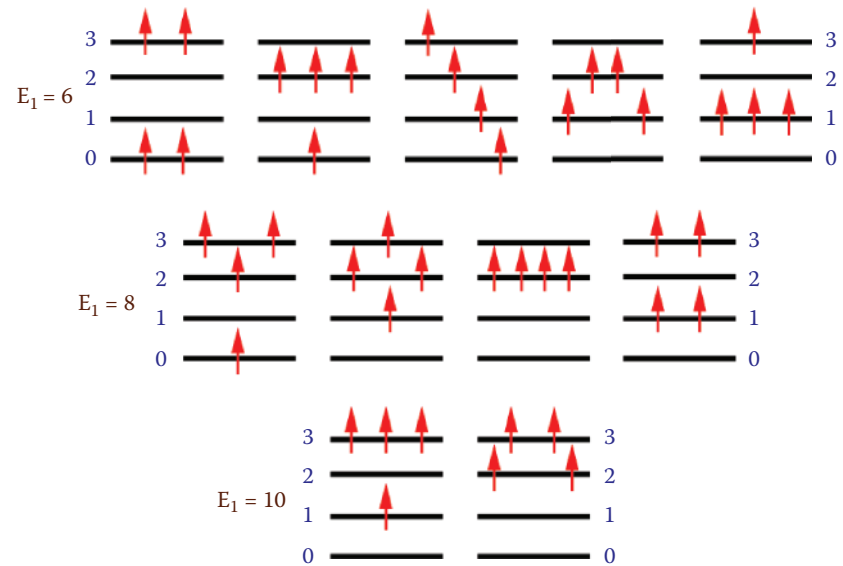
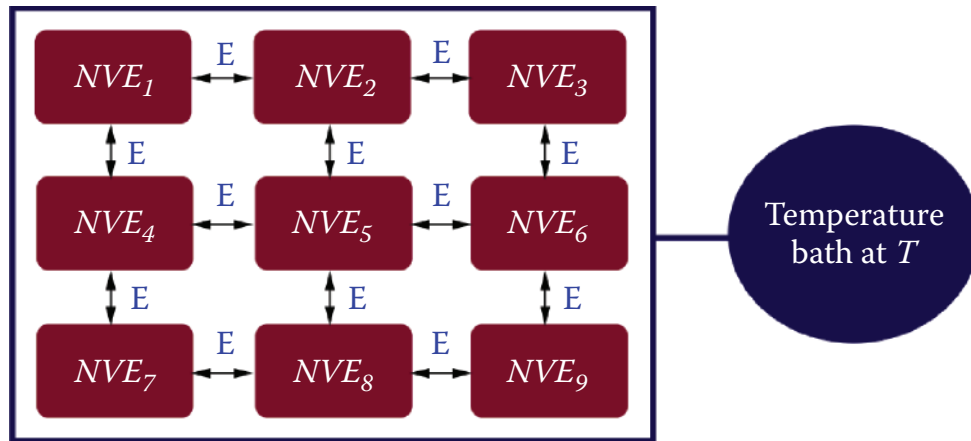


A **microcanonical ensemble** consists of mental replicas of the original (NVE) system (**Isolated System**)



Four arrangements giving rise to same total energy (these are microstates)

Canonical ensemble



Canonical ensemble, where energy of each system can fluctuate. The systems are kept at temperature T by putting the super-system in a bath and establishing thermal contacts between the individual systems

$N_s \rightarrow$ Total no. of Systems

$N \rightarrow$ No. of particles in each System

Super-System
(Isolated)

$$N_t = N_s N$$

\rightarrow total no. of particles

$$V_t = N_s V$$

\rightarrow total volume

$$E_t = \text{Total energy}$$

$$(N_t, V_t, E_t)$$

then we
can construct
a microcanonical
ensemble starting
here

$$\sum_j n_j = N_s$$

$$\sum_j n_j E_j = E_t$$

\rightarrow constraints

If the particles are distinguishable

$$\Omega(\{n_j\}) = \frac{N_s!}{n_1! n_2! n_3! \dots} \quad (\text{elaborate this later}) \dots \textcircled{1}$$

↘ No. of ways

* Biman Bagchi }
* McQuarrie }

Probability of observing a given state n_j with energy E_j

$$p_j = \frac{\bar{n}_j}{N_s} = \frac{1}{N_s} \left\{ \frac{\sum_j n_j \Omega(\{n_j\})}{\sum_j \Omega(\{n_j\})} \right\} \rightarrow \bar{n}_j \dots \textcircled{2}$$

[In general, there are many distributions that are consistent with $\sum_j n_j = N_s$.]

$$\simeq \frac{1}{N_s} \frac{n_j^* \Omega(\{n_j^*\})}{\Omega(\{n_j^*\})} = \frac{n_j^*}{N_s}$$

↘ Most probable value

$n_j^* \rightarrow$ is the value of n_j that maximises $\Omega(\{n_j\})$

In any particular distribution, $\frac{n_j}{N_s}$ is

the fraction of systems of the canonical ensemble

in the j -th energy state. The overall probability

p_j that a system is in the j -th energy state is

obtained by averaging $\frac{n_j}{N_s}$ over all the allowed distributions, giving equal weight to the each one
principle of equal a priori probabilities

$$\frac{\partial}{\partial n_j} \left(\ln \Omega(\{n_j\}) - \alpha \sum_j n_j - \beta \sum_j n_j E_j \right) = 0$$

(j)
 LUM

$$\left(0 - \frac{1}{n_j} - \ln n_j^* - \alpha - \beta E_j \right) = 0$$

$$-1 - \ln n_j^* - \alpha - \beta E_j = 0$$

$$\text{or, } \ln n_j^* = -(1 + \alpha) - \beta E_j \quad \alpha' = 1 + \alpha$$

$$= -\alpha' - \beta E_j$$

$$n_j^* = e^{-\alpha'} e^{-\beta E_j}$$

$$\sum n_j^* = N_s = e^{-\alpha'} \sum_j e^{-\beta E_j}$$

$$\text{or, } e^{+\alpha'} = \frac{\sum_j e^{-\beta E_j}}{N_s}$$

\Rightarrow

$$N_s = \sum n_j$$

$$\sum n_j E_j = E_t$$

$$\ln N! \approx N \ln N - N$$

(Stirling Approximation)

$$\ln \Omega(\{n_j\}) = \ln N_s! - \ln \prod_j n_j!$$

$$= N_s \ln N_s - N_s - \sum_j \ln n_j!$$

[$n_1! n_2! \dots$]

$$= N_s \ln N_s - N_s - \sum_j n_j \ln n_j + \sum_j n_j$$

$= \prod_j n_j!$

$\rightarrow N_s$

$$\ln \Omega = N_s \ln N_s - \sum_j n_j \ln n_j$$

$$P_j = \frac{n_j^*}{N_s} = \frac{N_s}{\sum_j e^{-\beta E_j}} \cdot \frac{e^{-\beta E_j}}{N_s}$$

$$P_j = \frac{e^{-\beta E_j}}{\sum_j e^{-\beta E_j}}$$

$$Q_N(V, T) = \sum_j e^{-\beta E_j}$$

Canonical
partition
function

\leadsto Boltzmann distribution

What about β ? \rightarrow LATER and show $\beta = \frac{1}{k_B T}$

$$\bar{E} = \sum_j P_j E_j = \frac{\sum_j E_j e^{-\beta E_j}}{\sum_j e^{-\beta E_j}}$$

$$d\bar{E} = \sum_j (E_j dP_j + P_j dE_j)$$

$$P_j = \frac{e^{-\beta E_j}}{Q}, \quad \ln P_j = -\beta E_j - \ln Q$$

or, $E_j = -\frac{1}{\beta} (\ln P_j + \ln Q)$

$$d\bar{E} = \sum_j \left(-\frac{1}{\beta} \right) (\ln P_j + \ln Q) dP_j + \sum_j P_j \left(\frac{\partial E_j}{\partial V} \right) dV$$

$$d\bar{E} = -\frac{1}{\beta} \sum_j \ln P_j dP_j - \frac{1}{\beta} \ln Q \sum_j dP_j + \sum_j P_j \left(\frac{\partial E_j}{\partial V} \right) dV$$

$$S = -k_B \sum_j P_j \ln P_j$$

$$dS = -k_B d \sum_j P_j \ln P_j$$

$$d\bar{E} = -\frac{1}{\beta} \sum_j \ln P_j dP_j + \sum_j P_j \left(\frac{\partial E_j}{\partial V} \right) dV$$

$$\beta = - \sum_j P_j \left(\frac{\partial E_j}{\partial V} \right) = -k_B \left(\sum_j \frac{P_j}{P_j} dP_j + \sum_j \ln P_j dP_j \right)$$

$$d\bar{E} = \left(\frac{1}{\beta} \right) \left(-\frac{dS}{k_B} \right) + \sum_j P_j \left(\frac{\partial E_j}{\partial V} \right) dV$$

$$d\bar{E} = \frac{1}{\beta k_B} dS + \sum_j P_j \left(\frac{\partial E_j}{\partial V} \right) dV \quad \text{--- (A)}$$

Thermodynamics \uparrow

$$dS = -k_B \sum_j \ln P_j dP_j$$

$$dE = T dS - p dV \quad \text{--- (B)} \Rightarrow T = \frac{1}{\beta k_B} \Rightarrow \beta = \frac{1}{k_B T}$$

$$Q = \sum_j e^{-\beta E_j}$$

$$\ln Q = \ln \left(\sum_j e^{-\beta E_j} \right)$$

$$\frac{\partial \ln Q}{\partial \beta} = \frac{1}{\sum_j e^{-\beta E_j}} \sum_j (-E_j) e^{-\beta E_j}$$

$$= - \frac{\sum_j E_j e^{-\beta E_j}}{\sum_j e^{-\beta E_j}} = -\bar{E}$$

$$\bar{E} = \sum_j P_j E_j \quad P_j = \frac{e^{-\beta E_j}}{Q}$$

$$\beta = \frac{1}{k_B T}$$

$$d\beta = -\frac{1}{k_B T^2} dT$$

$$\text{or } \frac{\partial}{\partial \beta} = -k_B T^2 \frac{\partial}{\partial T}$$

$$\bar{E} = - \frac{\partial \ln Q}{\partial \beta}$$

$$\bar{E} = k_B T^2 \left(\frac{\partial \ln Q}{\partial T} \right)_{N,V}$$

Internal energy

$$\frac{\partial \ln Q}{\partial V} = \frac{1}{\sum_j e^{-\beta E_j}} (-\beta) \sum_j e^{-\beta E_j} \left(\frac{\partial E_j}{\partial V} \right)$$

$$= \beta \left(- \frac{\sum_j e^{-\beta E_j} (\partial E_j / \partial V)}{Q} \right) = \beta p \leftarrow \text{Pressure}$$

$$p = k_B T \left(\frac{\partial \ln Q}{\partial V} \right)_{N,T}$$

$$S = -k_B \sum_j P_j \ln P_j$$

$$= -k_B \sum_j \frac{e^{-\beta E_j}}{Q} \ln \left(\frac{e^{-\beta E_j}}{Q} \right) = -k_B \sum_j \frac{e^{-\beta E_j}}{Q} (-\beta E_j - \ln Q)$$

$$= k_B \beta \left(\sum_j \frac{e^{-\beta E_j} E_j}{Q} \right) + k_B \ln Q$$

$$S = \frac{1}{T} \bar{E} + k_B \ln Q = k_B T \left(\frac{\partial \ln Q}{\partial T} \right)_{N,V} + k_B \ln Q$$

$$A = k_B T^2 \left(\frac{\partial \ln Q}{\partial T} \right)_{N,V} - k_B T^2 \left(\frac{\partial \ln Q}{\partial T} \right)_{N,V} - k_B T \ln Q$$

$$A = E - TS$$

↓
Helmholtz
free energy

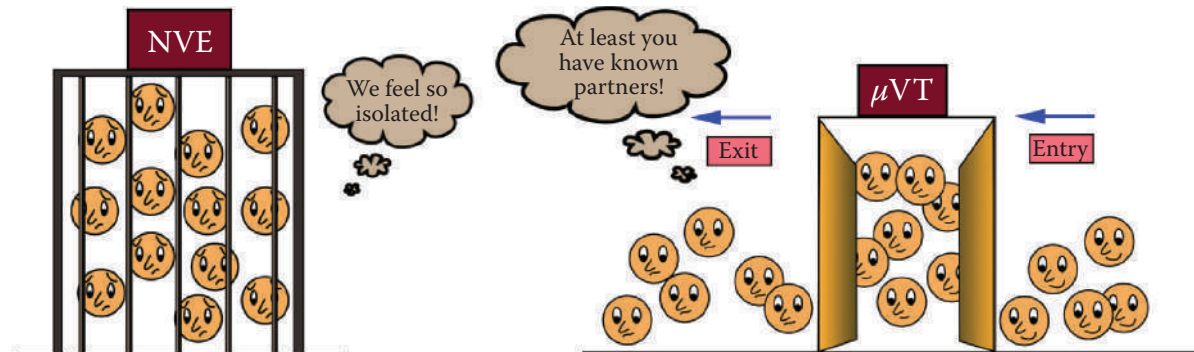
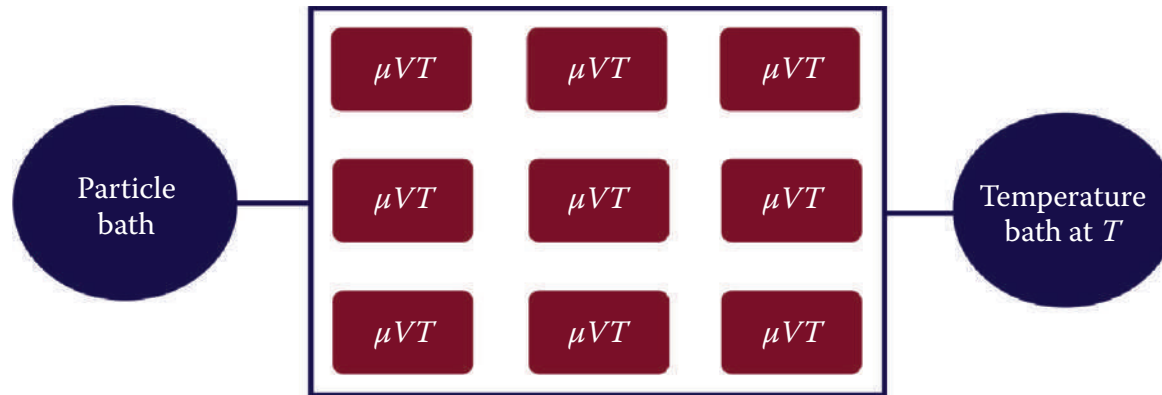
$$A = -k_B T \ln Q$$

→ Thermodynamic potential
in canonical ensemble

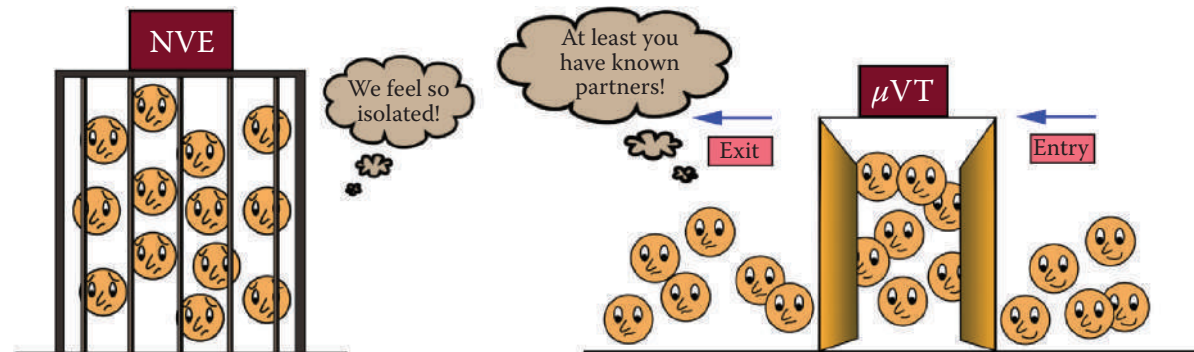
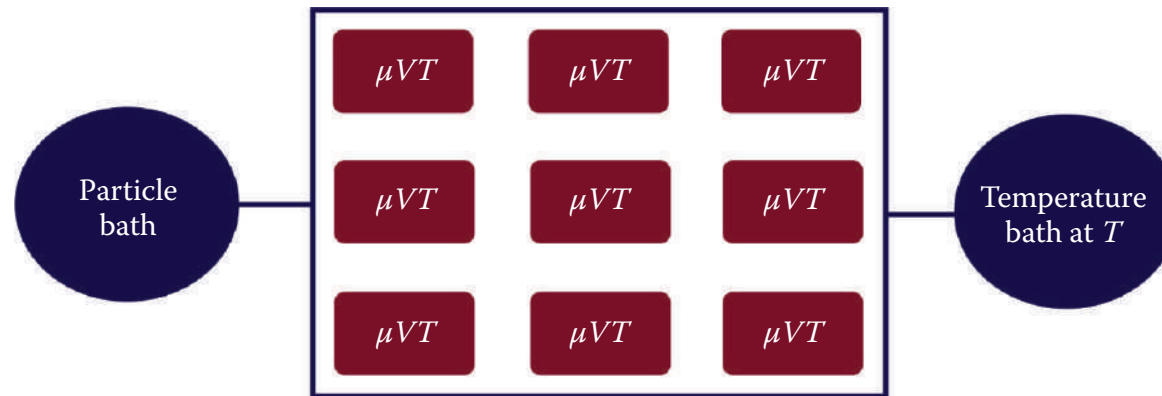
$$S = k_B \ln \Omega$$

↳
T.P. in
microcanonical
ensemble

Grand-canonical ensemble



Grand-canonical ensemble



$n_{Nj} \rightarrow$ No. of Systems in the ensemble
that contain N molecules/particles
that are in the state j

$\{n_{Nj}\}$ is a distribution

$$\sum_N \sum_j n_{Nj} = N_S \rightarrow \text{Total no. of Systems in the ensemble}$$

$$\sum_N \sum_j n_{Nj} E_{Nj} = E_t \rightarrow \text{Total energy of the ensemble}$$

$$\sum_N \sum_j n_{Nj} N = N^0 \rightarrow \text{Total no. of molecules (particles) in the ensemble}$$

No. of States

$$\Omega(\{n_{Nj}\}) = \frac{N_S!}{\prod_N \prod_j n_{Nj}!}$$

The distribution that maximizes $\Omega(\{n_{Nj}\})$ subjected to appropriate constraints completely dominates all others.

$$\ln \Omega(\{n_{Nj}\}) = \ln N_S! - \ln \prod_N \prod_j n_{Nj}!$$

$$\ln \Omega(\{n_{Nj}\}) = N_S \ln N_S - N_S - \sum_N \sum_j \ln n_{Nj}!$$

$$\ln \Omega = N_S \ln N_S - N_S - \sum_N \sum_j n_{Nj} \ln n_{Nj} + \sum_N \sum_j n_{Nj}$$

$$\ln \Omega(\{n_{Nj}\}) = N_S \ln N_S - \sum_N \sum_j n_{Nj} \ln n_{Nj} \quad N_S$$

$$\frac{\partial}{\partial n_{Nj}} \left[N_S \ln N_S - \sum_N \sum_j n_{Nj} \ln n_{Nj} - \alpha \sum_N \sum_j n_{Nj} - \beta \sum_N \sum_j n_{Nj} E_{Nj} - \gamma \left(\sum_N \sum_j n_{Nj} N \right) \right] = 0$$

$$\Rightarrow 0 - \frac{n_{Nj}}{n_{Nj}} - \ln n_{Nj}^* - \alpha - \beta E_{Nj} - \gamma N = 0$$

$$\ln n_{Nj}^* = -(1+\alpha) - \beta E_{Nj} - \gamma N = -\alpha' - \beta E_{Nj} - \gamma N$$

$$n_{Nj}^* = e^{-\alpha'} e^{-\beta E_{Nj}} e^{-\gamma N}$$

$$P_{Nj}(V, \beta, \gamma) = \frac{n_{Nj}^*}{N_S}$$

$$N_s = \sum_N \sum_i n_{Nj}^*$$

$$N_s = \sum_N \sum_j e^{-\alpha'} e^{-\beta E_{Nj}(V)} e^{-\gamma N} = e^{-\alpha'} \sum_N \sum_j e^{-\beta E_{Nj}(V)} e^{-\gamma N}$$

$$P_{Nj}(V, \beta, \gamma) = \frac{n_{Nj}^*}{N_s} = \frac{\cancel{e^{-\alpha'}} e^{-\beta E_{Nj}(V)} e^{-\gamma N}}{\cancel{e^{-\alpha'}} \sum_N \sum_j e^{-\beta E_{Nj}(V)} e^{-\gamma N}} = \frac{e^{-\beta E_{Nj}(V)} e^{-\gamma N}}{\Xi(V, \beta, \gamma)}$$

→ Grand Canonical partition fⁿ

$$\bar{E} = \langle E \rangle = \hat{E}(V, \beta, \gamma) = \frac{1}{\Xi} \sum_N \sum_j E_{Nj} e^{-\beta E_{Nj}(V)} e^{-\gamma N}$$

$$\ln \Xi = \ln \left(\sum_N \sum_j e^{-\beta E_{Nj}(V)} e^{-\gamma N} \right)$$

$$\left(\frac{\partial \ln \Xi}{\partial \beta} \right)_{V, \gamma} = \frac{1}{\sum_N \sum_j e^{-\beta E_{Nj}(V)} e^{-\gamma N}} \times \sum_N \sum_j (-E_{Nj}) e^{-\beta E_{Nj}(V)} e^{-\gamma N} = -\bar{E}(V, \beta, \gamma)$$

$$\bar{E} = \langle E \rangle = - \left(\frac{\partial \ln \Xi}{\partial \beta} \right)_{V, \gamma}$$

$$\left(\frac{\partial \ln \Xi}{\partial V} \right)_{\beta, \gamma} = \frac{1}{\sum_N \sum_j e^{-\beta E_{N,j}(V)} e^{-\gamma N}} \beta \sum_N \sum_j \left(-\frac{\partial E_{N,j}}{\partial V} \right) e^{-\beta E_{N,j}(V)} e^{-\gamma N} = \bar{p} = \langle p \rangle \rightarrow \text{Thermodynamic pressure}$$

↘ $\langle -\beta \frac{\partial E_{N,j}}{\partial V} \rangle$

$$\bar{N} = \langle N \rangle$$

$$\left(\frac{\partial \ln \Xi}{\partial \gamma} \right)_{\beta, V} = \frac{1}{\Xi} (-1) \sum_N \sum_j N e^{-\beta E_{N,j}(V)} e^{-\gamma N} = -\bar{N} = -\langle N \rangle$$

$$\bar{N} = \langle N \rangle = - \left(\frac{\partial \ln \Xi}{\partial \gamma} \right)_{\beta, V}$$

$$\beta = \frac{1}{k_B T} \quad (\text{can be shown})$$

$$\ln \Xi = f(\beta, \gamma, \{E_{Nj}(V)\}) = \ln \sum_N \sum_j e^{-\beta E_{Nj}(V)} e^{-\gamma N}$$

$$df = \left(\frac{\partial f}{\partial \beta} \right)_{\gamma, \{E_{Nj}\}} d\beta + \left(\frac{\partial f}{\partial \gamma} \right)_{\beta, \{E_{Nj}\}} d\gamma + \sum_N \sum_j \left(\frac{\partial f}{\partial E_{Nj}} \right)_{\beta, \gamma, E_{Nj}} dE_{Nj}$$

$$df = -\bar{E} d\beta - \bar{N} d\gamma - \beta \underbrace{\sum_N \sum_j P_{Nj} dE_{Nj}}_{\substack{\text{Ensemble average} \\ \text{reversible work} \\ \text{done by the system}}}$$

$$df = -\bar{E} d\beta - \bar{N} d\gamma + \beta \bar{P} dV$$

Adding $d(\beta \bar{E}) + d(\gamma \bar{N})$ to both sides of this eqⁿ.

$$d(f + \beta \bar{E} + \gamma \bar{N}) = -\cancel{\bar{E} d\beta} - \cancel{\bar{N} d\gamma} + \beta \bar{P} dV + \beta d\bar{E} + \cancel{\bar{E} d\beta} + \gamma d\bar{N} + \cancel{\bar{N} d\gamma}$$

$$d(f + \beta \bar{E} + \gamma \bar{N}) = \beta \bar{P} dV + \beta d\bar{E} + \gamma d\bar{N}$$

$$\ln \Xi = \sum_N \sum_j e^{-\beta E_{Nj}(V)} e^{-\gamma N}$$

$$\frac{\partial \ln \Xi}{\partial E_{Nj}} = \frac{1}{\Xi} (-\beta) e^{-\beta E_{Nj}(V)} e^{-\gamma N}$$

$$= -\beta P_{Nj}$$

$$d(f + \beta \bar{E} + \gamma \bar{N}) = \beta \bar{p} dV + \beta d\bar{E} + \gamma d\bar{N}$$

$$\beta = \frac{1}{k_B T}$$

$$T dS = d\bar{E} + \bar{p} dV - \mu d\bar{N} \quad (\text{thermodynamics}) \quad \dots \dots \textcircled{I}$$

$$k_B T d(f + \beta \bar{E} + \gamma \bar{N}) = \bar{p} dV + d\bar{E} + k_B T \gamma d\bar{N}$$

$$T d(k_B f + \frac{\bar{E}}{T} + k_B \gamma \bar{N}) = \bar{p} dV + d\bar{E} + k_B T \gamma d\bar{N} \quad \dots \textcircled{II}$$

Compare \textcircled{I} & \textcircled{II}

$$-\mu = k_B T \gamma \Rightarrow \gamma = -\mu / k_B T$$

$$S = k_B f + \frac{\bar{E}}{T} + k_B \bar{N} (-\mu / k_B T) = k_B \ln \Omega + \frac{\bar{E}}{T} - \frac{\bar{N} \mu}{T}$$

$$S = \frac{\bar{E}}{T} - \frac{\bar{N} \mu}{T} + k_B \ln \Omega \quad \rightsquigarrow \text{-Entropy in Grand Canonical ensemble-}$$

$$\Xi(V, T, \mu) = \sum_N \sum_j e^{-E_{Nj}(V)/k_B T} e^{\mu N/k_B T}$$

↳ grand canonical ensemble partition function

$$S = \frac{\bar{E}}{T} + k_B \ln \Xi - \frac{\mu \bar{N}}{T}$$

$$\Xi(V, T, \mu) = \sum_{N=0}^{\infty} Q(N, V, T) e^{\mu N/k_B T}$$

$$\Xi = \sum_{N=0}^{\infty} Q(N, V, T) \lambda^N$$

$$TS = \bar{E} + k_B T \ln \Xi - \mu \bar{N}$$

$$k_B T \ln \Xi = TS + \mu \bar{N} - \bar{E} = \bar{p} V$$

$$\Rightarrow \bar{p} V = k_B T \ln \Xi$$

$$Q(N, V, T) = \sum_j e^{-E_j/k_B T}$$

↳ canonical partition function

$$S = \frac{\bar{E}}{T} + k_B \ln Q$$

$$A = -k_B T \ln Q$$

$$\lambda = e^{\mu/k_B T}$$

$$G = \mu N = \bar{E} + \bar{p} V - TS$$

↳ Gibbs free energy $\Rightarrow \mu N + TS - \bar{E} = \bar{p} V$

Isothermal-Isobaric ensemble:

$$(N, p, T)$$

$$\Delta(N, T, p) = \sum_E \sum_V \Omega(N, V, E) e^{-E/k_B T - pV/k_B T} \leftarrow \text{"check"}$$

↓

Isothermal-Isobaric partition function

$$G = -k_B T \ln \Delta(N, T, p)$$

↳ Gibbs free energy

Canonical Ensemble
 (N, V, T)

$$\bar{E} \equiv \langle E \rangle = \frac{1}{Q(N, V, T)} \sum_j E_j e^{-\beta E_j}$$

$$\overline{E^2} = \langle E^2 \rangle = \frac{1}{Q(N, V, T)} \sum_j E_j^2 e^{-\beta E_j}$$

$$C_V(T) = \left(\frac{d\langle E \rangle}{dT} \right)_V$$

↳ Heat Capacity at constant Volume

Mean Square Standard deviation

$$\sigma_E^2 = \langle (E - \langle E \rangle)^2 \rangle$$

$$\sigma_E^2 = \langle E^2 \rangle - \langle E \rangle^2$$

$$\langle E^2 \rangle = \overline{E^2}$$

$$= \frac{1}{Q} \sum_j E_j^2 e^{-\beta E_j}$$

$$= -\frac{1}{Q} \frac{\partial}{\partial \beta} \underbrace{\sum_j E_j e^{-\beta E_j}}_{\langle E \rangle Q}$$

$$= -\frac{1}{Q} \frac{\partial}{\partial \beta} (\langle E \rangle Q)$$

$$= -\frac{1}{Q} Q \frac{\partial \langle E \rangle}{\partial \beta} - \frac{1}{Q} \langle E \rangle \frac{\partial Q}{\partial \beta}$$

$$= -\frac{\partial \langle E \rangle}{\partial \beta} - \langle E \rangle \left(\frac{\partial \ln Q}{\partial \beta} \right) \rightarrow -\langle E \rangle$$

$$\langle E^2 \rangle = -\frac{\partial \langle E \rangle}{\partial \beta} + \underbrace{\langle E \rangle \langle E \rangle}_{\langle E \rangle^2}$$

$$\Rightarrow \langle E^2 \rangle - \langle E \rangle^2 = \sigma_E^2 = -\frac{\partial \langle E \rangle}{\partial \beta}$$

$$\Rightarrow \sigma_E^2 = k_B T^2 \left(\frac{\partial \langle E \rangle}{\partial T} \right)_{N,V} = k_B T^2 C_V$$

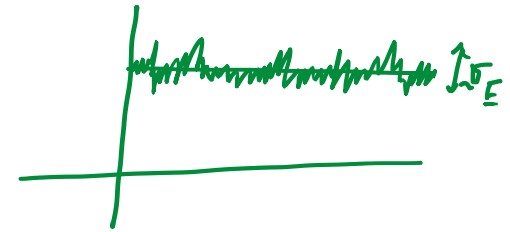
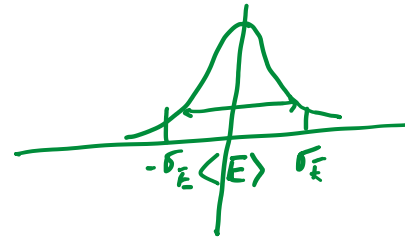
$$\Rightarrow C_V(k_B T) = \sigma_E^2$$

Heat capacity at constant volume.

$$\langle E \rangle = - \left(\frac{\partial \ln Q}{\partial \beta} \right)_{N,V}$$

$$\beta = \frac{1}{k_B T}$$

$\frac{\sigma_E}{\langle E \rangle} \sim$ How does it depend on
 the no. of particles ($\sim \frac{1}{\sqrt{N}}$)
 \downarrow
 Relative fluctuation



$$\frac{\sigma_E}{\langle E \rangle} \sim \frac{(k_B T^2 C_V)^{\frac{1}{2}}}{\langle E \rangle} \sim \frac{\sqrt{N}}{N} \sim \frac{1}{\sqrt{N}} \quad (\text{signature of equilibrium fluctuation})$$

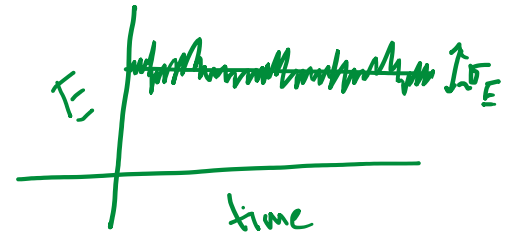
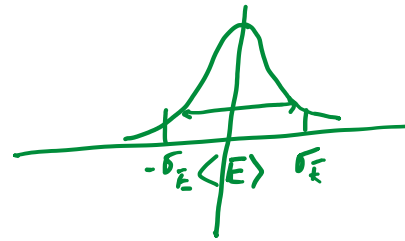
$\frac{\sigma_E}{\langle E \rangle} \sim$ How does it depend on
 the no. of particles ($\sim \frac{1}{\sqrt{N}}$)
 \downarrow
 Relative fluctuation

$$\frac{\sigma_E}{\langle E \rangle} \sim \frac{(k_B T^2 C_V)^{1/2}}{\langle E \rangle} \sim \frac{\sqrt{N}}{N} \sim \frac{1}{\sqrt{N}}$$

$C_V \rightarrow$ Heat capacity at constant vol^m
 (Response function)
 (N, V, T)

$$\Delta = \sum_{j, V} e^{-\beta(E_j + pV)}$$

Δ A partition fⁿ
 in (N, p, T) ensemble



(signature of equilibrium fluctuation)

$C_p \rightarrow$ Heat capacity at constant pressure
 (N, p, T)

$$k_B T^2 C_V = \sigma_E^2$$

$$\Delta(N, \beta, \tau) = \sum_{j,v} e^{-\beta(E_j + \epsilon v)}$$

$$\langle H \rangle = \frac{\sum_{j,v} (E_j + \epsilon v) e^{-\beta(E_j + \epsilon v)}}{\sum_{j,v} e^{-\beta(E_j + \epsilon v)}} \leftarrow \text{Average enthalpy} \quad \langle H \rangle = \langle E_j + \epsilon v \rangle$$

$$\frac{\partial \langle H \rangle}{\partial \beta} = \frac{\sum_{j,v} -(E_j + \epsilon v) e^{-\beta(E_j + \epsilon v)}}{\left(\sum_{j,v} e^{-\beta(E_j + \epsilon v)} \right)^2} - \frac{\sum_{j,v} (E_j + \epsilon v) e^{-\beta(E_j + \epsilon v)}}{\sum_{j,v} e^{-\beta(E_j + \epsilon v)}} \left/ \frac{d}{dx} \left(\frac{f(x)}{g(x)} \right) \right/ = \frac{f'(x)g(x) - g'(x)f(x)}{g(x)^2}$$

$$= \frac{-\sum_{j,v} (E_j + \epsilon v)^2 e^{-\beta(E_j + \epsilon v)}}{\sum_{j,v} e^{-\beta(E_j + \epsilon v)}}$$

$$= -\langle H^2 \rangle$$

$$+ \left(\frac{\sum_{j,v} (E_j + \epsilon v) e^{-\beta(E_j + \epsilon v)}}{\left(\sum_{j,v} e^{-\beta(E_j + \epsilon v)} \right)} \right)^2$$

↓

$$+ \langle H \rangle^2$$

$$\frac{\partial \langle H \rangle}{\partial \beta} = \langle H^2 \rangle - \langle H \rangle^2$$

$$\beta = \frac{1}{k_B T}$$

$$\frac{\partial}{\partial \beta} = -k_B T^2 \frac{\partial}{\partial T}$$

$$k_B T^2 \frac{\partial \langle H \rangle}{\partial T} = \langle H^2 \rangle - \langle H \rangle^2$$

$\underbrace{\hspace{1cm}}_{C_p} = \sigma_H^2$

$$k_B T^2 C_p = \sigma_H^2$$

Average Vol^m

$$\langle V \rangle = \frac{\sum_{j,v} v e^{-\beta(E_j + p v)}}{\sum_{j,v} e^{-\beta(E_j + p v)}}$$

$$\frac{\partial \langle V \rangle}{\partial p} = \frac{\sum_{j,v} v (-\beta v) e^{-\beta(E_j + p v)} \sum_{j,v} e^{-\beta(E_j + p v)} - \sum_{j,v} e^{-\beta(E_j + p v)} (-\beta V) \sum_{j,v} v e^{-\beta(E_j + p v)}}{\left(\sum_{j,v} e^{-\beta(E_j + p v)} \right)^2}$$

$$= (-\beta) \left[\frac{\sum_{j,v} v^2 e^{-\beta(E_j + p v)}}{\sum_{j,v} e^{-\beta(E_j + p v)}} - \left(\frac{\sum_{j,v} v e^{-\beta(E_j + p v)}}{\sum_{j,v} e^{-\beta(E_j + p v)}} \right)^2 \right] = -\beta [\langle v^2 \rangle - \langle v \rangle^2]$$

Isothermal Compressibility

$$\kappa = -\frac{1}{V} \left(\frac{\partial V}{\partial p} \right)_T \Rightarrow \boxed{\kappa = \frac{\sigma_v^2}{V k_B T}}$$

Canonical Ensemble

1 single particle \rightarrow Structureless (point particle)

N such particles (Non-interacting) \rightarrow vol^m V , temperature T

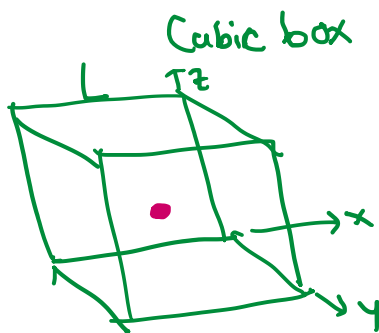
"ideal monatomic gas"

Single particle
 $q_{\text{trans}} = ?$

$$q_{\text{trans}} = \sum_j e^{-\beta E_j}$$

Only translational motion

How to calculate
 q_{trans} ?



$$q_{\text{trans}} = \sum_{n_x, n_y, n_z} e^{-\beta \frac{(n_x^2 + n_y^2 + n_z^2) \hbar^2}{8mL^2}}$$

$$= \sum_{n_x} e^{-\beta \frac{n_x^2 \hbar^2}{8mL^2}} \sum_{n_y} \dots \sum_{n_z} \dots$$

(N, V, T)

$$Q(N, V, T) = \sum_j e^{-\beta E_j^a} \cdot \sum_j e^{-\beta E_j^b} \dots$$

\leftarrow particle index

If the particles are distinguishable

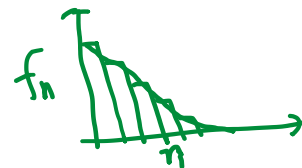
$$Q(N, V, T) = q_{\text{trans}}^a q_{\text{trans}}^b \dots$$

$$Q(N, V, T) = (q_{\text{trans}})^N$$

$m \Rightarrow$ mass of the particle

$$e^{abc} = e^a e^b e^c$$

$$q_{\text{trans}} = \sum_{n_x} e^{-\frac{\beta n_x^2 h^2}{8mL^2}} \sum_{n_y} e^{-\frac{\beta n_y^2 h^2}{8mL^2}} \sum_{n_z} e^{-\frac{\beta n_z^2 h^2}{8mL^2}} = \left(\sum_n e^{-\frac{n^2 h^2 \beta}{8mL^2}} \right)^3$$



$$= \left\{ \int_0^{\infty} dn e^{-\frac{n^2 h^2 \beta}{8mL^2}} \right\}^3$$

$$= \left\{ \int_0^{\infty} dn e^{-\frac{n^2 h^2}{8mL^2 k_B T}} \right\}^3$$

$$q_{\text{trans}} = \left\{ \frac{1}{2} \int_{-\infty}^{\infty} \dots \right\}^3 = \left\{ \frac{1}{2} \left(\frac{\pi 8mL^2 k_B T}{h^2} \right)^{\frac{1}{2}} \right\}^3$$

Then C_V (per particle)

$$C_V = \frac{\partial \langle E \rangle}{\partial T}$$

Equipartition theorem

$$\langle E \rangle = k_B T^2 \frac{\partial \ln q_{\text{trans}}}{\partial T} \quad \text{independent of } T$$

$$= k_B T^2 \frac{\partial}{\partial T} \left[\ln T^{\frac{3}{2}} + \dots \right]$$

$$\langle E \rangle = k_B T^2 \frac{3}{2} \frac{1}{T} = \frac{3}{2} k_B T$$

Recall

$$\int_{-\infty}^{+\infty} e^{-\alpha x^2} dx = (\pi/\alpha)^{\frac{1}{2}}$$

$$8^{\frac{1}{2}} = (2 \cdot 2 \cdot 2)^{\frac{1}{2}} = 2\sqrt{2} = 2^{\frac{3}{2}}$$

$V \rightarrow L^3$ (vol of cubic box)

$$q_{\text{trans}} = \frac{1}{8} 8^{\frac{3}{2}} \left(\frac{\pi m k_B T}{h^2} \right)^{\frac{3}{2}} V$$

$$q_{\text{trans}} = \left(\frac{2\pi m k_B T}{h^2} \right)^{\frac{3}{2}} V$$

$$Q(N, V, T) = q_{\text{trans}}^N$$

$$Q = V^N \left(\frac{2\pi m k_B T}{h^2} \right)^{\frac{3N}{2}}$$

$$\langle E \rangle = \frac{3}{2} N k_B T$$

$$C_V = \frac{\partial}{\partial T} \langle E \rangle = \frac{3}{2} N k_B$$

$$\text{if } N = N_A \quad C_V = \frac{3}{2} R$$

Entropy:

$$S = k_B \ln Q + k_B T \left(\frac{\partial \ln Q}{\partial T} \right)_{N, V}$$

$$= k_B \ln(q_{\text{trans}}^N) + k_B T \frac{\partial}{\partial T} \ln(q_{\text{trans}}^N)$$

$$= N k_B \ln q_{\text{trans}} + N k_B T \frac{\partial}{\partial T} (\ln q_{\text{trans}})$$

$$= N k_B \ln q_{\text{trans}} + N k_B T \frac{3}{2} \frac{1}{T}$$

$$= N k_B \ln q_{\text{trans}} + \frac{3}{2} N k_B$$

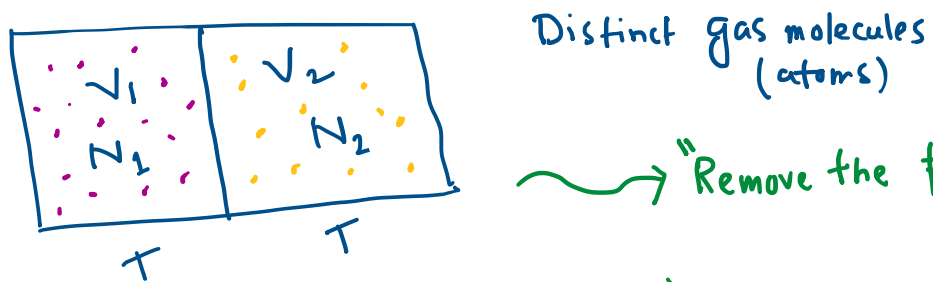
$$N k_B \ln(e^{3/2})$$

$$S = N k_B \ln \left[\left(\frac{2\pi m k_B T e}{h^2} \right)^{3/2} V \right] = N k_B \ln V$$

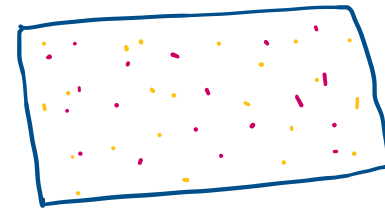
$$+ N k_B \ln \left[\left(\frac{2\pi m k_B T e}{h^2} \right)^{3/2} \right]$$

(say) σ

$$S = N k_B \ln V + N k_B \sigma$$



Remove the partition



"Mixing"

$\Delta S_{\text{mix}} > 0$ (should be positive)

When the partition is on:

$$S_i = S_1 + S_2 = N_1 k_B \ln V_1 + \sigma_1 N_1 k_B + N_2 k_B \ln V_2 + \sigma_2 N_2 k_B$$

↓

Entropy of
the initial
state

$$S_f = N_1 k_B \ln(V_1 + V_2) + N_2 k_B \ln(V_1 + V_2) + (N_1 \sigma_1 + N_2 \sigma_2) k_B$$

(after
removing
the partition)

$$\Delta S_{\text{mix}} = S_f - S_i$$

$$= N_1 k_B \ln(V_1 + V_2) + N_2 k_B \ln(V_1 + V_2) - N_1 k_B \ln V_1 - N_2 k_B \ln V_2$$

$$= N_1 k_B \ln\left(\frac{V}{V_1}\right) + N_2 k_B \ln\left(\frac{V}{V_2}\right) > 0$$

$$\begin{aligned} V &> V_1 \\ V &> V_2 \end{aligned}$$

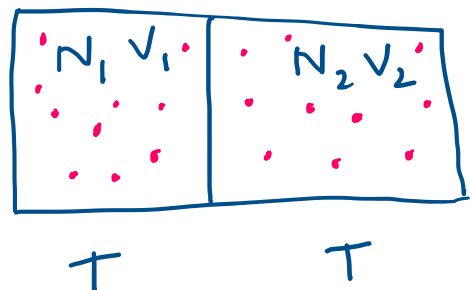
$$\left. \begin{aligned} V_1 + V_2 &= V \\ N_1 + N_2 &= N \end{aligned} \right\}$$

$$\Delta S_{\text{mix}} = -N k_B \left[\frac{N_1}{N} \ln\left(\frac{V_1}{V}\right) + \frac{N_2}{N} \ln\left(\frac{V_2}{V}\right) \right]$$

one can generalize the above expression

$$\Delta S_{\text{mix}} = -N k_B \sum_{\alpha} \left(\frac{N_{\alpha}}{N} \right) \ln\left(\frac{V_{\alpha}}{V}\right)$$

monatomic identical (indistinguishable)
ideal gas



$$\frac{N_1}{V_1} = \frac{N_2}{V_2} = \frac{N}{V}$$

$$\left. \begin{aligned} N &= N_1 + N_2 \\ V &= V_1 + V_2 \end{aligned} \right\}$$

~> Now we remove the partition

What do you expect in terms of
entropy (change)?

$$\Delta S_{\text{mix}} = 0 \quad \wedge \text{ (should be)}$$

But if we use the formula (we used few
minutes back)

$$\Delta S > 0$$

"Gibbs Paradox"

$$Q(N, V, T) = \frac{q_{\text{trans}}^N}{N!} \leftarrow \text{Divide by } N!$$

Non-interacting

indistinguishable
(identical)

particles

$$S = k_B \ln Q + k_B T \left(\frac{\partial \ln Q}{\partial T} \right)_{N, V}$$

$$S = k_B \ln \left(\frac{q_{\text{trans}}^N}{N!} \right) + k_B T \frac{\partial}{\partial T} \ln \left(\frac{q_{\text{trans}}^N}{N!} \right)$$

$$= N k_B \ln q_{\text{trans}} - k_B \ln N! + \frac{3}{2} N k_B$$

$$S = N k_B \ln q_{\text{trans}} - N k_B \ln N + N k_B + \frac{3}{2} N k_B$$

$$S = N k_B \ln \left[\left(\frac{2\pi m k_B T}{h^2} \right)^{3/2} \left(\frac{V}{N} \right) \right] + \frac{5}{2} N k_B$$

$$S = N k_B \ln \left(\frac{V e}{N} \right) + N k_B \underbrace{\ln \left[\left(\frac{2 \pi m k_B T}{h^2} \right)^{3/2} \right]}_{\sigma}$$

$$S = N k_B \ln \left(\frac{V e}{N} \right) + N k_B \sigma$$

$$V = V_1 + V_2$$

$$N = N_1 + N_2$$

$$S_i = N_1 k_B \ln \left(\frac{V_1 e}{N_1} \right) + N_1 k_B \sigma + N_2 k_B \ln \left(\frac{V_2 e}{N_2} \right) + N_2 k_B \sigma$$

$$S_f = N_1 k_B \ln \left(\frac{V e}{N_1} \right) + N_1 k_B \sigma + N_2 k_B \ln \left(\frac{V e}{N_2} \right) + N_2 k_B \sigma$$

(if distinct gas particles) (particle type 1 has available vol^m $V_1 + V_2 = V$)

$$\Delta S_{mix} = S_f - S_i = N_1 k_B \ln \left(\frac{V}{N_1} \cdot \frac{N_1}{V_1} \right) + N_2 k_B \ln \left(\frac{V}{N_2} \cdot \frac{N_2}{V_2} \right)$$

$$= -N k_B \left[\frac{N_1}{N} \ln \left(\frac{V_1}{V} \right) + \frac{N_2}{N} \ln \left(\frac{V_2}{V} \right) \right] > 0$$

But particles are identical !! \checkmark

$$S_f = (N_1 + N_2) k_B \ln \left(\frac{e \overbrace{(V_1 + V_2)}^V}{\underbrace{(N_1 + N_2)}_N} \right) + N_1 k_B \sigma_1 + N_2 k_B \sigma_2$$

$$\sigma_1 = \sigma_2$$

↓

$$S_i = N_1 k_B \ln \left(\frac{V_1 e}{N_1} \right) + N_1 k_B \sigma_1 + N_2 k_B \ln \left(\frac{V_2 e}{N_2} \right) + N_2 k_B \sigma_2$$

$$\Delta S_{mix}$$

$$\hookrightarrow = N_1 k_B \ln \left(\frac{eV}{N} \cdot \frac{N_1}{eV_1} \right) + N_2 k_B \ln \left(\frac{eV}{N} \cdot \frac{N_2}{eV_2} \right)$$

Now

$$\frac{N_1}{V_1} = \frac{N_2}{V_2} = \frac{N}{V} \quad (\text{density is uniform})$$

$$\Delta S_{mix} = N_1 k_B \ln \left(\underbrace{\left(\frac{V}{N} \cdot \frac{N_1}{V_1} \right)}_1 \right) + N_2 k_B \ln \left(\underbrace{\left(\frac{V}{N} \cdot \frac{N_2}{V_2} \right)}_1 \right) = 0$$

(What we expect)

N - non-interacting, indistinguishable (identical) particles/molecules

$$Q = \frac{q^N}{N!}$$

q = Molecular partition function

Canonical ensemble

Non-interacting but distinguishable

Total energy of the system

$$E_{\lambda}(N, V) = \epsilon_1^a(V) + \epsilon_j^b(V) + \dots + \epsilon_k^c(V) + \dots$$

particle no.

N such terms

Energy levels/particle

Total partition fn
of the system

$$q(V, T) = \sum_i e^{-\beta \epsilon_i}$$

Molecular p. fn

$$Q(N, V, T) = \sum_{\lambda} e^{-\beta E_{\lambda}} = \sum_{i, j, k} e^{-\beta (\epsilon_i^a(V) + \epsilon_j^b(V) + \dots)}$$

Sum is all possible configurations / states

$$= \sum_i e^{-\beta \epsilon_i^a} \sum_j e^{-\beta \epsilon_j^b} \sum_k e^{-\beta \epsilon_k^c} = q_a(V, T) q_b(V, T) q_c(V, T) \dots$$

particles are distinguishable

Non-interacting (independent) indistinguishable particles:

Two non-interacting Fermions (example: electron) $\rightarrow \epsilon_1, \epsilon_2, \epsilon_3, \epsilon_4 \rightarrow$ Available particle energy levels

$$Q(2, V, T) = \sum_{i,j=1}^4 e^{-\beta(\epsilon_i + \epsilon_j)}$$

\hookrightarrow 16 terms in this sum

* Additional restriction

\rightarrow for a given ϵ_i , occupation no. is 0 or 1

But 4 terms are unacceptable \rightarrow As these are Fermions

12 terms seem acceptable

$$\begin{array}{cc} \epsilon_1 + \epsilon_2 & \epsilon_2 + \epsilon_3 \\ \epsilon_1 + \epsilon_3 & \epsilon_2 + \epsilon_4 \\ \epsilon_1 + \epsilon_4 & \epsilon_3 + \epsilon_4 \end{array}$$

6

$$\begin{array}{cc} \epsilon_2 + \epsilon_1 & \epsilon_3 + \epsilon_2 \\ \epsilon_3 + \epsilon_1 & \epsilon_4 + \epsilon_2 \\ \epsilon_4 + \epsilon_1 & \epsilon_4 + \epsilon_3 \end{array}$$

6

Are they acceptable

As these particles are indistinguishable

$$\frac{12}{2!} = 6 \text{ terms}$$

Bosons? no. restriction on occupation no.
But they are still indistinguishable

In general
Collection of N particles

Position of ϵ_2 is not important
as particles are indistinguishable

ϵ_2 can be located in any of the
 N positions

$$E = \epsilon_1 + \epsilon_2 + \epsilon_3 + \dots + \epsilon_N$$

$$E = \underbrace{\epsilon_2 + \epsilon_{10} + \epsilon_{10} + \epsilon_{10} + \dots + \epsilon_{10}}_{N \text{ particles, hence } N \text{ terms}}$$

$$E = \underbrace{\epsilon_{10} + \epsilon_2 + \epsilon_{10} + \epsilon_{10} + \dots + \epsilon_{10}}_{N \text{ terms}}$$

↓
It represents the
same state

$N!$ possible permutations !!

$$Q = \frac{q^N}{N!} \quad (\text{to restrict overcounting})$$

Electronic partition function:

$$q_{elec} = \sum_i (\omega_{e_i}) e^{-\beta E_i}$$

ω_{e_i} → Degeneracy

$$q_{elec} = \omega_{e_1} e^0 + \omega_{e_2} e^{-\beta \underbrace{(E_2 - E_1)}_{\Delta E_{12} = E_2}} + \dots$$

$E_1 = 0$ (by definition)

Fix the arbitrary zero of energy

Say the lowest electronic state
≡ zero state

He atom

fraction of atoms in the lowest triplet state $3S_1$ is given by
(first excited state)

$$f_2 = \frac{\omega_{e_2} e^{-\beta \Delta E_{12}}}{\omega_{e_1} e^{-0} + \omega_{e_2} e^{-\beta \Delta E_{12}} + \dots} = \frac{3 e^{-\beta \Delta E_{12}}}{\omega_{e_1} + 3 e^{-\beta \Delta E_{12}} + \omega_{e_3} e^{-\beta \Delta E_{13}} + \dots}$$

At 300K, $\beta \Delta E_{12} \approx 770 \rightarrow f_2 \sim 10^{-334}$ (almost zero) 3000K, $f_2 \sim 10^{-33}$ (still small)

A good approximation

$$q_{elec} = \omega_{e1} + \underbrace{\omega_{e2} e^{-\beta \Delta E_{12}}}_{\substack{\text{may or} \\ \text{may not} \\ \text{be required}}}$$

$$q_{elec} \approx \omega_{e1} + \omega_{e2} e^{-\beta \Delta E_{12}}$$

F atoms

$T(K)$	f_2
200	0.027
400	0.105
:	
2000	0.272

Nuclear partition function:

Nuclear energy levels are separated by millions of electron volts

$$q_n = \omega_{n1} \quad \left(\begin{array}{l} \text{ground state (nuclear) energy is taken to be zero} \\ \rightarrow \text{Many times omitted} \end{array} \right)$$

Ideal Monatomic Gas

$$Q = \frac{(q_{trans} q_{elec} q_{nuc})^N}{N!}$$

$$q_{trans} = \left(\frac{2\pi m k_B T}{h^2} \right)^{3/2} V$$

$$q_{trans} = \frac{V}{\lambda^3}$$

Helmholtz free energy:

$$A = -k_B T \ln Q = -k_B T \ln \left(\frac{q_{\text{trans}} q_{\text{elec}}}{N!} \right)^N$$

$$\ln \left(\frac{1}{N!} \right) \approx -\ln N! = -N \ln N + N \\ = -(N \ln N - N) \rightarrow$$

$$= -N k_B T \ln \left[\left(\frac{2\pi m k_B T}{h^2} \right)^{3/2} \frac{V e}{N} \right] - \underbrace{N k_B T \ln (w_1 + w_2 e^{-\beta \Delta E_{12}})}_{\text{Small contribution}}$$

$$q_{\text{elec}} = w_1 + w_2 e^{-\beta \Delta E_{12}} \Rightarrow \ln q_{\text{elec}} = \ln (\dots)$$

Thermodynamic (internal) energy:

$$E = k_B T^2 \left(\frac{\partial \ln Q}{\partial T} \right)_{N,V} = + N k_B T^2 \left(\frac{3}{2} \right) \frac{1}{T} + \underbrace{\text{Electronic part}}$$

$$\beta = \frac{1}{k_B T}$$

$$N k_B T^2 \frac{1 (0 + w_2 (-\Delta E_{12}) e^{-\beta \Delta E_{12}} (-\frac{1}{k_B T^2}))}{(w_1 + w_2 e^{-\beta \Delta E_{12}})}$$

$$E = \frac{3}{2} N k_B T + \underbrace{\frac{N w_2 \Delta E_{12} e^{-\beta \Delta E_{12}}}{q_{\text{elec}}}}_{\text{Very small contribution}}$$

Entropy

$$S = k_B \ln Q + k_B T \left(\frac{\partial \ln Q}{\partial T} \right)_{N, V}$$

$$S = \frac{3}{2} N k_B + N k_B \ln \left[\left(\frac{2 \pi m k_B T}{h^2} \right)^{3/2} \frac{V e}{N} \right] + N k_B \ln (w_{e1} + w_{e2} e^{-\beta \Delta \epsilon_{12}}) + \frac{N k_B w_{e2} \beta \Delta \epsilon_{12} e^{-\beta \Delta \epsilon_{12}}}{q_{elec}}$$

"Sackur-Tetrode
equation"

S_{elec}

Contribution from
electronic part
of the partition
function.

The Chemical potential

$$q = q_{\text{trans}} q_{\text{elec}} q_n$$

$$\mu(T, p) = -k_B T \left(\frac{\partial \ln Q}{\partial N} \right)_{V, T} = -k_B T \frac{\partial}{\partial N} \left(\ln \left(\frac{q^N}{N!} \right) \right)$$

$$= -k_B T \frac{\partial}{\partial N} [N \ln q - \ln N!]$$

$$= -k_B T \frac{\partial}{\partial N} [N \ln q - N \ln N + N]$$

$$\mu(T, p) = -k_B T [\ln q - \cancel{N/N} - \ln N + \cancel{1}] = -k_B T \ln \left(\frac{q}{N} \right)$$

$$\mu(T, p) = -k_B T \ln \left[\left(\frac{2\pi m k_B T}{h^2} \right)^{3/2} \left(\frac{V}{N} \right) \right] - k_B T \ln q_{\text{elec}} q_n$$

$$\mu(T, p) = -k_B T \ln \left[\left(\frac{2\pi m k_B T}{h^2} \right)^{3/2} k_B T \right] - \frac{k_B T}{p} \ln q_{\text{elec}} q_n + k_B T \ln p$$

$$\underbrace{\quad}_{\mu_0(T)}$$

$$\mu(T, p) = \mu_0(T) + k_B T \ln p$$

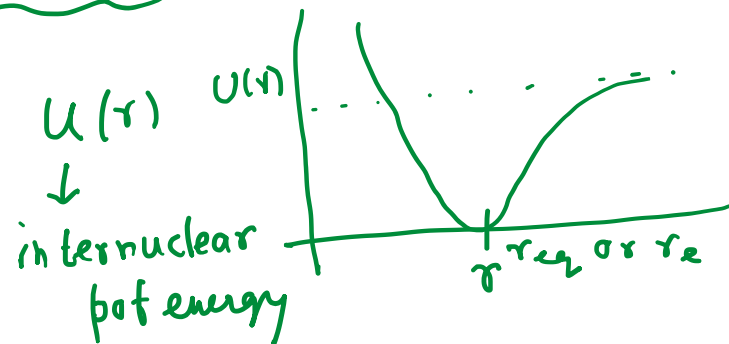
Ideal (non-interacting) diatomic gas:

Rigid Rotor - Harmonic Oscillator approximation

Born-Oppenheimer Approximation
fixed nuclei (heavy)



$$\frac{m_1 + m_2}{\text{Total mass}}$$



$$H = H_{\text{trans}} + H_{\text{int (internal)}}$$

↓ Hamiltonian for the motion of the whole body Relative co-ordinate

$$E = E_{\text{trans}} + E_{\text{int}}$$

$$U(r) = U(r_e) + (r - r_e) \left(\frac{\partial U}{\partial r} \right)_{r=r_e} + \frac{1}{2} \left(\frac{\partial^2 U}{\partial r^2} \right)_{r=r_e} (r - r_e)^2 + \dots$$

↓
0

$$q_{\text{trans}} = \left[\left(\frac{2\pi (m_1 + m_2) k_B T}{h^2} \right)^{3/2} \right] V$$

(Diatomic)

$$\frac{1}{2} \left(\frac{\partial^2 U}{\partial r^2} \right)_{r=r_e} (r - r_e)^2 + \dots$$

$$U(r) \approx U(r_e) + \frac{k}{2} (r - r_e)^2 + \dots$$

harmonic

(Taylor expansion)
 $k = \text{force constant}$

Anharmonic (not included)

for the internal motion

$$H_{int} = H_{rot, vib} = H_{rot} + H_{vib}$$

↓
Rigid
Rotor

↘ Simple Harmonic
Oscillator

$$E_{rot, vib} = E_{rot} + E_{vib}$$

$$q_{rot, vib} = q_{rot} q_{vib}$$

Let us consider vibration first

$$\nu = \frac{1}{2\pi} \left(\frac{k}{\mu} \right)^{\frac{1}{2}}$$

↓
 1000 cm^{-1}

$$\mu = \frac{m_1 m_2}{m_1 + m_2}$$

↓ Reduced mass

$$E_{vib} = \left(n + \frac{1}{2} \right) h \nu$$

$n = 0, 1, 2, \dots$
vibrational
quantum
number

$$k \sim 10^5 - 10^6 \text{ dynes/cm}$$

The vibrational partition function:

$$E_n = \left(n + \frac{1}{2}\right) h\nu$$

$$q_{\text{vib}}(T) = \sum_{n=0}^{\infty} e^{-\beta \left(n + \frac{1}{2}\right) h\nu} = e^{-\frac{\beta h\nu}{2}} \sum_{n=0}^{\infty} e^{-n\beta h\nu}$$

Let us assume $y = \sum_{n=0}^{\infty} e^{-n\beta h\nu} = (1 + e^{-\beta h\nu} + e^{-2\beta h\nu} + e^{-3\beta h\nu} + \dots)$

Let, $e^{-\beta h\nu} = x$

$$y = 1 + x + x^2 + x^3 + \dots$$
$$= 1 + x(1 + x + x^2 + \dots)$$

$$y = 1 + xy$$

$$\Rightarrow y(1-x) = 1 \Rightarrow y = \frac{1}{1-x}$$

$$q_{\text{vib}} = e^{-\beta h\nu/2} \left(\frac{1}{1 - e^{-\beta h\nu}} \right)$$

If temperature is high $\beta h\nu \ll 1$

$$q_{\text{vib}}(T) \approx \int_0^{\infty} e^{-\beta h\nu/2} e^{-n\beta h\nu} dn = e^{-\beta h\nu/2} \int_0^{\infty} e^{-n\beta h\nu} dn$$

$$= e^{-\beta h\nu/2} \left. \frac{e^{-\beta h\nu n}}{(-\beta h\nu)} \right|_0^{\infty}$$

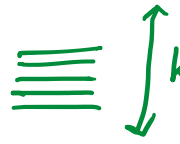
$$= e^{-\beta h\nu/2} \left(\frac{e^{-\infty} - e^0}{(-\beta h\nu)} \right) = \frac{e^{-\beta h\nu/2}}{\beta h\nu}$$

$$q_{\text{vib}}(T) \approx \frac{1}{\beta h\nu} = \left(\frac{k_B T}{h\nu} \right)$$

$$q_{\text{vib}}(T) = \frac{e^{-\beta h\nu/2}}{1 - e^{-\beta h\nu}}$$

$$\beta h\nu \ll 1 \quad \approx \frac{1 - \cancel{\beta h\nu/2}^0}{1 - 1 + \beta h\nu} \approx \frac{1}{\beta h\nu} = \left(\frac{k_B T}{h\nu} \right)$$

$\frac{h\nu}{k_B T}$ is small



$e^{-\beta h\nu/2} \approx 1$

Vibrational Contribution to the internal energy:

$$E_{\text{vib}} = N k_B T^2 \frac{d \ln q_{\text{vib}}}{dT} = -N \frac{d \ln q_{\text{vib}}}{d\beta}$$

$$\ln q_{\text{vib}} = -\beta h\nu/2 - \ln(1 - e^{-\beta h\nu})$$

$$q_{\text{vib}} = \frac{e^{-\beta h\nu/2}}{1 - e^{-\beta h\nu}}$$

$$\frac{d \ln q_{\text{vib}}}{d\beta} = -h\nu/2 - \frac{1}{(1 - e^{-\beta h\nu})} (0 + h\nu e^{-\beta h\nu})$$

$$-N \frac{d \ln q_{\text{vib}}}{d\beta} = +\frac{N h\nu}{2} + \frac{N h\nu e^{-\beta h\nu}}{1 - e^{-\beta h\nu}} = E_{\text{v}}$$

$$E_{\text{v}} = N k_B \left(\frac{\Theta_{\text{v}}}{2} + \frac{\Theta_{\text{v}}}{e^{\Theta_{\text{v}}/T} - 1} \right)$$

Let, $\Theta_{\text{v}} = \frac{h\nu}{k_B}$
↓
vibrational temp

$$\beta h\nu = \frac{\Theta_{\text{v}}}{T}$$

$C_{V,\text{vib}} \rightarrow \text{Next}$

$$E_v = Nk_B \left(\frac{\Theta_v}{2} + \frac{\Theta_v}{e^{\Theta_v/T} - 1} \right)$$

Vibrational Contribution to the heat Capacity

$$C_{v,vib} = \left(\frac{\partial E_v}{\partial T} \right) N, v$$

$$\left(\frac{\partial E_v}{\partial T} \right) N, v = Nk_B \times 0 + Nk_B \frac{(0 - (e^{\Theta_v/T} (-\frac{\Theta_v}{T^2}) - 0))}{(e^{\Theta_v/T} - 1)^2}$$

$$C_{v,vib} = Nk_B \left(\frac{\Theta_v}{T} \right)^2 \frac{e^{\Theta_v/T}}{(e^{\Theta_v/T} - 1)^2}$$

$$C_{v,vib} \approx (Nk_B) \left(\frac{\Theta_v}{T} \right)^2 \left(\frac{1 + \Theta_v/T + \dots}{(1 + \Theta_v/T + \dots - 1)^2} \right)$$

$$\approx (Nk_B) \left(\frac{\Theta_v}{T} \right)^2 \left(\frac{T}{\Theta_v} \right)^2 (1 + \dots) = Nk_B \quad (\text{checked})$$

$$\frac{d}{dx} \left(\frac{f(x)}{g(x)} \right) = \frac{f'(x)g(x) - g'(x)f(x)}{g(x)^2}$$

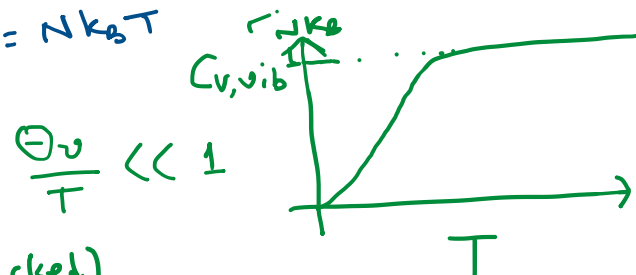
$$\left| \begin{aligned} E_v &= Nk_B \left(\frac{\Theta_v}{2} + \frac{\Theta_v}{1 + \Theta_v/T - 1 \dots} \right) \\ T \rightarrow \infty &= Nk_B \left(\frac{\Theta_v}{2} + \frac{\cancel{\Theta_v}}{\cancel{\Theta_v} T} \right) \\ &\approx Nk_B T \quad T \gg \Theta_v \\ &(\text{checked}) \end{aligned} \right.$$

What happens at high temperature?

Equipartition theorem (classical)

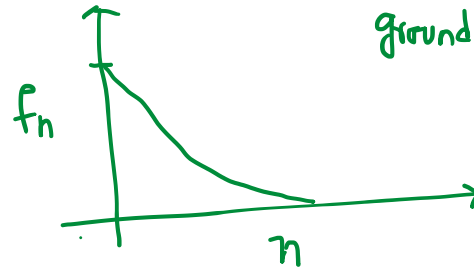
$$C_{v,vib} = Nk_B \quad T \rightarrow \infty$$

$$E_{vib} = Nk_B T$$



Fraction of molecules in the excited vibration state(s)

In state "n"
$$f_n = \frac{e^{-\beta h \nu (n + \frac{1}{2})}}{q_{vib}(T)}$$



Most molecules are in the ground vibrational state

$$\Theta_v = \frac{h\nu}{k_B}$$
 ← Molecular property

$$f_{n>0} = \sum_{n=1}^{\infty} \frac{e^{-\beta h \nu (n + \frac{1}{2})}}{q_{vib}(T)}$$

$$= \frac{e^{-\beta h \nu / 2}}{q_{vib}(T)} \sum_{n=1}^{\infty} e^{-\beta h \nu n}$$

$$= \frac{e^{-\beta h \nu / 2}}{q_{vib}(T)} \left(1 + \sum_{n=1}^{\infty} e^{-\beta h \nu n} - 1 \right)$$

$$= \frac{e^{-\beta h \nu / 2}}{q_{vib}(T)} \left(\frac{1}{1 - e^{-\beta h \nu}} - 1 \right)$$

$$= \frac{1}{q_{vib}(T)} \left(\frac{e^{-\beta h \nu / 2}}{1 - e^{-\beta h \nu}} - e^{-\beta h \nu / 2} \right)$$

Θ_v	Molecule
H ₂	6215
HCl	4227
I ₂	310
Cl ₂	810

$$f_0 = \frac{e^{-\beta h \nu / 2}}{q_{vib}(T)}$$

$$1 - f_0 = 1 - \frac{e^{-\beta h \nu / 2}}{\frac{e^{-\beta h \nu / 2}}{1 - e^{-\beta h \nu}}}$$

$$= 1 - (1 - e^{-\beta h \nu})$$

$$f_{n>0} = e^{-\beta h \nu} = e^{-\Theta_v / T}$$

$$= 1 - \frac{e^{-\beta h \nu / 2}}{q_{vib}(T)} = (1 - f_0) =$$

Rotational partition function of a heteronuclear Diatomic:

$$q_{\text{rot}}(T) = \sum_{J=0}^{\infty} (2J+1) e^{-\beta J(J+1) \bar{B}}$$

$\frac{\bar{B}}{k_B} = \Theta_r \rightarrow$ characteristic temperature of rotation

Usually, $\frac{\Theta_r}{T}$ is quite small at ordinary temperatures

then above sum can be replaced by an integral

$$q_{\text{rot}}(T) = \int_0^{\infty} dJ (2J+1) e^{-\frac{\Theta_r J(J+1)}{T}}$$

$$= \int_0^{\infty} dy \left(\frac{T}{\Theta_r} \right) e^{-y} = \frac{T}{\Theta_r} \int_0^{\infty} e^{-y} dy = \frac{T}{\Theta_r} \left. \frac{e^{-y}}{-1} \right|_0^{\infty} = \frac{T}{\Theta_r} (-0 + 1)$$

$$q_{\text{rot}}(T) = \frac{T}{\Theta_r} \quad \frac{\Theta_r}{T} \ll 1$$

"High temperature limit"

For low temperatures (molecules with high value of θ_r , HD, $\theta_r = 42.7\text{K}$)

One cannot use integration to evaluate the partition function

$$q_{\text{rot}}(T) = (1 + 3e^{-2\theta_r/T} + 5e^{-6\theta_r/T} + \dots)$$

Euler-Maclaurin Summation
formula

$$\frac{\theta_r}{T} \gg 1$$

$$\frac{\theta_r}{T} \ll 1$$

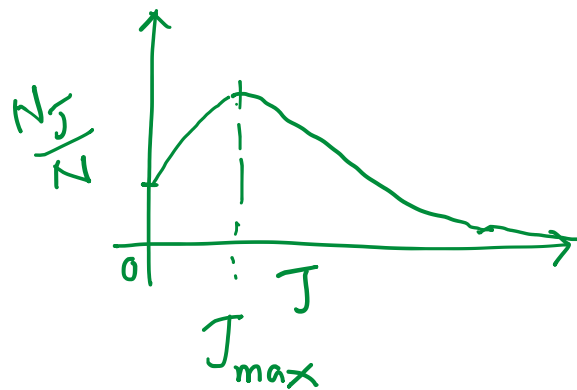
$$E_{\text{rot}} = N \frac{\partial \ln q_{\text{rot}}(T)}{\partial T} (k_B T^2) = N (k_B T^2) \frac{\partial}{\partial T} (\ln(T/\theta_r))$$

$$E_{\text{rot}} = N k_B T^2 \left(\frac{1}{T} \right) = N k_B T \quad (\text{equipartition theorem})$$

$$C_{V,\text{rot}} = N k_B$$

$$\frac{N_J}{N} = \frac{(2J+1)e^{-\theta_r J(J+1)/T}}{q_{\text{rot}}(T)}$$

↳ A plot of this vs J



J_{max} ? How to calculate?

$$\frac{\partial}{\partial J} \left(\frac{N_J}{N} \right) = 0$$

$$\frac{1}{q_{\text{rot}}(T)} \frac{\partial}{\partial J} \left((2J+1)e^{-\theta_r J(J+1)/T} \right) = 0$$

$$(2 \cdot 1 + 0) e^{-\theta_r J(J+1)/T} + (2J+1) e^{-\theta_r J(J+1)/T} \left(-\frac{\theta_r}{T} \right) (2J+1) = 0$$

$$J_{\text{max}} = \left(\frac{T}{2\theta_r} \right)^{1/2} - \frac{1}{2}$$

$$= \left(\frac{T}{2B/k_B} \right)^{1/2} - \frac{1}{2} \approx \left(\frac{k_B T}{2B} \right)^{1/2}$$

$$2 + (2J+1) \left(-\frac{\theta_r}{T} \right) = 0$$

$$\Rightarrow (2J+1)^2 \left(\frac{\theta_r}{T} \right) = 2$$

$$\Rightarrow (2J+1) = \left(\frac{2T}{\theta_r} \right)^{1/2}$$

$$J_{\text{max}} = \frac{1}{2} \left(\frac{2T}{\theta_r} \right)^{1/2} - \frac{1}{2}$$

Homonuclear Diatomic Molecule (partition f^n for rotation):

$$H_2 \quad \psi_{total} = \psi_{trans} \psi_{rot} \psi_{vib} \psi_{elec} \psi_{nucl}$$

Let us do a two step operation

- ① An inversion of all the particles, electrons and nuclei, through the origin
- ② An inversion of just electrons back through the origin

This two step process is equivalent to an exchange of nuclei

Translational wavefunction: \rightarrow Not affected by inversion

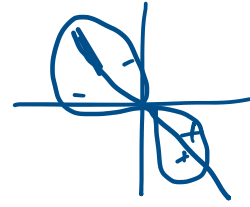
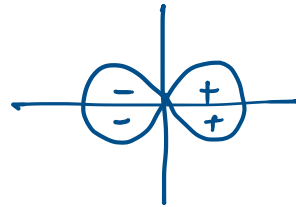
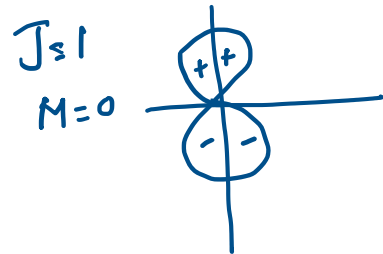
Electronic wavefunction: \rightarrow that depends on the symmetry of that electronic state

Ground electronic state \sum_g^+ (Symmetric under both the operations)

ψ_{vib} depends on \downarrow $(r - r_e)$ \rightarrow Unaffected by any inversion operation
magnitude
of

Rotation

rotational wavefunctions (Solve rigid-rotor problem) These wavefⁿs will get affected by inversion



$J = \text{even}$
 $J = \text{odd}$

$$(\tau, \theta, \phi) \rightarrow (\tau, \pi - \theta, \phi + \pi)$$

nuclear part of the wavefⁿ (nuclear spin)

H-H molecule
 $\uparrow \uparrow$

$\alpha\alpha$
 $\beta\beta$
 $\frac{1}{\sqrt{2}}(\alpha\beta + \beta\alpha)$

} Symmetric

$\frac{1}{\sqrt{2}}(\alpha\beta - \beta\alpha) \rightarrow \text{Antisymmetric}$

H₂ molecule in the ground electronic state Σ_g^+

In general

$I \rightarrow$ nuclei Spin

Each nuclei would have $(2I+1)$ states (Spin)

Diatomic molecule

Total no. of spin state
(Diatomic)

$$(2I+1) \quad (2I+1)$$

$$(2I+1)^2$$

$(2I+1)\left(\frac{2I}{2}\right) \rightarrow$ No. of antisymmetric spin states

$$\frac{1}{\sqrt{2}} (\alpha_i(1)\alpha_j(2) - \alpha_j(1)\alpha_i(2))$$

No. of Symmetric Spin states

$$i < j \quad j < 2I+1$$

$$(2I+1)^2 - (2I+1)\left(\frac{2I}{2}\right) = (I+1)(2I+1)$$

for H₂ $I = \frac{1}{2}$

$$\left(2 \cdot \frac{1}{2} + 1\right) \left(\frac{2 \cdot \frac{1}{2}}{2}\right) = \frac{2}{2} = 1$$

$$\left(\frac{1}{2} + 1\right) \left(2 \cdot \frac{1}{2} + 1\right) = 3$$

Integral Spins: (Total wavefn has to be symmetric)
 nuclear

(Bosons)

$I(2I+1)$ Antisymmetric spin f's combine with J odd

$(I+1)(2I+1)$ Symmetric nuclear spin f's combine with J even

$$\rightarrow q_{\text{rot,nuc}}(T) = (I+1)(2I+1) \sum_{J \text{ even}} (2J+1) e^{-J(J+1)\theta_r/T} + I(2I+1) \sum_{J \text{ odd}} \dots$$

Half-Integral Spins

(Fermions)

$I(2I+1)$ Antisymmetric nuclear spin f's \times J even

$(I+1)(2I+1)$ Symmetric " \times J odd

$$q_{\text{rot,nuc}}(T) = (I+1)(2I+1) \sum_{J \text{ odd}} \dots + I(2I+1) \sum_{J \text{ even}} \dots$$

(Antisymmetric)

Example H_2 molecule

$$q_{\text{rot,nuc}} \neq q_{\text{rot}} q_{\text{nuc}}$$

At ordinary temperatures $\theta_r \ll T$

$$\sum_{J \text{ even}} \approx \sum_{J \text{ odd}} \approx \frac{1}{2} \int_0^{\infty} (2J+1) e^{-\theta_r J(J+1)/T} = \frac{1}{2} \frac{T}{\theta_r}$$

For Integral Spins

$$q_{\text{rot, nucl}}(T) = (I+1)(2I+1) \cdot \frac{1}{2} \frac{T}{\theta_r} + I(2I+1) \frac{1}{2} \frac{T}{\theta_r}$$

$$= \frac{1}{2} \frac{T}{\theta_r} (2I+1) (I+1+I)$$

"High temp"

$$q_{\text{rot, nucl}}(T) = \underbrace{\frac{1}{2} \frac{T}{\theta_r}}_{q_{\text{rot}}} \underbrace{(2I+1)^2}_{q_{\text{nucl}}}$$

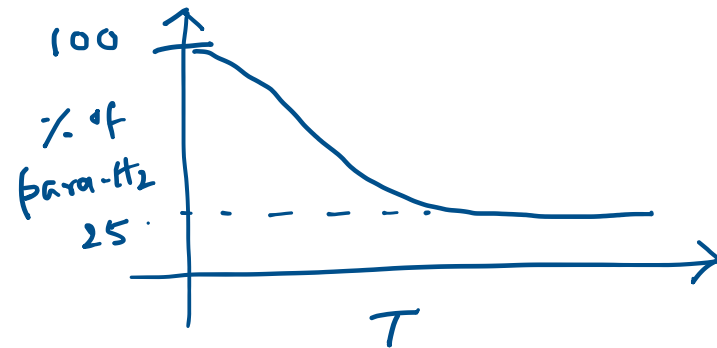
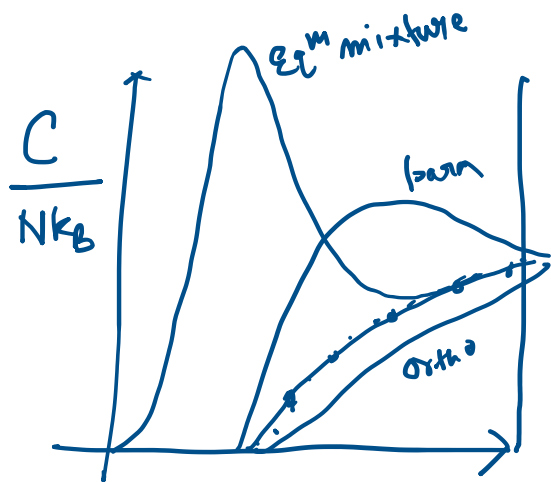
$$q_{\text{rot, nucl}}(T) = I(2I+1) \frac{1}{2} \frac{T}{\theta_r} + (I+1)(2I+1) \frac{1}{2} \frac{T}{\theta_r}$$

$$q_{\text{rot}}(\text{Homonuclear diatomic}) = \frac{1}{2} \frac{T}{\theta_r} = \frac{1}{2} \frac{T}{\theta_r} (2I+1)^2$$

 Symmetry no.

$$\underline{H_2} \quad q_{\text{rot, nucl}} = \underbrace{\sum_{J \text{ even}} (2J+1) e^{-J(J+1)\theta_r/T}}_{\text{para}} + 3 \underbrace{\sum_{J \text{ odd}} (2J+1) e^{-J(J+1)\theta_r/T}}_{\text{ortho}}$$

$$\frac{N_{\text{ortho}}}{N_{\text{para}}} = \frac{3 \sum_{J \text{ odd}} \dots}{1 \sum_{J \text{ even}} \dots} \rightarrow 3 \text{ as } T \text{ increases}$$



"SM
Mc.Quarrie"

Quantum Statistics

N indistinguishable particles
(Quantum)

Microcanonical Ensemble
(N, V, E)

"Single particle energy levels"

"Fermions"

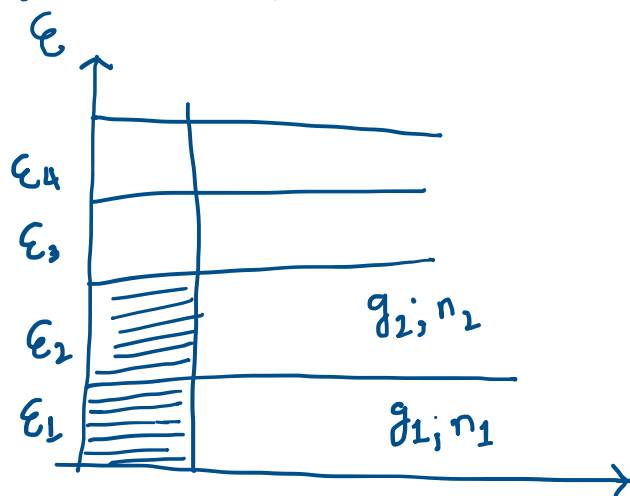
"Bosons"

"grouping into cells"

n_1 particles in first cell

n_2 " " " Second cell

$g_i \rightarrow$ Degeneracy



$$\sum_i n_i = N ; \textcircled{I} \quad \sum_i n_i E_i = E \textcircled{II}$$

$$\Omega(N, V, E) = \sum_{\{n_i\}} W\{n_i\}$$

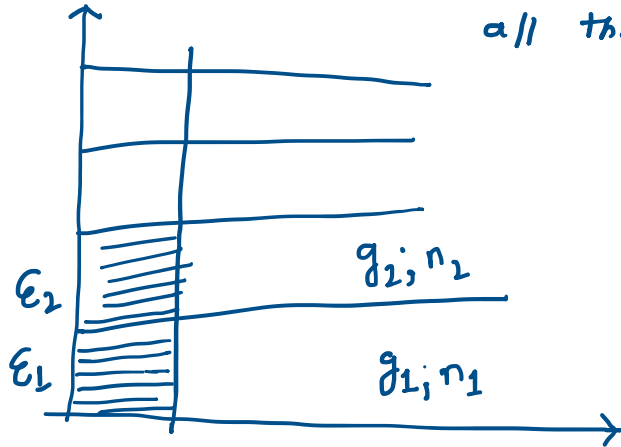
\downarrow No. of distinct microstates accessible to the system under the macrostate N, V, E

$W\{n_i\} \rightarrow$ No. of distinct microstates associated with the distribution set $\{n_i\}$

\downarrow Due to constraints $\textcircled{I} \& \textcircled{II}$

$$W\{n_i\} = \prod_i w(i)$$

→ The product goes over



all the cells
in the spectrum

$w(i)$ is the no. of distinct microstates associated with i th cell of the spectrum (the cell contains n_i particles to be accommodated among g_i levels)

$w(i)$ is the number of distinct ways in which the n_i identical and indistinguishable particles can be distributed among the g_i levels of the cell.

Fermions (follow Fermi-Dirac distribution)

$$n_i = 0 \text{ or } 1$$

$g_i \gg n_i$ (Fermions)

$$w_{F.D}(i) = \frac{g_i!}{n_i! (g_i - n_i)!}$$

$$W_{F.D} = \prod_i \frac{g_i!}{n_i! (g_i - n_i)!}$$

Bosons (follow Bose-Einstein Statistics):

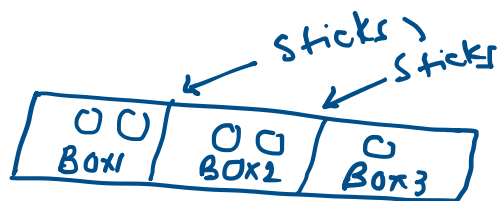
$$n_i = 0, 1, 2, 3, \dots$$

$$n_i^{\max} = \infty \quad (N \rightarrow \infty)$$

$$n_i = 5$$

$$g_i = 3$$

$(g_i - 1)$ no.
of sticks



ONE possibility

(These balls are
indistinguishable)

→ Other possibilities → can be generated
by moving sticks

$(n_i + g_i - 1)$ no. of things to be arranged
to generate configurations
 \downarrow \downarrow
 No. of particles No. of sticks

$$\omega_{B.E.}(i) = \frac{(n_i + g_i - 1)!}{n_i! (g_i - 1)!}$$

$$n_i = 2, g_i = 3$$

$$n_i = 5, g_i = 3$$

$$\omega_{B.E.} = \frac{(5+3-1)!}{5! 2!} = 21$$

↓
Possibilities

$$\omega_{B.E.} = \frac{4!}{2! 2!} = 6; \quad \omega_{F.D.} = \frac{3!}{2! 1!} = 3$$

Classical particles:
(Distinguishable)

n_i particles may be put into any of the g_i levels, independently of one another and the resulting states may be counted as distincts

The no. of these states is clearly $(g_i)^{n_i}$

$$\frac{N!}{n_1! n_2! \dots} \rightsquigarrow \text{weight factor} \quad \frac{1}{n_1! n_2! \dots} = \prod_i \frac{1}{n_i!}$$

$$W_{MB}\{n_i\} = \prod_i \frac{(g_i)^{n_i}}{n_i!}$$

↓
Maxwell-Boltzmann

Entropy: $S = k_B \ln \Omega(N, V, E)$

$$= k_B \ln \left[\sum_{\{n_i\}} W\{n_i\} \right] \simeq k_B \ln W(\{n_i^*\})$$

(Approximately equal to the
logarithm of the
longest term in
the sum)
"Most probable".

$\{n_i^*\}$ is the distribution set

that maximizes the no. $W\{n_i^*\}$

$n_i^* \rightarrow$ is the the Most probable n_i

(Do not forget the constraints)

Employing Lagrange's method of undetermined multipliers

$$\delta \ln W\{n_i\} - \alpha \sum_i \delta n_i - \beta \sum_i \epsilon_i \delta n_i = 0$$

$$\delta \ln W\{n_i\} = \sum_i \delta \ln W(i)(n_i)$$

F. D Statistics

$$\omega_{F.D.}^{(i)} = \frac{g_i!}{n_i! (g_i - n_i)!}$$

$$\begin{aligned} \ln \omega_{F.D.}^{(i)} &= \ln g_i! - \ln n_i! - \ln (g_i - n_i)! \\ &= g_i \ln g_i - \cancel{g_i} - n_i \ln n_i + \cancel{n_i} - (g_i - n_i) \ln (g_i - n_i) \\ &\quad + \cancel{g_i - n_i} \end{aligned}$$

$$= -g_i \ln \left(\frac{g_i - n_i}{g_i} \right) + n_i \ln \left(\frac{g_i - n_i}{n_i} \right)$$

$$\ln \omega_{F.D.}^{(i)} = -g_i \ln \left(1 - \frac{n_i}{g_i} \right) + n_i \ln \left(\frac{g_i}{n_i} - 1 \right)$$

$$\omega_{B.E.}^{(i)} = \frac{(n_i + g_i - 1)!}{n_i! (g_i - 1)!}$$

$$\begin{aligned} \ln \omega_{B.E.}^{(i)} &= \ln (n_i + g_i - 1)! - \ln n_i! - \ln (g_i - 1)! \\ &= (n_i + g_i - 1) \ln (n_i + g_i - 1) - n_i \ln n_i + \cancel{n_i} - (g_i - 1) \ln (g_i - 1) + \cancel{g_i - 1} \\ &\quad - (n_i + g_i - 1) \end{aligned}$$

$$g_i - 1 \approx g_i$$

$$\begin{aligned} \ln \omega_{B.E.}^{(i)} &\approx \\ &\approx n_i \ln \left(\frac{n_i + g_i - 1}{n_i} \right) + (g_i - 1) \ln \left(\frac{n_i + g_i - 1}{g_i - 1} \right) \end{aligned}$$

$$\ln \omega_{B.E.}^{(i)} \approx n_i \ln \left(1 + \frac{g_i}{n_i} \right) + g_i \ln \left(\frac{n_i}{g_i} + 1 \right)$$

In general

$$\ln \omega(i) = n_i \ln \left(\frac{g_i}{n_i} - a \right) - \frac{g_i}{a} \ln \left(1 - a \cdot \frac{n_i}{g_i} \right)$$

$$\ln \omega_{F.D}(i) = -g_i \ln \left(1 - \frac{n_i}{g_i} \right) + n_i \ln \left(\frac{g_i}{n_i} - 1 \right)$$

$$\ln \omega_{BE}(i) \simeq n_i \ln \left(1 + g_i/n_i \right) + g_i \ln \left(\frac{n_i}{g_i} + 1 \right)$$

$a = -1$ for B.E.
 $a = 1$ for F.D.
 $a = 0$ for MB.

$$\omega_{MB}(i) = \frac{(g_i)^{n_i}}{n_i!}$$

$$\begin{aligned}
 \ln \omega_{MB}(i) &= n_i \ln g_i - n_i \ln n_i + n_i \\
 &= n_i \left(\ln \left(\frac{g_i}{n_i} \right) + 1 \right)
 \end{aligned}$$

$$\lim_{x \rightarrow 0} \ln(1-x) \simeq -x$$

$$\lim_{a \rightarrow 0} \ln \left(1 - a \frac{n_i}{g_i} \right) = -a \frac{n_i}{g_i}$$

$$\left(-\frac{g_i}{n_i} \right) \left(-a \frac{n_i}{g_i} \right) = n_i$$

$$\ln w(i) = n_i \ln \left(\frac{g_i}{n_i} - a \right) - \frac{g_i}{a} \ln \left(1 - a \cdot \frac{n_i}{g_i} \right)$$

$$\delta \ln w(i)$$

$$\hookrightarrow \left[1 \cdot \ln \left(\frac{g_i}{n_i} - a \right) + n_i \frac{1}{\left(\frac{g_i}{n_i} - a \right)} \times \left(-\frac{g_i}{n_i^2} \right) - \frac{g_i}{a} \frac{1}{\left(1 - a \frac{n_i}{g_i} \right)^2} \left(-a \frac{1}{g_i} \right) \right] \delta n_i$$

$$= \left[\ln \left(\frac{g_i}{n_i} - a \right) - \frac{g_i}{n_i} \left(\frac{1}{\frac{g_i}{n_i} - a} \right) + \frac{1}{1 - a \frac{n_i}{g_i}} \right] \delta n_i$$

$$= \ln \left(\frac{g_i}{n_i} - a \right) \delta n_i$$

Once we put the constraints

$$\ln W(n_i) = \sum_i \delta \ln w(i) (n_i)$$

$$\sum_i \left\{ \ln \left(\frac{g_i}{n_i} - a \right) - \alpha - \beta \varepsilon_i \right\} \delta n_i = 0 \rightarrow \text{Should be true for any arbitrary } \delta n_i$$

Each term in the sum = 0

$$\ln\left(\frac{g_i}{n_i^*} - a\right) - \alpha - \beta \epsilon_i = 0$$

$$\text{or, } \ln\left(\frac{g_i}{n_i^*} - a\right) = \alpha + \beta \epsilon_i$$

$$\text{or, } \frac{g_i}{n_i^*} - a = e^{\alpha + \beta \epsilon_i}$$

$$\text{or, } \frac{g_i}{n_i^*} = a + e^{\alpha + \beta \epsilon_i}$$

$$\text{or, } \frac{n_i^*}{g_i} = \frac{1}{a + e^{\alpha + \beta \epsilon_i}} \Rightarrow n_i^* = \frac{g_i}{(a + e^{\alpha + \beta \epsilon_i})}$$

$a = 1$ for FD

$$n_{i\text{FD}}^* = \frac{g_i}{1 + e^{\alpha + \beta \epsilon_i}}$$

$a = -1$ for BE

$$n_{i\text{BE}}^* = \frac{g_i}{e^{\alpha + \beta \epsilon_i} - 1}$$

$a = 0$ for M.B

$$n_{i\text{MB}}^* = g_i e^{-\alpha - \beta \epsilon_i}$$

$$n_i^* = \frac{g_i}{(\alpha + e^{\alpha + \beta \epsilon_i})}$$

F.D. $\alpha = 1$
 B.E $\alpha = -1$
 M.B $\alpha = 0$

$$\ln w(i) = n_i \ln \left(\frac{g_i}{n_i} - \alpha \right) - \frac{g_i}{\alpha} \ln \left(1 - \alpha \cdot \frac{n_i}{g_i} \right)$$

$$\begin{aligned} S/k_B &= \ln \Omega \\ \downarrow \\ \text{Entropy} &= \ln \sum_{\{n_i\}} W(\{n_i\}) \simeq \ln W(n_i^*) = \sum_i \ln w(i)(n_i^*) \end{aligned}$$

$$[\ln(x/y) = -\ln(y/x)]$$

$$\begin{aligned} \frac{S}{k_B} &= \sum_i n_i^* \ln \left(\frac{g_i}{n_i^*} - \alpha \right) - \frac{g_i}{\alpha} \ln \left(1 - \alpha \frac{n_i^*}{g_i} \right) \\ &= \sum_i \left\{ n_i^* \ln \left(\frac{g_i}{\alpha + e^{\alpha + \beta \epsilon_i}} - \alpha \right) - \frac{g_i}{\alpha} \ln \left(1 - \frac{\alpha}{\alpha + e^{\alpha + \beta \epsilon_i}} \right) \right\} \\ &= \sum_i \left\{ n_i^* (\alpha + \beta \epsilon_i) - \frac{g_i}{\alpha} \ln \left(\frac{\alpha + e^{\alpha + \beta \epsilon_i} - \alpha}{\alpha + e^{\alpha + \beta \epsilon_i}} \right) \right\} = \sum_i \left\{ n_i^* (\alpha + \beta \epsilon_i) + \frac{g_i}{\alpha} \ln \left(\frac{\alpha + e^{\alpha + \beta \epsilon_i}}{e^{\alpha + \beta \epsilon_i}} \right) \right\} \end{aligned}$$

$$\frac{S}{k_B} = \sum_i \{ n_i^* (\alpha + \beta \epsilon_i) + \frac{g_i}{a} \ln(1 + a e^{-\alpha - \beta \epsilon_i}) \}$$

$$\frac{S}{k_B} = \alpha N + \beta E + \sum_i \frac{g_i}{a} \ln(1 + a e^{-\alpha - \beta \epsilon_i})$$

$$\sum_i \frac{g_i}{a} \ln(1 + a e^{-\alpha - \beta \epsilon_i}) = \frac{S}{k_B} - \alpha N - \beta E$$

$$\alpha = -\mu/k_B T ; \beta = \frac{1}{k_B T} \quad (\text{Done earlier})$$

$$\frac{S}{k_B} + \frac{\mu N}{k_B T} - \frac{E}{k_B T} = \sum_i \frac{g_i}{a} \ln(1 + a e^{-\alpha - \beta \epsilon_i})$$

$$\underline{TS + \mu N - E} = \frac{k_B T}{a} \sum_i g_i \ln(1 + a e^{-\alpha - \beta \epsilon_i})$$

$$\Rightarrow \quad \boxed{PV = \frac{k_B T}{a} \sum_i g_i \ln(1 + a e^{-\alpha - \beta \epsilon_i})}$$

$$\sum_i n_i^* = N$$

$$\alpha \sum_i n_i^* = \sum_i \alpha n_i^* = \alpha N$$

$$\beta \sum_i n_i^* \epsilon_i = \beta E$$

$$\begin{aligned} \mu N &= G \quad (\text{Gibbs Free energy}) \\ &= E + PV - TS \end{aligned}$$

Equation of state for a quantum ideal gas

$$PV = \frac{k_B T}{\alpha} \sum_i g_i \ln(1 + \underbrace{a e^{-\alpha - \beta \epsilon_i}}_{\text{"Small"}})$$

In the Maxwell-Boltzmann case
 $\alpha \rightarrow 0$

$$\lim_{x \rightarrow 0} \ln(1+x) = x$$

$$\simeq \frac{k_B T}{\alpha} \sum_i g_i a e^{-\alpha - \beta \epsilon_i}$$

$$= k_B T \sum_i g_i \underbrace{e^{-\alpha - \beta \epsilon_i}}_{n_i^*(MB)}$$

$$= k_B T \sum_i n_i^* = N k_B T$$

$$PV = N k_B T$$

"CLASSICAL"
 ideal gas law
 "MB"

Ideal Bose Gas (non-interacting collection of Bosons)

$$\alpha = -1$$

$$\frac{PV}{k_B T} = - \sum_i g_i \ln(1 - e^{-\alpha - \beta \epsilon_i})$$

$$\frac{PV}{k_B T} = - \sum_i g_i \ln(1 - z e^{-\beta \epsilon_i})$$

\rightsquigarrow Sum is over energy levels

$$e^{-\alpha} \equiv e^{\mu/k_B T} \equiv z \downarrow \text{fugacity}$$

Replacing the sum over energy levels with the sum over states and replacing the sum with an integral giving a statistical weight to each state

[Either high temp or large volume]

$$\frac{PV}{k_B T} = - \sum_i g_i \ln(1 - e^{-\alpha - \beta \epsilon_i})$$

$$= - \sum_i g_i \ln(1 - z e^{-\beta \epsilon_i})$$

$g(\epsilon) \cdot d\epsilon \rightarrow$ No. of States betⁿ ϵ and $\epsilon + d\epsilon$

$$\frac{PV}{k_B T} = - \int d\epsilon \underbrace{g(\epsilon)}_{\text{Density of states}} \ln(1 - z e^{-\beta \epsilon})$$

$$N = \sum_i \langle n_i \rangle = \sum_i \frac{g_i}{(z^{-1} e^{\beta \epsilon_i} - 1)} = \sum_i \frac{g_i}{(z^{-1} e^{\beta \epsilon_i} - 1)}$$

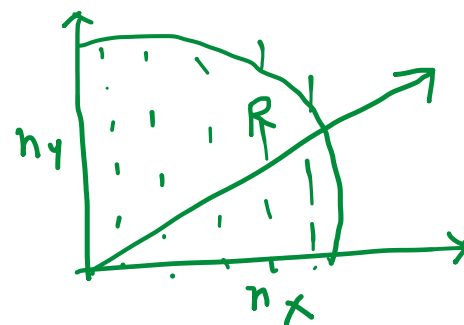
$$N = \int \frac{d\epsilon g(\epsilon)}{(z^{-1} e^{\beta \epsilon} - 1)}$$

How to calculate the density of states?

3D Box
Particle
in

$$E = \epsilon = \frac{(n_x^2 + n_y^2 + n_z^2) \hbar^2}{8mL^2}$$

$$R^2 = \frac{8mL^2 E}{\hbar^2} = n_x^2 + n_y^2 + n_z^2$$



Calculate the
no. of lattice
points

No. of States
with energy

$$\phi(E) = \frac{1}{8} \left(\frac{4}{3} \pi R^3 \right) = \frac{1}{8} \frac{4}{3} \pi \left(\frac{8mL^2 E}{\hbar^2} \right)^{3/2} = \frac{\pi}{6} \left(\frac{8mL^2 E}{\hbar^2} \right)^{3/2}$$

↓ only one octant

$n_x, n_y, n_z \rightarrow$ cannot be negative

$$\phi(E + \Delta E) =$$

$$\Delta E/E \ll 1$$

$$V = L^3$$

$$\phi'(E) = \frac{d\phi}{dE}$$

$$W(E, \Delta E) = \phi(E + \Delta E) - \phi(E)$$

$$\simeq \phi(E) + \phi'(E) \Delta E - \phi(E) = \phi'(E) \Delta E$$

$$= \frac{\pi}{6} \left(\frac{8mL^2}{\hbar^2} \right)^{3/2} \frac{3}{2} E^{3/2-1} \Delta E = \frac{\pi}{4} \left(\frac{8m}{\hbar^2} \right)^{3/2} L^3 E^{1/2} \Delta E$$

↑ = V

$$\frac{PV}{k_B T} = - \int \uparrow \ln(1 - z e^{-\beta E})$$

$$g(E) dE = \left(\frac{2\pi V}{h^3} \right) (2m)^{3/2} E^{1/2} dE \quad (\text{Please check})$$

$$\langle n_k \rangle_{BE} = \frac{z e^{-\beta E_k}}{(1 - z e^{-\beta E_k})} \quad \text{Bosons}$$

$$\langle n_k \rangle_{FD} = \frac{z e^{-\beta E_k}}{1 + z e^{-\beta E_k}} \quad \text{Fermions}$$

What happens if T is high or density is low ($\rho = N/V$). \Rightarrow No. of available energy states \gg No. of particles \rightarrow "CLASSICAL"

$$\langle n_k \rangle_{FD/BE} \simeq 0 \quad \sim \quad z \rightarrow 0$$

(Weakly degenerate case)

$$\langle n_k \rangle = z e^{-\beta E_k}$$

$$\sum_k \langle n_k \rangle = z \sum_k e^{-\beta E_k} \simeq z q = N \quad \hookrightarrow \quad z = N/q$$

$$\langle n_k \rangle = z e^{-\beta E_k}$$

$$\langle n_k \rangle = \frac{N}{q} e^{-\beta E_k}$$

$$z = e^{\mu/k_B T}$$

$z \rightarrow 0$ (Classical)

$z \gg 0$ (Quantum)

$$\boxed{\frac{\langle n_k \rangle}{N} = \frac{e^{-\beta E_k}}{q}} \rightarrow \text{M.B}$$

Equation of State for the Bose gas (ideal Bose)

$$\frac{PV}{k_B T} = - \left(\frac{2\pi V}{h^3} \right) (2m)^{3/2} \int_0^\infty d\epsilon \ln(1 - ze^{-\beta\epsilon}) \epsilon^{1/2}$$

ONE Problem

The state $\epsilon=0$ has been given zero weight

We should take this state out and then replace the sum with the integral

$$\frac{P}{k_B T} = - \left(\frac{2\pi}{h^3} \right) (2m)^{3/2} \int_0^\infty d\epsilon \ln(1 - ze^{-\beta\epsilon}) \epsilon^{1/2} - \frac{\ln(1-z)}{V}$$

Similarly

$$\frac{N}{V} = \left(\frac{2\pi}{h^3} \right) (2m)^{3/2} \int_0^\infty \frac{\epsilon^{1/2}}{(z^{-1}e^{\beta\epsilon} - 1)} + \frac{1}{V} \frac{1}{(z^{-1} - 1)}$$

$$N = N_e + N_0$$

No. of particles in the ground state
 $N_0 \leq \frac{z}{1-z}$

$$N_e = \left(\frac{2\pi V}{h^3} \right) (2m)^{3/2} \int \dots \dots T \text{ is small } \xrightarrow{T \rightarrow 0}$$

BEC

→ Bose-Einstein Condensation

$z \ll 1$ (classical)

$$\frac{1}{V} \left(\frac{z}{1-z} \right) \sim \frac{1}{N} z$$

~ vanishing (small)

$z \gg 0 \rightarrow$ approaching 1
 $\frac{1}{V} \left(\frac{z}{1-z} \right)$ will blow up!!

New state of matter

As $T \downarrow$, the $\epsilon=0$ level gets densely populated.

$$\frac{P}{k_B T} = - \left(\frac{2\pi}{h^3} \right) \frac{(2m)^{3/2}}{\beta^{3/2}} \int_0^\infty dx x^{1/2} \ln(1 - ze^{-x})$$

$$\boxed{\frac{P}{k_B T} = \frac{1}{\lambda^3} g_{5/2}(z)} + \dots \dots$$

$$\frac{N_e}{V} = \frac{N - N_0}{V} = \left(\frac{2\pi}{h^3} \right) (2m)^{3/2} \frac{1}{\beta^{3/2}} \int_0^\infty \frac{x^{1/2} dx}{(z^{-1}e^x - 1)}$$

$$\boxed{\frac{N_e}{V} = \frac{1}{\lambda^3} g_{3/2}(z)}$$

$$\lambda = \frac{h}{(2\pi m k_B T)^{1/2}}$$

$g_\nu(z) \rightarrow$ Bose-Einstein functions

$$g_\nu(z) = \frac{1}{\Gamma(\nu)} \int_0^\infty \frac{x^{\nu-1} dx}{(z^{-1}e^x - 1)} = z + \frac{z^2}{2^\nu} + \frac{z^3}{3^\nu} + \dots$$

$$\begin{aligned} \frac{z}{1-z} &= N_0 \Rightarrow \ln(1-z) \\ z &= \frac{N_0}{N_0+1} \Rightarrow -\ln(1-z) = \ln \frac{1}{N_0+1} \\ &= -\ln \frac{1}{N} \\ &\sim \frac{1}{N} \ln N \end{aligned}$$

~ 0 (thermodynamic limit)

$$\beta \epsilon = \frac{p^2}{2m k_B T} = x$$

$$\beta d\epsilon = dx$$

$$d\epsilon = \frac{1}{\beta} dx$$

$$\epsilon'^{1/2} = \frac{x^{1/2}}{\beta^{1/2}}$$

Upper bound of $(\nu = 3/2)$
 $g_{3/2}(z)$

$$g_{3/2}(1) = \zeta(3/2)$$

$$= 1 + \frac{1}{2^{3/2}} + \frac{1}{3^{3/2}} + \dots$$

$$\zeta(3/2) \simeq 2.62$$

$$N_e = \frac{V}{\lambda^3} g_{3/2}(z)$$

if
fugacity is 1 (quantum)

$$N_e = \frac{V}{\lambda^3} \zeta(3/2)$$

$$N_e < \frac{V}{\lambda^3} \zeta(3/2)$$

\swarrow
 $T > 0$

$\underbrace{\hspace{1.5cm}}$
 No. of particles
 in the excited states
 at $z=1$ ($T=0$)

if $N_e = \frac{V}{\lambda^3} g_{3/2}(z)$
 fugacity is 1 (quantum)

$$N_e = \frac{V}{\lambda^3} \zeta(3/2)$$

$$N_e \leq \frac{V}{\lambda^3} \zeta(3/2)$$

\downarrow $T > 0$
 $\underbrace{\hspace{10em}}$ No. of particles in the excited states at $z=1$ ($T=0$)

$$g_{3/2}(z) = 1 + \frac{z^2}{2^{3/2}} + \frac{z^3}{3^{3/2}} + \dots$$

$z = 1$ (quantum)
 $z = 0$ (classical)

"upper bound"

$$g_{3/2}(z) = 1 + \frac{z^2}{2^{3/2}} + \frac{z^3}{3^{3/2}} + \dots$$

$$z=1 \quad g_{3/2}(1) \equiv \zeta(3/2) = 1 + \frac{1}{2^{3/2}} + \frac{1}{3^{3/2}} + \dots > g_{3/2}(z)$$

$$g_{3/2}(z) \leq \zeta(3/2)$$

$$N_e = \frac{V}{\lambda^3} g_{3/2}(z) = V \left(\frac{2\pi m k_B T}{h^2} \right)^{3/2} g_{3/2}(z) \leq \frac{V}{\lambda^3} \zeta(3/2)$$

\swarrow No. of particles in the excited state ($\epsilon \neq 0, > 0$)
 \downarrow Any z
 \downarrow $z=1$

No. of particles occupying the ground state ($\epsilon=0$)

$$N_0 = N - V \frac{(2\pi m k_B T)^{3/2}}{h^3} \zeta(3/2)$$

Condition for the onset of BEC

$$(*) \quad N > V \frac{(2\pi m k_B T)^{3/2}}{h^3} \zeta(3/2)$$

↓
Total no. of particles exceeds the "upper bound" of the excited state

In terms of temperature??

"CRITICAL TEMPERATURE"
for a given N and V

Below this C.T (T_c) → Two phases exist

① No particles in the excited state ② $N - N_e$ in the ground state

$$\epsilon \rightarrow p$$

$$p_{\text{rms}}^2 \sim \epsilon$$

What is special about BEC?

"Phase Transition"

* No intermolecular interaction!!
(Purely quantum mechanical in origin)

* This is happening in momentum space
(Not in co-ordinate space)

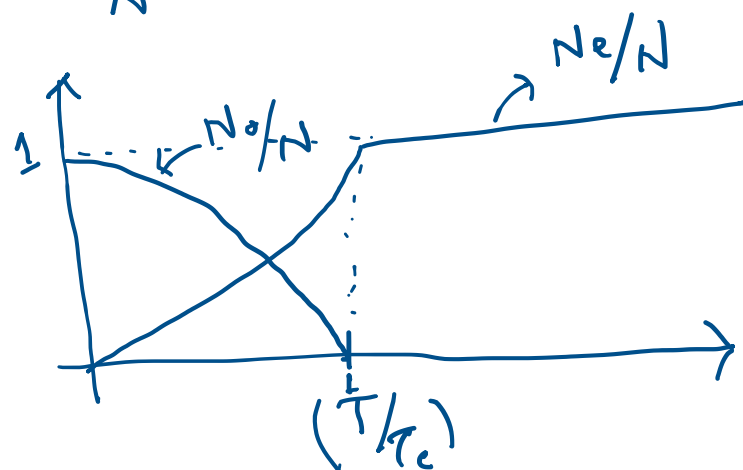
$$N = V \frac{(2\pi m k_B)^{3/2}}{h^3} \zeta(3/2) T_c^{3/2}$$

$$\Rightarrow T_c = \left(\frac{h^2}{2\pi m k_B} \right) \left(\frac{N}{V} \frac{1}{\zeta(3/2)} \right)^{2/3}$$

Complementary fractions

$$\frac{N_0}{N} = 1 - \left(\frac{V}{N} \right) \frac{(2\pi m k_B T)^{3/2}}{h^3} \zeta(3/2)$$

$$\frac{N_0}{N} = 1 - \left(T/T_c \right)^{3/2}$$



$$T_c = \frac{h^2}{2\pi m k_B} \left(\frac{N}{V} \frac{1}{\zeta(3/2)} \right)^{2/3}$$

$$T_c^{3/2} = \left(\frac{h^2}{2\pi m k_B} \right)^{3/2} \left(\frac{N}{V} \right) \frac{1}{\zeta(3/2)}$$

[R.K. Pathria
Statistical Mechanics]

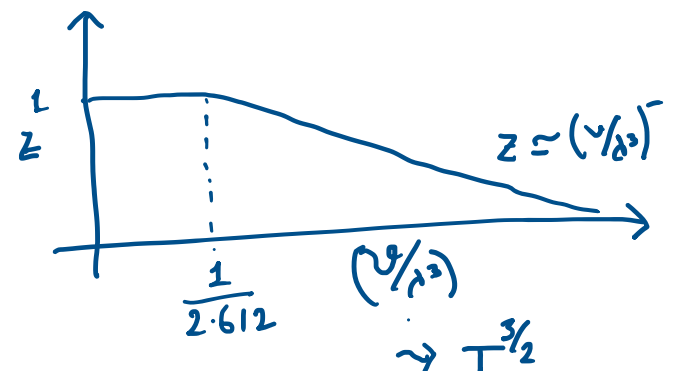
Biman Baychi
SM for Chemists
.....

$$\frac{PV}{Nk_B T} = \sum_{l=1}^{\infty} \left(a_l \left(\lambda^3 / v \right)^{l-1} \right)$$

"Virial Coefficients" $v \equiv \frac{1}{n}$
↓
"vol^m/particle"

$z \rightarrow$ can be expanded in
powers of $n\lambda^3$

$$\zeta(3/2) = 2.612$$



$$\frac{P}{k_B T} = \frac{1}{\lambda^3} g_{5/2}(z)$$

for unit fugacity

at the critical temperature? T_c

$$P(T) = \frac{k_B T}{\lambda^3} \zeta(5/2)$$

$$P(T_c) = k_B T_c \zeta(5/2) \frac{(2\pi m k_B T_c)^{3/2}}{h^3}$$

$$= k_B T_c \zeta(5/2) \frac{(2\pi m k_B)^{3/2}}{h^3} \cdot \left(\frac{h^2}{2\pi m k_B} \right)^{3/2} \left(\frac{N}{V} \frac{1}{\zeta(3/2)} \right)^2$$

$$P(T_c) = k_B T_c \left(\frac{N}{V} \right) \frac{\zeta(5/2)}{\zeta(3/2)}$$

$$P_{(T_c)} = k_B T_c \left(\frac{N}{V} \right) 0.5134$$

$$\approx \frac{1}{2} k_B T_c \left(\frac{N}{V} \right)$$

$$P = k_B T \left(\frac{N}{V} \right)$$

ideal gas
classical

Ideal Boson gas at $T=T_c$
exerts pressure $\frac{1}{2}$ of

that of a classical (Boltzmannian) ideal gas

Internal Energy:

$$\frac{PV}{k_B T} = \ln \Xi \quad \leftarrow \text{grand canonical partition function}$$

$$E = k_B T^2 \frac{\partial}{\partial T} \left(\frac{PV}{k_B T} \right)_{z, V}$$

$$= k_B T^2 V g_{5/2}(z) \left(\frac{\partial}{\partial T} \left(\frac{1}{\lambda^3} \right) \right)$$

$$= k_B T^2 V g_{5/2}(z) \frac{3}{2} \frac{1}{\lambda^3}$$

$$E = \frac{3}{2} k_B T V g_{5/2}(z) \frac{1}{\lambda^3}$$

$$P = \frac{k_B T}{\lambda^3} g_{5/2}(z)$$

$$\boxed{E = \frac{3}{2} P V}$$

→ familiar
→ It satisfies the relationship for the ideal gas!!

$$\frac{PV}{k_B T} = \frac{V}{\lambda^3} g_{5/2}(z)$$

$$\frac{\partial}{\partial T} \left[\frac{(2\pi m k_B T)^{3/2}}{h^3} \right]$$

$$= \frac{(2\pi m k_B)^{3/2}}{h^3} \frac{3}{2} T^{1/2}$$

$$= \left(\frac{(2\pi m k_B T)^{3/2}}{h^3} \right)^{3/2} \frac{3}{2} \frac{T^{1/2}}{T^{3/2}}$$

$$= \frac{1}{\lambda^3} \frac{3}{2} \frac{1}{T}$$

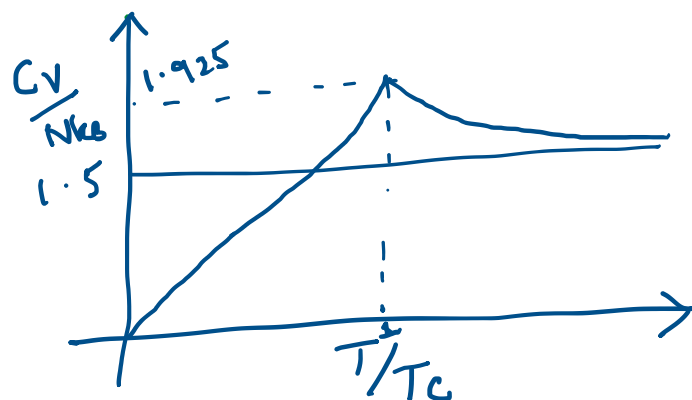
$$E \approx \frac{3}{2} P V \quad (\text{ideal gas})$$

$$PV = N k_B T$$

$$E = \frac{3}{2} N k_B T$$

Specific Heat of an ideal Bose Gas °

For unit fugacity ($T \sim T_c$)



$$\frac{T}{\lambda^3} = \left(\frac{2\pi m k_B}{h^3} \right)^{3/2} T^{5/2}$$

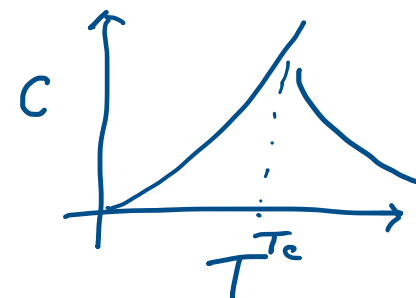
$$\frac{C_V}{N k_B} = \frac{1}{N k_B} \left(\frac{\partial E}{\partial T} \right)_{V,z}$$

$$= \frac{1}{N k_B} \frac{3}{2} k_B V \zeta(5/2) \left\{ \frac{d}{dT} \left(\frac{T}{\lambda^3} \right) \right\}$$

$$= \frac{3V}{2N} \zeta(5/2) \frac{(2\pi m k_B)^{3/2}}{h^3} \frac{5}{2} T^{5/2-1}$$

$$= \frac{15}{4} \frac{V}{N} \zeta(5/2) \frac{(2\pi m k_B)^{3/2}}{h^3} T^{3/2}$$

He I and He II
T
He-4
(Boson).



$$\frac{C_V(T=T_c)}{N k_B} = \frac{15}{4} \frac{V}{N} \zeta(5/2) \frac{(2\pi m k_B)^{3/2}}{h^3} \frac{h^3}{(2\pi m k_B)^{3/2}} \left(\frac{N}{V} \frac{1}{\zeta(3/2)} \right)^1 = \frac{15}{4} \frac{\zeta(5/2)}{\zeta(3/2)} \simeq 1.925$$

$$T \gg T_c \text{ (High temperature limit)} \Rightarrow \frac{C_V}{N k_B} = \frac{3}{2}$$

Fermi (ideal) gas

$$\langle n_k \rangle = \frac{1}{1 + e^{\alpha + \beta \epsilon_k}} = \frac{1}{1 + \underbrace{e^{\beta(\epsilon_k - \mu)}}_{\langle n_k \rangle}} \rightarrow \begin{array}{l} \beta \text{ is positive} \\ (\epsilon_k - \mu) \text{ is real} \\ e^{\beta(\epsilon_k - \mu)} \text{ is always } > 0 \end{array}$$

$\langle n_k \rangle$ will vary betⁿ 0 and 1
(Pauli-Exclusion principle)

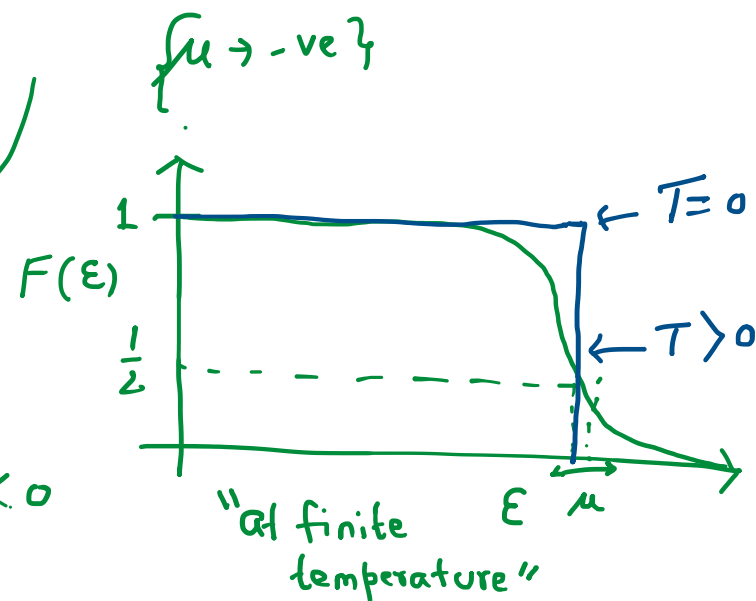
Let us define a fⁿ (Fermi function)

$$F(\epsilon) = \frac{1}{1 + e^{\beta(\epsilon - \mu)}} = \frac{1}{1 + e^{\beta\epsilon - \beta\mu}} \quad \left\{ \begin{array}{l} \beta\mu < 0 \text{ (condition)} \\ e^{\beta(\epsilon - \mu)} \gg 1 \end{array} \right.$$

$\sim e^{-\beta(\epsilon - \mu)}$ (Boltzmann).

$\epsilon = \mu \quad F(\epsilon) = \frac{1}{1+1} = \frac{1}{2}$
 ↳ "Fermi Energy" (chemical pot).

- Let us consider
- i) $\beta\mu \gg 1$
 - ii) $\epsilon \ll \mu; \beta(\epsilon - \mu) < 0$
 $F(\epsilon) \sim 1$





1. The ensemble average of a function $B(p, q)$ is given by

$$\langle B \rangle = \int d\Gamma \rho(p, q, t) B(p, q)$$

Show

$$\frac{d}{dt} \langle B \rangle = \langle \{B, H\} \rangle,$$

where p and q represent momenta and position in $6N$ dimensional phase space, integration is over the phase space co-ordinates (momenta and positions) and H is the Hamiltonian of the classical system.

Answer: $d\Gamma = \prod_{i=1}^N d^3q_i d^3p_i$.

The ensemble average of a function $B(p, q)$ can be expressed as

$$\langle B \rangle = \int d\Gamma \rho(p, q, t) B(p, q).$$

Therefore,

$$\frac{d}{dt} \langle B \rangle = \int d\Gamma B(p, q) \frac{\partial}{\partial t} \rho(p, q, t), \quad (1)$$

as $\frac{\partial}{\partial t} B(p, q) = 0$. Using the Liouville equation,

$$\frac{\partial}{\partial t} \rho(p, q, t) + \{\rho, H\} = 0,$$

where

$$\{, H\} = \sum_{i=1}^{3N} \left(\frac{\partial H}{\partial p_i} \frac{\partial}{\partial q_i} - \frac{\partial H}{\partial q_i} \frac{\partial}{\partial p_i} \right),$$

one can explicitly write Eq. (1) as

$$\frac{d}{dt} \langle B \rangle = - \int d\Gamma B(p, q) \sum_{i=1}^{3N} \left(\frac{\partial H}{\partial p_i} \frac{\partial \rho}{\partial q_i} - \frac{\partial H}{\partial q_i} \frac{\partial \rho}{\partial p_i} \right). \quad (2)$$

Doing the integration by parts in Eq. (2) and then taking boundary terms as zero, one can arrive at

$$\begin{aligned} \frac{d}{dt} \langle B \rangle &= \int d\Gamma \rho(p, q) \sum_{i=1}^{3N} \left(\frac{\partial H}{\partial p_i} \frac{\partial B}{\partial q_i} - \frac{\partial H}{\partial q_i} \frac{\partial B}{\partial p_i} \right) + \int d\Gamma \rho(p, q) \sum_{i=1}^{3N} B \left(\frac{\partial}{\partial p_i} \frac{\partial H}{\partial q_i} - \frac{\partial}{\partial q_i} \frac{\partial H}{\partial p_i} \right) \\ &= \langle \{B, H\} \rangle. \end{aligned} \quad (3)$$

2. For a classical “ N ” particle system, let other than the Hamiltonian the angular momentum (l_n) is also conserved. Then show mathematically that $\frac{dl_n}{dt} = 0$.

Answer: As both the Hamiltonian (H) and the angular momentum (l_n) are conserved, $\{H, l_n\} = 0$. We know, $l_n = l_n(\{p_i\}, \{q_i\})$. Therefore,

$$\frac{dl_n}{dt} = \sum_{i=1}^{3N} \left(\frac{\partial l_n}{\partial q_i} \frac{dq_i}{dt} + \frac{\partial l_n}{\partial p_i} \frac{dp_i}{dt} \right)$$

Using relations, $\frac{dq_i}{dt} = \dot{q}_i = \frac{\partial H}{\partial p_i}$ and $\frac{dp_i}{dt} = \dot{p}_i = -\frac{\partial H}{\partial q_i}$, one can rewrite the above equation as

$$\frac{dl_n}{dt} = \sum_{i=1}^{3N} \left(\frac{\partial l_n}{\partial q_i} \frac{\partial H}{\partial p_i} - \frac{\partial l_n}{\partial p_i} \frac{\partial H}{\partial q_i} \right) = \{l_n, H\} = 0. \quad (4)$$

3. For a single particle subjected to a potential $V(q)$, write down the action, $A[q(t)]$, where $A[q(t)] = \int_{t_a}^{t_b} dt L(q, \dot{q})$, and $L(q, \dot{q})$ is the Lagrangian of the system. Show that the extremization of this action results the Newton’s equation of motion.

Answer: The action is given by

$$A[q(t)] = \int_{t_a}^{t_b} dt L(q, \dot{q}).$$

The extremization of action, *i.e.*, $\delta A = 0$ leads to the Euler–Lagrange equation,

$$\frac{\partial L}{\partial q} - \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}} \right) = 0. \quad (5)$$

For a single particle of mass m subjected to a potential $V(q)$, the action is given by

$$L(q, \dot{q}) = \frac{m}{2} \dot{q}^2 - V(q).$$

So using Eq. (5), one gets

$$m\ddot{q} = -V'(q),$$

which is the Newton’s equation of motion.

4. For a free particle, the Lagrangian, $L = \frac{m}{2} \dot{q}^2$. Show that the action A_q corresponding to the classical motion of a free particle is

$$A_q = \frac{m}{2} \frac{(q_b - q_a)^2}{(t_b - t_a)},$$

where end points are q_a and q_b at time t_a and t_b , respectively.

Answer: For a free particle, the Lagrangian is $L = \frac{m}{2} \dot{q}^2$. So the action is given by $A[q(t)] = \frac{m}{2} \int_{t_a}^{t_b} dt \dot{q}^2$. Using the Euler–Lagrange equation given in Eq. (5), one finds $m\ddot{q}_{cl} = 0$, which is

the equation for classical motion followed by the particle. From this, we have $\dot{q}_{cl} = \text{constant} = (q_b - q_a)/(t_b - t_a)$. Putting the value of \dot{q}_{cl} in the expression of action, one gets

$$A_q = A[q(t)]_{q(t)=q_{cl}} = \frac{m}{2} \frac{(q_b - q_a)^2}{(t_b - t_a)}.$$

5. For a harmonic oscillator, the Lagrangian, $L = \frac{m}{2}(\dot{q}^2 - \omega^2 q^2)$. Show that the classical action is

$$A_q = \frac{m\omega}{2\sin\omega T} [(q_b^2 + q_a^2)\cos\omega T - 2q_a q_b],$$

where $T = t_b - t_a$.

Answer: The Lagrangian for a harmonic oscillator is given by $L = \frac{m}{2}(\dot{q}^2 - \omega^2 q^2)$. So by virtue of Eq. (5), one can find the equation of motion for the classical trajectory, which is given as

$$\ddot{q}_{cl} + \omega^2 q_{cl} = 0.$$

The general solution of the above equation is: $q_{cl}(t) = c_1 \sin\omega t + c_2 \cos\omega t$. The boundary conditions are: $q_{cl}(t)|_{t=t_b} = q_b$ and $q_{cl}(t)|_{t=t_a} = q_a$. Using them, one can get the following matrix equation for c_1, c_2 :

$$\begin{bmatrix} \sin\omega t_b & \cos\omega t_b \\ \sin\omega t_a & \cos\omega t_a \end{bmatrix} \begin{bmatrix} c_1 \\ c_2 \end{bmatrix} = \begin{bmatrix} q_b \\ q_a \end{bmatrix} \implies \begin{bmatrix} c_1 \\ c_2 \end{bmatrix} = \begin{bmatrix} \sin\omega t_b & \cos\omega t_b \\ \sin\omega t_a & \cos\omega t_a \end{bmatrix}^{-1} \begin{bmatrix} q_b \\ q_a \end{bmatrix}. \quad (6)$$

Doing the matrix calculation, one obtains, $c_1 = -q_a \frac{\cos\omega t_b}{\sin\omega T} + q_b \frac{\cos\omega t_a}{\sin\omega T}$ and $c_2 = q_a \frac{\sin\omega t_b}{\sin\omega T} - q_b \frac{\sin\omega t_a}{\sin\omega T}$, where $T = t_b - t_a$. Therefore the classical action can be calculated as

$$A_q = \frac{m}{2} \int_{t_a}^{t_b} dt (\dot{q}_{cl}^2 - \omega^2 q_{cl}^2) = \frac{m\omega}{2\sin\omega T} [(q_b^2 + q_a^2)\cos\omega T - 2q_a q_b]$$

Detailed mathematical Steps:

$$\begin{aligned} \frac{m}{2} \int_{t_a}^{t_b} dt \dot{q}_{cl}^2 &= \frac{m}{2} [q(t)_{cl} \dot{q}_{cl}(t)]_{t=t_a}^{t=t_b} - \frac{m}{2} \int_{t_a}^{t_b} dt q_{cl}(t) \ddot{q}_{cl}(t) \\ &= \frac{m}{2} [q_{cl}(t) \dot{q}_{cl}(t)]_{t=t_a}^{t=t_b} - \frac{m\omega^2}{2} \int_{t_a}^{t_b} dt q_{cl}(t) q_{cl}(t) \end{aligned} \quad (7)$$

Therefore,

$$\begin{aligned} &\frac{m}{2} \int_{t_a}^{t_b} dt (\dot{q}_{cl}^2 - \omega^2 q_{cl}^2) \\ &= \frac{m}{2} [q_{cl}(t) \dot{q}_{cl}(t)]_{t=t_a}^{t=t_b} \\ &= \frac{m}{2} [q_b(c_1\omega \cos\omega t_b - c_2\omega \sin\omega t_b) - q_a(c_1\omega \cos\omega t_a - c_2\omega \sin\omega t_a)] \\ &= \frac{m\omega}{2} \left[q_b^2 \left(\frac{\cos\omega t_b \cos\omega t_a}{\sin\omega T} + \frac{\sin\omega t_b \sin\omega t_a}{\sin\omega T} \right) + q_a^2 \left(\frac{\cos\omega t_b \cos\omega t_a}{\sin\omega T} + \frac{\sin\omega t_b \sin\omega t_a}{\sin\omega T} \right) \right] \\ &\quad - \frac{m\omega}{2} \frac{q_b q_a}{\sin\omega T} [\cos^2\omega t_b + \sin^2\omega t_b + \cos^2\omega t_a + \sin^2\omega t_a] \\ &= \frac{m\omega}{2\sin\omega T} [(q_b^2 + q_a^2)\cos\omega T - 2q_a q_b] \end{aligned} \quad (8)$$

6. The energy levels of a rigid rotor of moment of inertia I are given by

$$E_j = \frac{\hbar^2}{2I} j(j+1),$$

where $j = 0, 1, 2, \dots$;

- (i) Using the Boltzmann statistics (with proper degeneracy factor), find an expression for the thermodynamic internal energy of the system.
- (ii) Under what conditions can the sum in the part (i) be *approximated* by an integral? In this case evaluate the specific heat C_v of the system.

Answer: For a rigid rotor, the energy of j -th state is

$$E_j = \frac{\hbar^2}{2I} j(j+1),$$

with degeneracy $2j+1$, where $j = 0, 1, 2, \dots$. Therefore the partition function is given by

$$q(T) = \sum_{j=0}^{\infty} (2j+1) e^{-\frac{\beta \hbar^2}{2I} j(j+1)}, \quad (9)$$

where $\beta = 1/k_B T$.

The internal energy can be expressed in terms of partition function as $\bar{E} = k_B T^2 \left(\frac{\partial \ln q(T)}{\partial T} \right)$. So using Eq. (9), one has

$$\begin{aligned} \bar{E} &= k_B T^2 \left(\frac{\partial \ln q(T)}{\partial T} \right) = k_B T^2 \sum_{j=0}^{\infty} \frac{(2j+1)}{q(T)} \frac{\hbar^2}{2k_B T^2 I} j(j+1) e^{-\frac{\hbar^2}{2k_B T I} j(j+1)} \\ &= \sum_{j=0}^{\infty} E_j \frac{1}{q(T)} (2j+1) e^{-\frac{\hbar^2}{2k_B T I} j(j+1)}. \end{aligned} \quad (10)$$

For heavy molecules or at high temperature, *i.e.*, if $k_B T \gg B$, where $B = \frac{\hbar^2}{2I}$, then the sum can be replaced by an integral, and the partition function takes the following form:

$$\begin{aligned} \lim_{k_B T/B \gg 1} q(T) &\approx \int_0^{\infty} dj (2j+1) e^{-\beta B j(j+1)} \\ &= -\frac{1}{\beta B} \int_0^{\infty} d \left(e^{-\beta B j(j+1)} \right) = k_B T/B. \end{aligned} \quad (11)$$

So, $\bar{E} = k_B T^2 \left(\frac{\partial \ln q(T)}{\partial T} \right) = k_B T$, and molar average internal energy is $\bar{E}_{molar} = N_A k_B T = RT$. Therefore, $C_v = \frac{\partial \bar{E}}{\partial T} = R$.

7. The partition function of a crystal can be approximated by

$$Q = \left(\frac{e^{-h\nu/2k_B T}}{1 - e^{-h\nu/2k_B T}} \right)^{3N} e^{\frac{U_0}{k_B T}},$$

where $h\nu/k_B = \Theta_E$ is a constant characteristic of the crystal, and U_0 is the sublimation energy of the crystal. Calculate the heat capacity from this simple partition function and show that at high temperatures, one obtains the law of Dulong and Petit, namely $C_v = 3Nk_B$ as $T \rightarrow \infty$.
There is a typo in the question.

Answer: The partition function of a crystal can be approximated by

$$Q = \left(\frac{e^{-h\nu/2k_B T}}{1 - e^{-h\nu/k_B T}} \right)^{3N} e^{\frac{U_0}{k_B T}}.$$

Taking $h\nu/k_B = \Theta_E$, the above can be rewritten as

$$Q = \left(\frac{e^{-\Theta_E/2T}}{1 - e^{-\Theta_E/T}} \right)^{3N} e^{\frac{U_0}{k_B T}}, \quad (12)$$

or, $\ln Q = -3N\Theta_E/2T - 3N\ln(1 - e^{-\Theta_E/T}) + \frac{U_0}{k_B T}$. So the average internal energy is calculated using $\bar{E} = k_B T^2 \left(\frac{\partial \ln Q}{\partial T} \right)$, as

$$\begin{aligned} \bar{E} &= 3Nk_B T^2 \frac{\Theta_E}{2T^2} + 3Nk_B T^2 \frac{\frac{\Theta_E}{T^2} e^{-\Theta_E/T}}{(1 - e^{-\Theta_E/T})} - k_B T^2 \frac{U_0}{k_B T^2} \\ &= \frac{3}{2} N h \nu + \frac{3N h \nu e^{-\Theta_E/T}}{(1 - e^{-\Theta_E/T})} - U_0. \end{aligned} \quad (13)$$

Therefore the heat capacity is given by

$$C_v = \frac{\partial \bar{E}}{\partial T} = 3Nk_B \left(\frac{\Theta_E}{T} \right)^2 \frac{e^{-\Theta_E/T}}{(1 - e^{-\Theta_E/T})^2}. \quad (14)$$

At high temperature, *i.e.*, for $T \rightarrow \infty$, the exponential can be well approximated as $e^{-\Theta_E/T} \approx 1 - \Theta_E/T + \mathcal{O}(1/T^2)$. Applying this, one obtains, $C_v \approx 3Nk_B \left(\frac{\Theta_E}{T} \right)^2 \frac{1}{(\Theta_E/T)^2} = 3Nk_B$, which is the mathematical form of the law of Dulong and Petit.

8. Consider a system of N non-interacting harmonic oscillators in three dimension with total energy E . Show that the microcanonical partition function is given by

$$Q(N, E) = \left(\frac{2\pi E}{3N\hbar} \right)^{3N} e^{3N} \prod_{i=1}^N \frac{1}{\omega_i^3},$$

where ω_i is the angular frequency of the i -th harmonic oscillator. Then show that the above partition function is consistent with the relation: $E = 3Nk_B T$. (Here, you can use: $\Gamma(3N) \approx (3N)! \approx (3N)^{3N} e^{-3N}$.)

Answer: The Hamiltonian for a system of N non-interacting harmonic oscillators in three-dimensional space is given by

$$\mathcal{H} = \sum_{i=1}^N \left(\frac{\mathbf{p}_i^2}{2m_i} + \frac{k_i}{2} \mathbf{r}_i^2 \right), \quad (15)$$

where the mass m_i and the spring constant k_i of the i -th harmonic oscillator are related to the angular frequency ω_i , as $\omega_i = \sqrt{k_i/m_i}$. So the microcanonical partition function can be expressed as

$$Q(N, E) = \frac{E_0}{h^{3N}} \int d^N \mathbf{p} \int d^N \mathbf{q} \delta \left(\sum_{i=1}^N \left(\frac{\mathbf{p}_i^2}{2m_i} + \frac{k_i}{2} \mathbf{r}_i^2 \right) - E \right) \quad (16)$$

The above equation can be rescaled by taking $\mathbf{J}_i = \mathbf{p}_i/\sqrt{2m_i}$ and $\mathbf{W}_i = \sqrt{k_i/2} \mathbf{r}_i$, which makes

$$\mathcal{H} = \sum_{i=1}^N (\mathbf{J}_i^2 + \mathbf{W}_i^2), \quad (17)$$

and subsequently, Eq. (16) becomes

$$Q(N, E) = E_0 \frac{2^{3N}}{h^{3N}} \sum_{i=1}^N \frac{1}{\omega_i^3} \int d^N \mathbf{J} \int d^N \mathbf{W} \delta \left(\sum_{i=1}^N (\mathbf{J}_i^2 + \mathbf{W}_i^2) - E \right). \quad (18)$$

Now defining $R^2 = \sum_{i=1}^N (\mathbf{J}_i^2 + \mathbf{W}_i^2)$, one can introduce a $6N$ -dimensional spherical coordinates consisting of one polar coordinate R and $(6N-1)$ number of angular coordinates, Ω . Here the volume element is $R^{6N-1} dR d^{6N-1} \Omega$. Therefore Eq. (18) becomes

$$Q(N, E) = E_0 \frac{2^{3N}}{h^{3N}} \sum_{i=1}^N \frac{1}{\omega_i^3} \int d^{6N-1} \Omega \int dR R^{6N-1} \delta(R^2 - E). \quad (19)$$

Using the formula, $\int d^n \omega = \frac{2\pi^{\frac{n+1}{2}}}{\Gamma(\frac{n+1}{2})}$ along with the relation, $\delta(R^2 - E) = \frac{1}{2\sqrt{E}} [\delta(R - \sqrt{E}) + \delta(R + \sqrt{E})]$ in Eq. (19), one can arrive at

$$Q(N, E) = \frac{E_0}{E} \frac{2^{3N} \pi^{3N}}{h^{3N} \Gamma(3N)} E^{3N} \sum_{i=1}^N \frac{1}{\omega_i^3}. \quad (20)$$

For large N , using the relation, $\Gamma(3N) \approx (3N)! \approx (3N)^{3N} e^{-3N}$, the above can be approximated as

$$Q(N, E) \approx \frac{E_0}{E} \frac{2^{3N} \pi^{3N}}{h^{3N} (3N)^{3N}} E^{3N} e^{3N} \sum_{i=1}^N \frac{1}{\omega_i^3}. \quad (21)$$

Ignoring the prefactor E_0/E and rearranging the terms, we obtain

$$Q(N, E) = \left(\frac{2\pi E}{3Nh} \right)^{3N} e^{3N} \prod_{i=1}^N \frac{1}{\omega_i^3}.$$

From the Boltzmann's equation, $S = k_B \ln Q$, and the definition of temperature, $\frac{1}{T} = \frac{\partial S}{\partial E}$, one can get

$$\frac{1}{T} = k_B \frac{3N}{E} \implies E = 3Nk_B T, \quad (22)$$

which is consistent with the equipartition theorem.

9. Consider a system consisting of N “indistinguishable”, non-interacting point particles (each of mass m) with total energy E and volume V . Show that the microcanonical partition function of this system is

$$Q(N, V, E) = \frac{1}{N!} \left[\frac{V}{h^3} \left(\frac{4\pi m E}{3N} \right)^{3/2} \right]^N e^{3N/2}.$$

Then derive the equation of state for this system.

Answer: The Hamiltonian for N particles (each of mass m) is

$$H = \sum_{i=1}^N \frac{\mathbf{p}_i^2}{2m}.$$

For a system consisting of N particles with total energy E and volume V , the microcanonical partition function is given by

$$Q(N, V, E) = \frac{E_0}{N! h^{3N}} \int d\mathbf{p}_1 \int d\mathbf{p}_2 \cdots \int d\mathbf{p}_N \int d\mathbf{q}_1 \int d\mathbf{q}_2 \cdots \int d\mathbf{q}_N \delta \left(\sum_{i=1}^N \frac{\mathbf{p}_i^2}{2m} - E \right), \quad (23)$$

where $\int d\mathbf{q}_i = \int dx_i \int dy_i \int dz_i$. The integration over q_i is straightforward and can be done easily as $\int d\mathbf{q}_i = V$. So Eq. (23) simplifies to

$$Q(N, V, E) = \frac{E_0 V^N}{N! h^{3N}} \int d\mathbf{p}_1 \int d\mathbf{p}_2 \cdots \int d\mathbf{p}_N \delta \left(\sum_{i=1}^N \frac{\mathbf{p}_i^2}{2m} - E \right). \quad (24)$$

Taking $\Phi_i^2 = \mathbf{p}_i^2/2m$, the above can be rewritten as

$$Q(N, V, E) = \frac{E_0 (2m)^{3N/2} V^N}{N! h^{3N}} \int d\Phi_1 \int d\Phi_2 \cdots \int d\Phi_N \delta \left(\sum_{i=1}^N \Phi_i^2 - E \right). \quad (25)$$

Eq. (25) can be calculated in an easier way if it is to be done in the polar coordinate. For that, we take, $r^2 = \sum_{i=1}^N \Phi_i^2$, where r is the radial coordinate of a $3N$ -dimensional spherical coordinates with $(3N-1)$ number of angular coordinates (ω). So the volume element is $r^{3N-1} dr d^{3N-1}\omega$. Now we can express Eq. (25) as

$$Q(N, V, E) = \frac{E_0 (2m)^{3N/2} V^N}{N! h^{3N}} \int d^{3N-1}\omega \int_0^\infty dr r^{3N-1} \delta(r^2 - E). \quad (26)$$

Using the formula, $\int d^n \omega = \frac{2\pi^{\frac{n+1}{2}}}{\Gamma(\frac{n+1}{2})}$, where $\Gamma(n) = \int_0^\infty dy y^{n-1} e^{-y}$, one can perform the integration over ω , which transforms Eq. (26) to

$$\begin{aligned} Q(N, V, E) &= \frac{E_0 (2m)^{3N/2} 2\pi^{3N/2} V^N}{N! h^{3N} \Gamma(3N/2)} \int_0^\infty dr r^{3N-1} \delta(r^2 - E) \\ &= \frac{E_0 (2m)^{3N/2} 2\pi^{3N/2} V^N}{N! h^{3N} \Gamma(3N/2)} \int_0^\infty dr r^{3N-1} \frac{1}{2\sqrt{E}} [\delta(r - \sqrt{E}) + \delta(r + \sqrt{E})] \\ &= \frac{E_0 (2m)^{3N/2} \pi^{3N/2} V^N}{E N! h^{3N} \Gamma(3N/2)} E^{3N/2} = \frac{1}{N!} \frac{E_0}{E} \frac{1}{\Gamma(3N/2)} \left[V \left(\frac{2\pi m E}{h^2} \right)^{3/2} \right]^N. \end{aligned} \quad (27)$$

For the above calculation, we used the fact that $r \in (0, \infty)$, and $E > 0$, so the second term in the integrand has no contribution in the second step. In the large N limit, one can use the Sterling's approximation, $\Gamma(n+1) = e^{-n} n^n$. Applying this, one obtains, $\Gamma(3N/2) = e^{-3N/2+1} (3N/2 - 1)^{3N/2-1} \approx e^{-3N/2} (3N/2)^{3N/2}$. Also the prefactor E_0/E can be neglected. Taking all these results into account, we get

$$Q(N, V, E) = \frac{1}{N!} \left[\frac{V}{h^3} \left(\frac{4\pi m E}{3N} \right)^{3/2} \right]^N e^{3N/2}. \quad (28)$$

The pressure is calculated from the partition function, using the formula

$$P = k_B T \left(\frac{\partial \ln Q(N, V, E)}{\partial V} \right)_{N, E}.$$

Since $Q(N, V, E) \sim V^N$, one gets $P = k_B T N/V$, or

$$PV = nRT,$$

where $n = N/N_A$. This is the equation of state of the system.

10. The canonical partition function of a monoatomic ideal gas is

$$Q(N, V, T) = \frac{1}{N!} \left(\frac{2\pi m k_B T}{h^2} \right)^{3N/2} V^N.$$

Derive expressions for the pressure and the energy from this partition function. Also show that the ideal gas equation of state is obtained if Q is of the form: $f(T)V^N$, where $f(T)$ is any function of temperature.

Answer: The canonical partition function of a monoatomic ideal gas is

$$Q(N, V, T) = \frac{1}{N!} \left(\frac{2\pi m k_B T}{h^2} \right)^{3N/2} V^N. \quad (29)$$

The pressure P is given by $P = k_B T \left(\frac{\partial \ln Q(N, V, T)}{\partial V} \right)_{N, T}$. So using Eq. (29), one gets

$$\begin{aligned} P &= k_B T N/V + k_B T \frac{\partial}{\partial V} \ln \left(\frac{1}{N!} \left(\frac{2\pi m k_B T}{h^2} \right)^{3N/2} V^N \right) \\ PV &= nRT. \end{aligned} \quad (30)$$

The average energy can be calculated using the formula, $\bar{E} = k_B T^2 \left(\frac{\partial \ln Q(N, V, T)}{\partial T} \right)_{N, V}$. So from Eq. (29), one has

$$\begin{aligned}\bar{E} &= k_B T^2 \frac{\partial}{\partial T} \left(\ln \left(\frac{2\pi m k_B T}{h^2} \right)^{3N/2} \right) + \cancel{k_B T^2 \frac{\partial}{\partial T} \left(\ln \left(\frac{V^N}{N!} \right) \right)}^0 \\ \bar{E} &= \frac{3}{2} N k_B T = \frac{3}{2} n R T.\end{aligned}\quad (31)$$

If the partition function is of the form: $Q(T, V) = f(T)V^N$, the pressure can be calculated as $P = k_B T \left(\frac{\partial \ln Q(N, V, T)}{\partial V} \right)_{N, T} = k_B T N/V$, which leads to ideal gas equation of state, $PV = nRT$.

11. The partition function of an ideal gas of diatomic molecules in an external electric field ξ is

$$Q(N, V, T, \xi) = \frac{[q(V, T, \xi)]^N}{N!}$$

where

$$q(V, T, \xi) = V \left(\frac{2\pi m k_B T}{h^2} \right)^{3/2} \left(\frac{8\pi^2 I k_B T}{h^2} \right) \left(\frac{e^{-h\nu/2k_B T}}{1 - e^{-h\nu/2k_B T}} \right) \left(\frac{k_B T}{\mu \xi} \right) \sinh \left(\frac{\mu \xi}{k_B T} \right).$$

Here I is the moment of inertia of the molecule; ν is its fundamental vibrational frequency, and μ is its dipole moment. Using this partition function along with the thermodynamic relation,

$$dA = -SdT - pdV - Md\xi,$$

where $M = N\bar{\mu}$, $\bar{\mu}$ is the average dipole moment of a molecule in the direction of the external field ξ , show that

$$\bar{\mu} = \mu \left[\coth \left(\frac{\mu \xi}{k_B T} \right) - \left(\frac{k_B T}{\mu \xi} \right) \right].$$

Sketch this result versus ξ from $\xi = 0$ to $\xi = \infty$, and interpret it.

There is a typo in the question.

Answer: From the thermodynamic relation, $dA = -SdT - pdV - Md\xi$, one can find the magnetization M as

$$M = - \left(\frac{\partial A}{\partial \xi} \right)_{V, T}.$$

Here A is the Helmholtz free energy, and it is related to partition function via the formula: $A = -k_B T \ln Q(N, V, T, \xi)$. With the given partition function, one can see

$$\ln Q(N, V, T, \xi) = N \ln \left(\frac{k_B T}{\mu \xi} \right) + N \ln \left[\sinh \left(\frac{\mu \xi}{k_B T} \right) \right] + g(T, V, N),$$

where $g[T, V, N] = N \ln \left[V \left(\frac{2\pi m k_B T}{h^2} \right)^{3/2} \left(\frac{8\pi^2 I k_B T}{h^2} \right) \left(\frac{e^{-h\nu/2k_B T}}{1 - e^{-h\nu/k_B T}} \right) \right] - \ln N!$. So we get

$$\begin{aligned} M &= -N k_B T \frac{\left(\frac{k_B T}{\mu \xi^2} \right)}{\left(\frac{k_B T}{\mu \xi} \right)} + N k_B T \left(\frac{\mu}{k_B T} \right) \frac{\cosh \left(\frac{\mu \xi}{k_B T} \right)}{\sinh \left(\frac{\mu \xi}{k_B T} \right)} \\ &= N \mu \left[\coth \left(\frac{\mu \xi}{k_B T} \right) - \left(\frac{k_B T}{\mu \xi} \right) \right]. \end{aligned} \quad (32)$$

Therefore,

$$\bar{\mu} = \frac{M}{N} = \mu \left[\coth \left(\frac{\mu \xi}{k_B T} \right) - \left(\frac{k_B T}{\mu \xi} \right) \right].$$

Interpretation: At low electric field, $\bar{\mu} = 0$, as spins are randomly oriented. With the

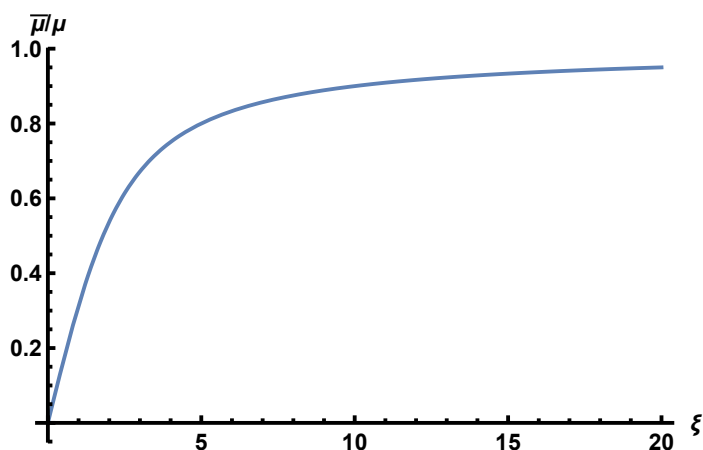


Figure 1: Plot of average dipole moment, $\bar{\mu}$ (rescaled by μ) as a function of electric field, ξ . In the given plot, we have taken $k_B T / \mu = 1$.

application of electric field, spins are continuously aligned in the field, thereby showing positive moment. At some point, all spins get aligned with the field, and the average moment saturates to μ .

12. An *approximate* partition function for a dense gas is

$$Q(N, V, T) = \frac{1}{N!} \left(\frac{2\pi m k_B T}{h^2} \right)^{3N/2} (V - Nb)^N e^{aN^2/Vk_B T},$$

where a and b are constants that are given in terms of molecular parameters. Calculate the equation of state from this partition function. What equation of state is this? Calculate the thermodynamic energy and the heat capacity.

Answer: The pressure of a gas can be calculated by using the formula, $P = k_B T \left(\frac{\partial \ln Q(N, V, T)}{\partial V} \right)_{N, T}$.

The given partition function is of the form of

$$Q(N, V, T) = f(N, T) (V - Nb)^N e^{aN^2/Vk_B T},$$

where $f(N, T) = \frac{1}{N!} \left(\frac{2\pi mk_B T}{h^2} \right)^{3N/2}$. Therefore,

$$P = k_B T \left(\frac{\partial \ln Q(N, V, T)}{\partial V} \right)_{N, T} = \frac{Nk_B T}{V - Nb} - k_B T \frac{aN^2}{V^2 k_B T}. \quad (33)$$

Rearranging the terms, we have

$$\left(P + \frac{aN^2}{V^2} \right) (V - Nb) = Nk_B T,$$

which is the equation state for Van der Waals gas.

The average energy can be computed as

$$\begin{aligned} \bar{E} &= k_B T^2 \left(\frac{\partial \ln Q(N, V, T)}{\partial T} \right)_{N, V} \\ &= k_B T^2 \frac{3N}{2} \left(\frac{\partial \ln T}{\partial T} \right)_{N, V} + k_B T^2 \left(\frac{\partial}{\partial T} \right)_{N, V} \left(\frac{aN^2}{V k_B T} \right) \\ &= \frac{3}{2} Nk_B T - \frac{aN^2}{V}. \end{aligned} \quad (34)$$

The heat capacity is given by

$$C_v = \left(\frac{\partial \bar{E}}{\partial T} \right)_{N, V} = \frac{3}{2} Nk_B. \quad (35)$$

13. A Material consists of n independent particles and is in a weak external field H . Each particle can have a magnetic moment $m\mu$ along the magnetic field, where $m = j, j-1, \dots, -j+1, -j$, j being an integer, and μ is a constant. The system is at temperature T .

- (i) Find the partition function for this system.
- (ii) Calculate the average magnetization, \bar{M} , of the material.
- (iii) For large values of T , find an asymptotic expression for \bar{M} .

Answer: The interaction energy of j -th moment with the field H is $\epsilon_j = -j\mu H$. So the partition function of a particle can be given as

$$\begin{aligned} q(T, H) &= \sum_{m=-j}^{+j} e^{\frac{m\mu H}{k_B T}} = \frac{e^{-\frac{j\mu H}{k_B T}} \left(1 - e^{\frac{(2j+1)\mu H}{k_B T}} \right)}{1 - e^{\frac{\mu H}{k_B T}}} \\ &= \frac{e^{-\frac{j\mu H}{k_B T} - \frac{\mu H}{2k_B T}} - e^{\frac{j\mu H}{k_B T} + \frac{\mu H}{2k_B T}}}{e^{-\frac{\mu H}{2k_B T}} - e^{\frac{\mu H}{2k_B T}}} \\ &= \frac{\sinh \left(\left(j + \frac{1}{2} \right) \frac{\mu H}{k_B T} \right)}{\sinh \left(\frac{1}{2} \frac{\mu H}{k_B T} \right)} \end{aligned} \quad (36)$$

For the system of n independent particles, the partition function is

$$Q(n, T, H) = q^n(T, H) = \left[\frac{\sinh \left((j + \frac{1}{2}) \frac{\mu H}{k_B T} \right)}{\sinh \left(\frac{1}{2} \frac{\mu H}{k_B T} \right)} \right]^n. \quad (37)$$

The average magnetization, \bar{M} can be expressed as

$$\bar{M} = - \left(\frac{\partial A}{\partial H} \right)_{n, T}, \quad (38)$$

where $A = -k_B T \ln Q(n, T, H)$ is the free energy. Thus,

$$\begin{aligned} \bar{M} &= nk_B T \left(\frac{\partial}{\partial H} \right)_{n, T} \left[\ln \left(\sinh \left((j + \frac{1}{2}) \frac{\mu H}{k_B T} \right) \right) - \ln \left(\sinh \left(\frac{\mu H}{2k_B T} \right) \right) \right] \\ &= nk_B T \frac{\mu}{k_B T} \left[(j + \frac{1}{2}) \frac{\cosh \left((j + \frac{1}{2}) \frac{\mu H}{k_B T} \right)}{\sinh \left((j + \frac{1}{2}) \frac{\mu H}{k_B T} \right)} - \frac{1}{2} \frac{\cosh \left(\frac{\mu H}{2k_B T} \right)}{\sinh \left(\frac{\mu H}{2k_B T} \right)} \right] \\ &= \frac{n\mu}{2} \left[(2j + 1) \coth \left((2j + 1) \frac{\mu H}{2k_B T} \right) - \coth \left(\frac{\mu H}{2k_B T} \right) \right]. \end{aligned} \quad (39)$$

For large values of T , i.e., in the limit $k_B T \gg \mu H$, one can approximate the hyperbolic function as $\coth \left(\frac{\mu H}{k_B T} \right) \approx \frac{1}{\frac{\mu H}{k_B T}} + \frac{1}{3} \frac{\mu H}{k_B T}$. Using this, Eq. (39) becomes

$$\bar{M} \approx \frac{n\mu}{2} \frac{1}{6} \frac{\mu H}{k_B T} [(2j + 1)^2 - 1] = \frac{n}{3} j(j + 1) \frac{\mu^2 H}{k_B T}. \quad (40)$$

14. The entropy of an ideal paramagnet in a magnetic field is given *approximately* by

$$S = S_0 - CU^2,$$

where U is the energy of the spin system and C is a constant with fixed mechanical parameters of the system.

- (i) Using fundamental definition of the temperature, determine the energy U of the spin system as a function of T .
- (ii) Sketch a graph of U versus T for all values of T ($-\infty < T < \infty$).
- (iii) Briefly tell what physical sense you can make of the negative temperature part of your result.

Answer: From the definition of temperature, $\frac{1}{T} = \frac{\partial S}{\partial U}$, one gets

$$\frac{1}{T} = -2CU \implies U = -\frac{1}{2CT},$$

where $C > 0$.

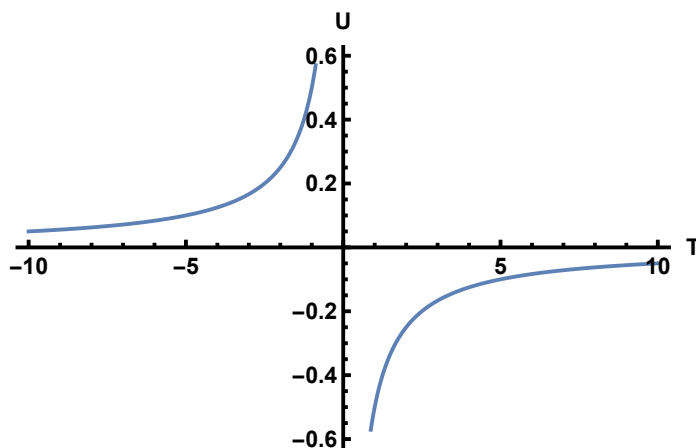


Figure 2: Plot of U versus T . Here, $C = 1$.

Interpretation: The negative temperature signifies that there are more number of particles in an excited state than the ground state (population inversion). For a spin system, one may have a situation where most of the spins are aligned anti-parallel to the applied field, thereby resulting negative temperature.

15. A one-dimensional quantum harmonic oscillator (whose ground state energy is $\hbar\omega/2$) is in thermal equilibrium with a heat bath at temperature T .

- (i) What is the mean value of the oscillator's energy, $\langle E \rangle$, as a function of T ?
- (ii) What is the value of ΔE , the root mean square fluctuation in energy about $\langle E \rangle$?
- (iii) How do $\langle E \rangle$ and ΔE behave in the limits $k_B T \leq \hbar\omega$ and $k_B T \gg \hbar\omega$?

Answer: The energy of ν -th state of a harmonic oscillator with angular frequency ω is given by $\epsilon_\nu = (\nu + \frac{1}{2}) \hbar\omega$. Therefore the partition function can be expressed as

$$q(T) = \sum_{\nu=0}^{\infty} e^{-\frac{\hbar\omega}{k_B T}(\nu + \frac{1}{2})} = \frac{e^{-\frac{\hbar\omega}{2k_B T}}}{1 - e^{-\frac{\hbar\omega}{k_B T}}} = \frac{2}{\sinh\left(\frac{\hbar\omega}{2k_B T}\right)} \quad (41)$$

So the mean value of energy is

$$\bar{E} = k_B T^2 \left(\frac{\partial \ln q(T)}{\partial T} \right) = +k_B T^2 \frac{\cosh\left(\frac{\hbar\omega}{2k_B T}\right)}{\sinh\left(\frac{\hbar\omega}{2k_B T}\right)} \left(\frac{\hbar\omega}{2k_B T^2} \right) = \frac{\hbar\omega}{2} \coth\left(\frac{\hbar\omega}{2k_B T}\right). \quad (42)$$

The root mean square fluctuation in energy, ΔE is computed as

$$\Delta E = \sqrt{k_B T^2 \frac{\partial \bar{E}}{\partial T}} = \sqrt{k_B T^2 \frac{\hbar\omega}{2} \left(\frac{\hbar\omega}{2k_B T^2} \right) \text{csch}^2\left(\frac{\hbar\omega}{2k_B T}\right)} = \frac{\hbar\omega}{2 \sinh\left(\frac{\hbar\omega}{2k_B T}\right)}. \quad (43)$$

In the limit $k_B T \ll \hbar\omega$, $\coth\left(\frac{\hbar\omega}{2k_B T}\right) = \frac{e^{\frac{\hbar\omega}{2k_B T}} + e^{-\frac{\hbar\omega}{2k_B T}}}{e^{\frac{\hbar\omega}{2k_B T}} - e^{-\frac{\hbar\omega}{2k_B T}}} \approx \frac{e^{\frac{\hbar\omega}{2k_B T}}}{e^{\frac{\hbar\omega}{2k_B T}}} \approx 1$, and $\sinh\left(\frac{\hbar\omega}{2k_B T}\right) = \frac{e^{\frac{\hbar\omega}{2k_B T}} - e^{-\frac{\hbar\omega}{2k_B T}}}{2} \approx \frac{1}{2}e^{\frac{\hbar\omega}{2k_B T}}$. So in this limit, $\bar{E} \approx \frac{1}{2}\hbar\omega$ and $\Delta E \approx \hbar\omega e^{-\frac{\hbar\omega}{2k_B T}}$.

In the limit $k_B T \gg \hbar\omega$, $\coth\left(\frac{\hbar\omega}{2k_B T}\right) = \frac{\cosh\left(\frac{\hbar\omega}{2k_B T}\right)}{\sinh\left(\frac{\hbar\omega}{2k_B T}\right)} \approx \frac{1 + \mathcal{O}(1/T^2)}{\left(\frac{\hbar\omega}{2k_B T}\right) + \mathcal{O}(1/T^3)} \approx \frac{2k_B T}{\hbar\omega}$, and $\sinh\left(\frac{\hbar\omega}{2k_B T}\right) \approx \frac{\hbar\omega}{2k_B T}$. Therefore, $\bar{E} \approx k_B T$ and $\Delta E \approx k_B T$.

Class 16: (04.03.2021)

Systems of interacting particles

Lattice vibrations and Normal modes:

Solid \rightarrow made of N atoms

Each atom can be described by its mass and position co-ordinates

To describe the displacements from the equilibrium we introduce the variable

$$\xi_{i\alpha} \equiv x_{i\alpha} - x_{i\alpha}^{(0)}; \quad \alpha = 1, 2, 3$$

\nwarrow Eq^m

The kinetic energy of vibration of the solid

i th atom
 $\hookrightarrow m_i$

r_i
 \downarrow

(x_{i1}, x_{i2}, x_{i3})
Like (x_i, y_i, z_i)

$\dot{x}_{i\alpha} \rightarrow \alpha$ component
of vel. of
 i th atom

$$K = \frac{1}{2} \sum_{i=1}^N \sum_{\alpha=1}^3 m_i \dot{x}_{i\alpha}^2 = \frac{1}{2} \sum_{i=1}^N \sum_{\alpha=1}^3 m_i \dot{\xi}_{i\alpha}^2$$

"Potential energy" $V(x_{11}, x_{12}, \dots, x_{N3})$

Potential energy in eq^m position (Can be expressed as a Taylor series (since the displacements are small))

$$V = V_0 + \sum_{i\alpha} \left(\frac{\partial V}{\partial x_{i\alpha}} \right)_0 \xi_{i\alpha} + \frac{1}{2} \sum_{i\alpha, j\gamma} \underbrace{\left(\frac{\partial^2 V}{\partial x_{i\alpha} \partial x_{j\gamma}} \right)_0}_{A_{i\alpha, j\gamma}} \xi_{i\alpha} \xi_{j\gamma} + \dots$$

i or j from 1 to N
 α or γ from 1 to 3

\downarrow
 "0"
 Evaluated at the equilibrium positions
 $x_{i\alpha} = x_{i\alpha}^{(0)}$

$$V = V_0 + \frac{1}{2} \sum_{i\alpha, j\gamma} A_{i\alpha, j\gamma} \xi_{i\alpha} \xi_{j\gamma}$$

$$H = \frac{1}{2} \sum_{i\alpha} m_i \dot{\xi}_{i\alpha}^2 + V_0 + \frac{1}{2} \sum_{i\alpha, j\gamma} A_{i\alpha, j\gamma} \xi_{i\alpha} \xi_{j\gamma}$$

\downarrow
 Hamiltonian

interacting

"Complicated Hamiltonian"

Change of variables \rightarrow eliminates the cross terms in the potential energy

"Classical Mechanics"

$\{x_i\}^{3N} \rightarrow q_r^{3N}$ (generalized co-ordinate)

$$\{x_i\} = \sum_{r=1}^{3N} B_{i,r} q_r$$

\rightarrow A proper choice of coefficients gives the following Hamiltonian

$$H = V_0 + \frac{1}{2} \sum_{r=1}^{3N} \dot{q}_r^2 + \frac{1}{2} \sum_{r=1}^{3N} \omega_r^2 q_r^2$$

$\omega_r^2 \rightarrow$ positive constants

$q_r \rightarrow$ normal co-ordinates

$$H_r = \frac{1}{2} (\dot{q}_r^2 + \omega_r^2 q_r^2) \xrightarrow{QM} E_r = (n_r + \frac{1}{2}) \hbar \omega_r$$

$$E_{n_1, \dots, n_{3N}} = V_0 + \sum_{r=1}^{3N} (n_r + \frac{1}{2}) \hbar \omega_r$$

independent of quantum no.

$$= -N\eta + \sum_{r=1}^{3N} n_r \hbar \omega_r$$

$$V_0 + \frac{1}{2} \sum_{r=1}^{3N} \hbar \omega_r$$

\rightarrow zero point energy

$\eta \rightarrow$ binding energy per atom in the solid at absolute zero

Partition function:

$$Q = \sum_{n_1, n_2, \dots, n_{3N}} e^{-\beta \left[-N\eta + \sum_{r=1}^{3N} \hbar \omega_r n_r \right]} \rightarrow E_{n_1, n_2, \dots}$$

$$Q = e^{\beta N \eta} \sum_{n_1, n_2, \dots, n_{3N}} e^{-\hbar \beta (n_1 \omega_1 + n_2 \omega_2 + \dots + n_{3N} \omega_{3N})}$$

$$= e^{\beta N \eta} \left(\sum_{n_1=0}^{\infty} e^{-\hbar \beta n_1 \omega_1} \right) \left(\sum_{n_2=0}^{\infty} e^{-\hbar \beta n_2 \omega_2} \right) \dots \left(\sum_{n_{3N}=0}^{\infty} e^{-\hbar \beta n_{3N} \omega_{3N}} \right)$$

$$Q = e^{\beta N \eta} \left(\frac{1}{1 - e^{-\beta \hbar \omega_1}} \right) \left(\frac{1}{1 - e^{-\beta \hbar \omega_2}} \right) \dots \left(\frac{1}{1 - e^{-\beta \hbar \omega_{3N}}} \right)$$

$$\ln Q = \beta N \eta - \sum_{r=1}^{3N} \ln(1 - e^{-\beta \hbar \omega_r})$$

If normal mode frequencies are closely spaced (ω_r)

$$\ln Q = \beta N \eta - \int_0^{\infty} \ln(1 - e^{-\beta \hbar \omega}) \underbrace{\sigma(\omega) d\omega}_e \rightarrow \text{A knowledge is required}$$

the no. of normal
modes with angular
frequency in the range
betⁿ ω and $\omega + d\omega$

$$\langle E \rangle = - \frac{\partial \ln Q}{\partial \beta} = -N\eta + \int_0^{\infty} \frac{(0 + \hbar \omega e^{-\beta \hbar \omega})}{(1 - e^{-\beta \hbar \omega})} \sigma(\omega) d\omega$$

$$\langle E \rangle = -N\eta + \int_0^{\infty} \frac{\hbar \omega}{(e^{\beta \hbar \omega} - 1)} \sigma(\omega) d\omega$$

$$\langle E \rangle = -N\eta + \int_0^{\infty} \frac{\hbar\omega}{(e^{\beta\hbar\omega} - 1)} \sigma(\omega) d\omega$$

$$C_V = \left(\frac{\partial \langle E \rangle}{\partial T} \right)_V = - \frac{1}{k_B T^2} \left(\frac{\partial \langle E \rangle}{\partial \beta} \right)_V = -k_B \beta^2 \left(\frac{\partial \langle E \rangle}{\partial \beta} \right)_V$$

$$= -k_B \beta^2 \left[0 + \int_0^{\infty} \frac{(\hbar\omega)(0 - \hbar\omega e^{\beta\hbar\omega})}{(e^{\beta\hbar\omega} - 1)^2} \sigma(\omega) d\omega \right]$$

$$C_V = k_B \int_0^{\infty} \frac{(\beta\hbar\omega)^2 \sigma(\omega) e^{\beta\hbar\omega}}{(e^{\beta\hbar\omega} - 1)^2} d\omega$$

$$\frac{d}{dx} \left(\frac{f(x)}{g(x)} \right) = \frac{f'(x)g(x) - g'(x)f(x)}{g(x)^2}$$

$$C_V = k_B \int_0^{\infty} \sigma(\omega) d\omega$$

if T is large
 β is small \nearrow "Dulong-Petit's result"

$$C_V = k_B \int_0^{\infty} \frac{(\beta\hbar\omega)^2 \sigma(\omega) (1 + \beta\hbar\omega + \dots)}{(1 + \beta\hbar\omega + \dots - 1)^2} d\omega$$

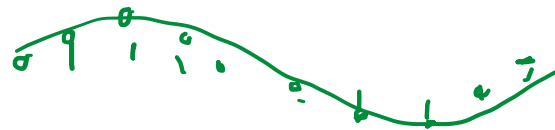
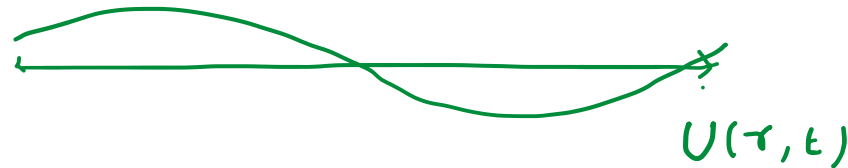
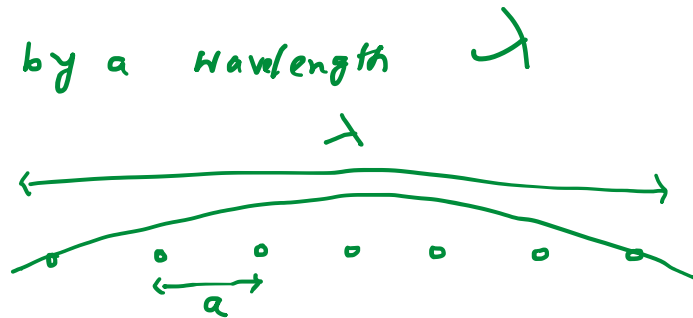
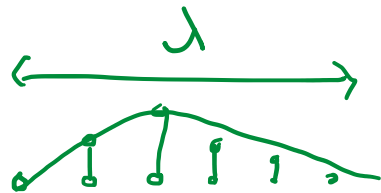
"Debye"

Solid \equiv Neglecting the discreteness of the atoms
and treating the latter as if it were
a continuous isotropic elastic medium

"a" \rightarrow mean interatomic separation in solid ($\sim 10^0$)

Elastic medium is characterised by a wavelength λ

$$\lambda \gg a$$



$$\lambda \sim a$$

Discreteness
becomes
important



$$\omega = c_s k$$

velocity of sound \downarrow 'Wave Vector'

$$\sigma_c(\omega) d\omega = 3 \frac{V}{(2\pi)^3} (4\pi k^2 dk) = 3 \frac{V}{(2\pi^3 c_s^3)} \omega^2 d\omega = \sigma(\omega) d\omega$$

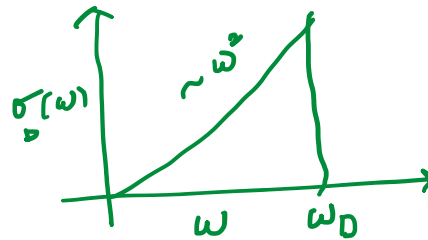
three possible polarizations

2 - transverse
1 - longitudinal

Debye approximation

$$\sigma_D(\omega) = \sigma_c(\omega) \left. \begin{array}{l} \omega < \omega_D \\ \omega > \omega_D \end{array} \right\} = 0$$

$\omega < \omega_D$
 $\omega > \omega_D$



Debye frequency spectrum

$$\int_0^{\infty} \sigma_D(\omega) d\omega = \int_0^{\omega_D} \sigma_c(\omega) d\omega = 3N$$

$$\left(\frac{3V}{2\pi^3 c_s^3} \right) \int_0^{\omega_D} \omega^2 d\omega = 3N$$

$$\Rightarrow \left(\frac{3V}{2\pi^3 c_s^3} \right) \frac{\omega_D^3}{3} = 3N$$

\Rightarrow

"Debye frequency"

$$\omega_D = \left(6\pi^2 \frac{N}{V} \right)^{\frac{1}{3}} c_s$$

$$c_s = 5 \times 10^5 \text{ cm/s}$$

$$\left(\frac{V}{N} \right)^{\frac{1}{3}} = a = 10^{-8} \text{ cm}$$

$$\omega_D \sim 10^{14} \text{ sec}^{-1}$$

"infrared"

Different for different solids

$$C_V = k_B \int_0^{\omega_D} \frac{e^{\beta \hbar \omega}}{(e^{\beta \hbar \omega} - 1)^2} (\beta \hbar \omega)^2 \left(\frac{3V}{2\pi^2 c_s^3} \right) \omega^3 d\omega$$

$$= k_B \int_0^{\beta \hbar \omega_D} \frac{e^x x^2}{(e^x - 1)^2} \left(\frac{3V}{2\pi^2 c_s^3} \right) \frac{1}{(\beta \hbar)^3} x^3 dx$$

$$C_V = k_B \frac{3V}{2\pi^2 (c_s \beta \hbar)^3} \int_0^{\beta \hbar \omega_D} \frac{e^x x^4}{(e^x - 1)^2} dx$$

$$C_V = k_B \frac{3 \left(\frac{3}{4\pi^2} N \left(\frac{c_s}{\omega_D} \right)^3 \right)}{4\pi^2 (c_s \beta \hbar)^3} \int_0^{\beta \hbar \omega_D} \frac{x^4 e^x}{(e^x - 1)^2} dx$$

$$C_V = (3N k_B) f_D \left(\frac{\Theta_D}{T} \right)$$

"Dulong-Petit" result

$$\Theta_D = \frac{\hbar \omega_D}{k_B}$$

$T \rightarrow \infty$

$$\frac{3}{(\beta \hbar \omega_D)^3} \int_0^{\beta \hbar \omega_D} \frac{x^4 (1+x+\dots)}{(1+x+\dots-1)^2} dx$$

$$= \frac{3}{(\beta \hbar \omega_D)^3} \int_0^{\beta \hbar \omega_D} x^4 dx \quad \uparrow$$

$$= (3N k_B) \frac{3}{(\beta \hbar \omega_D)^3} \int_0^{\beta \hbar \omega_D} \frac{x^4 e^x}{(e^x - 1)^2} dx$$

$$V = \frac{3N (2\pi^2 c_s^3)}{\omega_D^3}$$

$$V = 6\pi^2 N \left(\frac{c_s}{\omega_D} \right)^3$$

(F. Reif
Fundamentals
of Statistical
and Thermal
physics)

$f_D(\beta \hbar \omega_D)$
Debye f^n

$$\frac{3}{(\beta \hbar \omega_D)^3} \int_0^{\beta \hbar \omega_D} \frac{x^4 e^x}{(e^x - 1)^2} dx$$

$x = \beta \hbar \omega$

At high temperature? $T \gg 1$

$$\boxed{C_V = 3N k_B} \quad f_D(x) \xrightarrow{x \rightarrow 0} 1$$

Low temperature

$$\omega \ll \omega_0$$

$$\beta \hbar \omega \gg 1$$

$$\frac{\hbar \omega}{k_B T} \gg 1$$

$$C_V = \frac{k_B \times \text{constants}}{(\beta \hbar)^3}$$

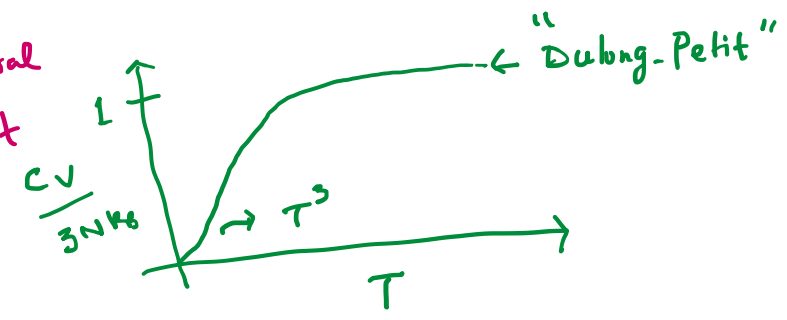
$$\left\{ \int_0^{\beta \hbar \omega_0 \rightarrow \infty} \frac{x^4 e^x}{(e^x - 1)^2} dx \right\}$$

$$\textcircled{I} \rightarrow I = \frac{4\pi^4}{15} \text{ (check.)}$$

$$\omega < \omega_0 \quad \sigma(\omega) \neq 0$$

Definite integral
= Constant

$C_V \sim T^3$ at low temperature
(Experimentally verified)



$C_V \sim 3Nk_B$ at high temperature

$$C_V = \left(\frac{2\pi^2}{5} \right) (V k_B) \left(\frac{k_B T}{\epsilon_s \hbar} \right)^3$$

$$\int_0^{\infty} \frac{x^4 e^x}{(e^x - 1)^2} dx = \int_0^{\infty} \frac{x^3}{(e^x - 1)} dx = \frac{4\pi^4}{15}$$

$\Theta_D(k) \rightarrow$ either from
specific heat
or elastic constants.

(F. Reif's book)

NaCl 308
KCl 230

Einstein's formulation

N no. of non-interacting independent harmonic oscillators
- each with frequency ω .

$$C_V = k_B \int_0^{\infty} \frac{e^{\beta \hbar \omega}}{(e^{\beta \hbar \omega} - 1)^2} (\beta \hbar \omega)^2 \sigma(\omega) d\omega$$

"General" but
in the normal mode
picture

Einstein

$$C_V = 3Nk_B (\beta \hbar \omega)^2 \frac{e^{\beta \hbar \omega}}{(e^{\beta \hbar \omega} - 1)^2}$$

$$= 3Nk_B \left(\frac{\Theta_E}{T}\right)^2 \frac{e^{\Theta_E/T}}{(e^{\Theta_E/T} - 1)^2}$$

$$\Theta_E = \frac{\hbar \omega}{k_B}$$

↓
"Einstein temp"

$T \rightarrow \infty$

$$C_V = 3Nk_B \left(\frac{\Theta_E}{T}\right)^2 \frac{(1 + \cancel{\Theta_E/T} + \dots)}{(1 + \Theta_E/T + \dots - 1)^2}$$

$$\approx 3Nk_B \left(\frac{\Theta_E}{T}\right)^2 \frac{1}{(\Theta_E/T)^2} = 3Nk_B$$

(Dulong-Petit's result)

Low temperature

$T \rightarrow 0$

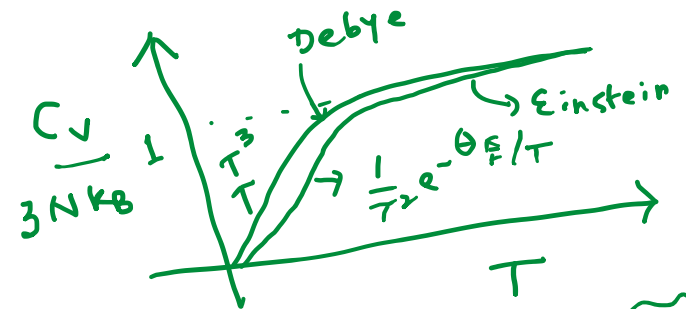
$\frac{\theta_E}{T} \gg 1$

$$C_V = 3Nk_B \left(\frac{\theta_E}{T}\right)^2 \frac{e^{-\theta_E/T}}{e^{2\theta_E/T}}$$

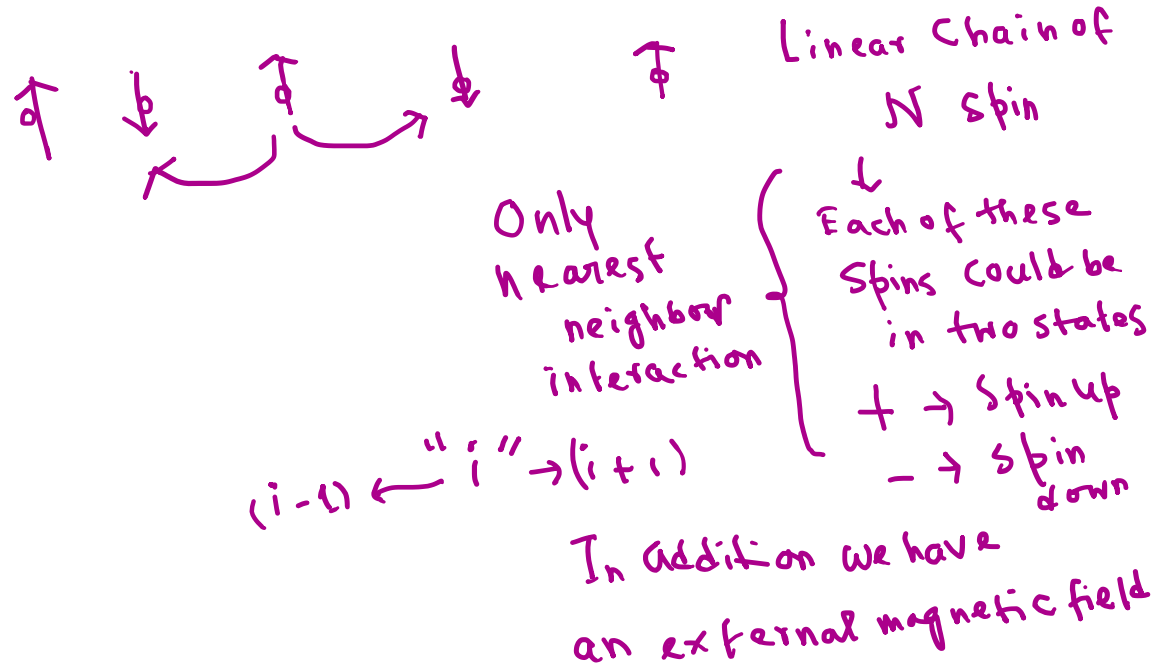
$$C_V = 3Nk_B \left(\frac{\theta_E}{T}\right)^2 e^{-\theta_E/T}$$

$$(e^{\theta_E/T} - 1)^{-2} \approx (e^{\theta_E/T})^{-2}$$

(Not T^3 !!!)



"Ising Model"
1925 (1-Dimension)



Hamiltonian

$$H = -J \sum_{i=1}^N \sigma_i \sigma_{i+1} - \sum_{i=1}^N B \sigma_i$$

Nearest neighbor interaction

External magnetic field

Mag. moment

$$B = \frac{\hbar \mu}{J} > 0$$

(interaction strength)

$$\sigma_i = \pm 1$$

$\uparrow \uparrow$

$$H = -J(1 \times 1) = -J$$

Energy is lowered

$\downarrow \downarrow$

$$H = -J(-1 \times -1) = -J$$

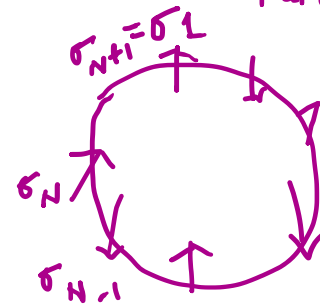
"ferromagnetic state"

All spins are up or
all spins are down

Canonical partition function:

$$Q(N, B, T) = \sum_{\substack{\sigma_1 \\ = \pm 1}} \sum_{\substack{\sigma_2 \\ = \pm 1}} \dots \sum_{\substack{\sigma_N \\ = \pm 1}} e^{+\beta \left(J \sum_{i=1}^N \sigma_i \sigma_{i+1} + B \sum_{i=1}^N \sigma_i \right)}$$

"Put a Periodic boundary Condition"



$$\sigma_{N+1} = \sigma_1$$

$$Q(B, T) = \sum_{\sigma_1} \sum_{\sigma_2} \cdots \sum_{\sigma_N} \exp \left[\beta \left(J \sum_{i=1}^N \sigma_i \sigma_{i+1} + \frac{B}{2} \sum_{i=1}^N (\sigma_i + \sigma_{i+1}) \right) \right]$$

$$Q(B, T) = \sum_{\{\sigma_i\}} \prod_{i=1}^N \exp \left[\beta \left(J \sigma_i \sigma_{i+1} + \frac{B}{2} (\sigma_i + \sigma_{i+1}) \right) \right]$$

$M_{\sigma_i \sigma_{i+1}}$ (Matrix elements)

$M \rightarrow 2 \times 2$ Matrix
with matrix elements

$$M_{++} \quad M_{--} \quad M_{+-} \quad M_{-+}$$

$$Q(B, T) = \sum_{\{\sigma_i\}} \prod_{i=1}^N M_{\sigma_i \sigma_{i+1}}$$

$$Q(B, T) = \sum_{\sigma_1} \sum_{\sigma_2} \sum_{\sigma_3} \cdots M_{\sigma_1 \sigma_2} M_{\sigma_2 \sigma_3} \cdots M_{\sigma_N \sigma_1}$$

$$M_{\sigma_i \sigma_{i+1}} = \langle \sigma_i | M | \sigma_{i+1} \rangle$$

Basis $\{\sigma_i\}$

Like quantum Mechanics

"Matrix Representation of Operators"

$$\hat{A} = \begin{pmatrix} A_{11} & A_{12} \\ A_{21} & A_{22} \end{pmatrix}$$

$$A_{11} = \langle \psi_1 | A | \psi_1 \rangle$$

$$A_{12} = \langle \psi_1 | A | \psi_2 \rangle$$

$$Q(B, T) = \sum_{\sigma_1} \sum_{\sigma_2} \dots \sum_{\sigma_N} M_{\sigma_1 \sigma_2} M_{\sigma_2 \sigma_3} \dots M_{\sigma_N \sigma_1}$$

$$Q(B, T) = \sum_{\sigma_1} \left(\sum_{\sigma_2} \sum_{\sigma_3} \dots \sum_{\sigma_N} \langle \sigma_1 | M | \sigma_2 \rangle \langle \sigma_2 | M | \sigma_3 \rangle \langle \sigma_3 | \dots \langle \sigma_{N-1} | M | \sigma_N \rangle \langle \sigma_N | M | \sigma_1 \rangle \right)$$

$$\sum_i |\sigma_i\rangle \langle \sigma_i| = I$$

↓
Resolution of identity

$$Q(B, T) = \sum_{\sigma_1} \langle \sigma_1 | M^N | \sigma_1 \rangle = \text{Tr}(M^N)$$

↳ "trace"

$$M_{\sigma_i \sigma_{i+1}} = \exp \left[\beta \left(J \sigma_i \sigma_{i+1} + \frac{B}{2} (\sigma_i + \sigma_{i+1}) \right) \right]$$

$$M = \begin{pmatrix} M_{++} & M_{+-} \\ M_{-+} & M_{--} \end{pmatrix} = \begin{pmatrix} \exp[\beta(J+B)] & e^{-\beta J} \\ e^{-\beta J} & \exp[\beta(J-B)] \end{pmatrix} = \begin{pmatrix} e^{x+y} & e^{-x} \\ e^{-x} & e^{x-y} \end{pmatrix}$$

Let us call $\beta J = x$ $\beta B = y$

$$M^\dagger = M \text{ (check)}$$

M is Hermitian
 ↳ One should be able to diagonalize it + real eigenvalues

$$U^{-1} M U \equiv D = \begin{bmatrix} \lambda_+ & 0 \\ 0 & \lambda_- \end{bmatrix} \quad \lambda_{\pm} \rightarrow \text{eigen values}$$

$$\underbrace{U U^{-1} M U U^{-1}}_D = M = U D U^{-1}$$

$$Q = \text{Tr}\{M^N\} = \text{Tr}\{(U D U^{-1})^N\}$$

$$= \text{Tr}\{U D \underbrace{U^{-1} U}_I D U^{-1} \dots U D U^{-1} \dots\}$$

$$= \text{Tr}\{U D^N U^{-1}\}$$

$$= \text{Tr}\{D^N U^{-1} U\} = \text{Tr}\{D^N\} = \lambda_+^N + \lambda_-^N$$

$$\text{Tr}(A) = (\lambda_1 + \lambda_2 + \dots) \quad \text{Sum of eigen values}$$

$$\text{Tr}(A^2) = (\lambda_1^2 + \lambda_2^2 + \dots)$$

$$\text{Tr}(A^N) = (\lambda_1^N + \lambda_2^N + \dots)$$

$$A = \begin{pmatrix} \lambda_1 & 0 \\ 0 & \lambda_2 \end{pmatrix} \rightarrow \text{Tr}(A) = \lambda_1 + \lambda_2$$

$$A^2 = \begin{pmatrix} \lambda_1 & 0 \\ 0 & \lambda_2 \end{pmatrix} \begin{pmatrix} \lambda_1 & 0 \\ 0 & \lambda_2 \end{pmatrix} = \begin{pmatrix} \lambda_1^2 & 0 \\ 0 & \lambda_2^2 \end{pmatrix}$$

$$\text{Tr}(A^2) = \lambda_1^2 + \lambda_2^2$$

Calculating eigenvalues:

$$|M - \lambda I| = 0$$

$$\begin{vmatrix} e^{x+y} - \lambda & e^{-x} \\ e^{-x} & e^{x-y} - \lambda \end{vmatrix} = 0$$

$$\cosh x = \frac{1}{2}(e^x + e^{-x})$$

$$\sinh x = \frac{1}{2}(e^x - e^{-x})$$

$$\Rightarrow (e^{x+y} - \lambda)(e^{x-y} - \lambda) - e^{-2x} = 0$$

$$\Rightarrow e^{x+y+x-y} - \lambda e^{x-y} - \lambda e^{x+y} + \lambda^2 - e^{-2x} = 0$$

$$\Rightarrow \lambda^2 - \lambda(e^{x-y} + e^{x+y}) + (e^{2x} - e^{-2x}) = 0$$

$$\Rightarrow \lambda^2 - \lambda e^x(e^y + e^{-y}) + (e^{2x} - e^{-2x}) = 0$$

$$\Rightarrow \lambda^2 - 2e^x \cosh y + 2 \sinh(2x) = 0$$

$$ax^2 + bx + c = 0$$

$$x_{\pm} = \frac{-b \pm \sqrt{b^2 - 4ac}}{2a}$$

$$\lambda_{\pm} = \frac{2 \cosh y e^x \pm \sqrt{4e^{2x} \cosh^2 y - 8 \sinh(2x)}}{2}$$

$$\lambda_{\pm} = e^x \cosh y \pm \sqrt{e^{2x} \cosh^2 y - 2 \sinh(2x)}$$

$$= e^x \cosh y \pm \sqrt{e^{2x}(1 + \sinh^2 y) - \frac{2}{2} e^{2x} + e^{-2x}}$$

$$\lambda_{\pm} = e^x \cosh y \pm \sqrt{e^{2x} \sinh^2 y + e^{-2x}}$$

$$Q = (\lambda_+)^N + (\lambda_-)^N$$

$$= \lambda_+^N$$

$\lambda_+ \gg \lambda_-$
 $N \rightarrow \infty$ (thermodynamic limit)

$$\frac{1}{N} \ln Q(B, T) = \ln \lambda_+ + \frac{1}{N} \ln \left(1 + \left(\frac{\lambda_-}{\lambda_+} \right)^N \right) \sim \ln \lambda_+$$

$$Q = (\lambda_+)^N = \left\{ e^x \cosh \gamma + \sqrt{e^{2x} \sinh^2 \gamma + e^{-2x}} \right\}^N$$

$$Q = \left\{ e^{\beta J} \cosh(\beta B) + \sqrt{e^{2\beta J} \sinh^2(\beta B) + e^{-2\beta J}} \right\}^N$$

Helmholtz free energy
 $A = -k_B T \ln Q$

$$-k_B T \ln Q = A = -N k_B T \ln \left\{ e^{\beta J} \cosh(\beta B) + \sqrt{e^{2\beta J} \sinh^2(\beta B) + e^{-2\beta J}} \right\}$$

Average Magnetization

$$\langle M(B, T) \rangle = \left\langle \sum_i \mu \sigma_i \right\rangle = - \left(\frac{\partial A}{\partial h} \right)_T = \frac{1}{\beta} \left(\frac{\partial \ln Q}{\partial h} \right)_T = \frac{N}{\beta} \left(\frac{\partial \ln \lambda +}{\partial h} \right)_T$$

$$= N \mu \left(\frac{\partial \ln \lambda +}{\partial (\beta \mu h)} \right)$$

$$= N \mu \left(\frac{\partial \ln \lambda +}{\partial y} \right)$$

Not Matrix M
Here, magnetization

$$\langle M(B, T) \rangle = N \mu \frac{\partial}{\partial y} \left\{ \ln \left(e^x \cosh y + \sqrt{e^{2x} \sinh^2 y + e^{-2x}} \right) \right\}$$

$$= N \mu \frac{\left(e^x \sinh y + \frac{1}{2} \frac{1}{\sqrt{e^{2x} \sinh^2 y + e^{-2x}}} (e^{2x} \sinh y \cosh y + 0) \right)}{\left(e^x \cosh y + \sqrt{e^{2x} \sinh^2 y + e^{-2x}} \right)}$$

Average Magnetization

$$\langle M(B, T) \rangle = \left\langle \sum_i \mu \sigma_i \right\rangle = - \left(\frac{\partial A}{\partial h} \right)_T = \frac{1}{\beta} \left(\frac{\partial \ln Q}{\partial h} \right)_T = \frac{N}{\beta} \left(\frac{\partial \ln \lambda +}{\partial h} \right)_T$$

$$= N \mu \left(\frac{\partial \ln \lambda +}{\partial (\beta \mu h)} \right)$$

$$= N \mu \left(\frac{\partial \ln \lambda +}{\partial y} \right)$$

Not Matrix M
Here, magnetization

$$\langle M(B, T) \rangle = N \mu \frac{\partial}{\partial y} \left\{ \ln \left(e^x \cosh y + \sqrt{e^{2x} \sinh^2 y + e^{-2x}} \right) \right\}$$

$$= N \mu \frac{\left(e^x \sinh y + \frac{1}{2} \frac{1}{\sqrt{e^{2x} \sinh^2 y + e^{-2x}}} (e^{2x} \sinh y \cosh y + 0) \right)}{\left(e^x \cosh y + \sqrt{e^{2x} \sinh^2 y + e^{-2x}} \right)}$$

$$\frac{\partial}{\partial y} \cosh y = \sinh y$$

$$\frac{\partial}{\partial y} \sinh y = \cosh y$$

$$\langle M(B, T) \rangle = N \mu \frac{\left(e^x \sinh y \sqrt{e^{2x} \sinh^2 y + e^{-2x}} + e^{2x} \sinh y \cosh y \right)}{\sqrt{e^{2x} \sinh^2 y + e^{-2x}} \left(e^x \cosh y + \sqrt{e^{2x} \sinh^2 y + e^{-2x}} \right)}$$

$$\langle M(B, T) \rangle = N\mu \frac{e^x \sinh y \sqrt{e^{2x} \sinh^2 y + e^{-2x}} + e^{2x} \sinh y \cosh y}{\sqrt{e^{2x} \sinh^2 y + e^{-2x}} (e^x \cosh y + \sqrt{e^{2x} \sinh^2 y + e^{-2x}})}$$

$$\langle M(B, T) \rangle = N\mu \frac{e^x \sinh y (\cancel{\sqrt{e^{2x} \sinh^2 y + e^{-2x}}} + e^x \cosh y)}{\sqrt{e^{2x} \sinh^2 y + e^{-2x}} (\cancel{e^x \cosh y + \sqrt{e^{2x} \sinh^2 y + e^{-2x}}})}$$

$$\langle M(B, T) \rangle = N\mu \frac{e^x \sinh y}{\sqrt{e^{2x} \sinh^2 y + e^{-2x}}} = N\mu \frac{\cancel{e^x} \sinh y}{\cancel{e^x} \sqrt{\sinh^2 y + e^{-4x}}}$$

$$\langle M(B, T) \rangle = N\mu \frac{\sinh(B\beta)}{\sqrt{\sinh^2(B\beta) + e^{-4\beta J}}}$$

↑
Average magnetization

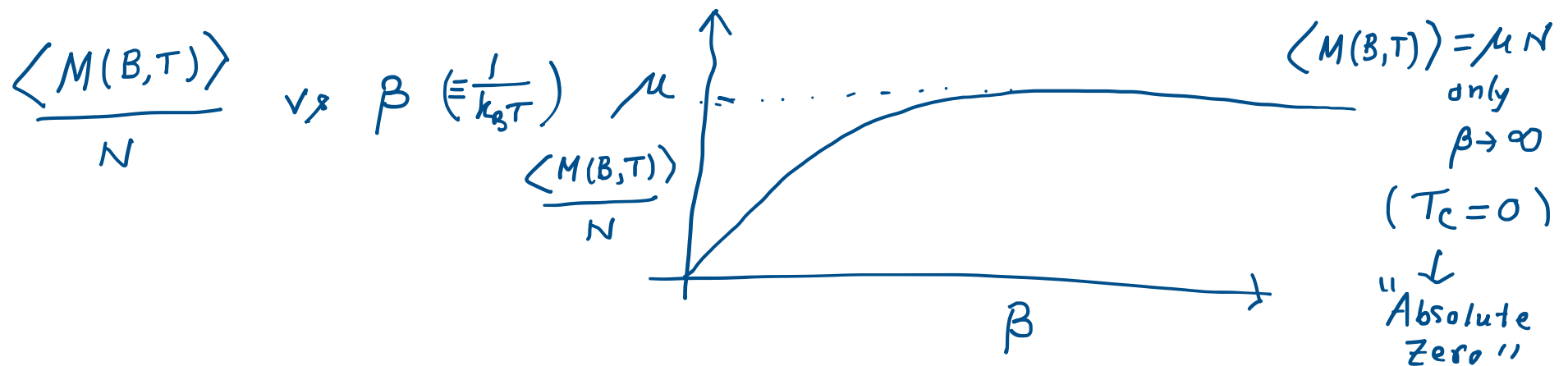
$\frac{\langle M(B, T) \rangle}{N} \Rightarrow$ Average magnetization per spin

$$\langle M(B, T) \rangle = N\mu \frac{\sinh(B\beta)}{\sqrt{\sinh^2(B\beta) + e^{-4\beta J}}}$$

If there is no external magnetic field $B=0$

$$\frac{\langle M(B, T) \rangle}{N} = 0 \quad \left(\text{No, net magnetization} \equiv \text{No. ferromagnetic phase for Ising model in 1D.} \right)$$

\hookrightarrow spontaneous at $(B=0)$ finite temperature
 \hookrightarrow for finite β



$$\langle M(B, T) \rangle = N\mu \frac{\sinh(B\beta)}{\sqrt{\sinh^2(B\beta) + e^{-4\beta J}}}$$

$J=0$ (Non-interacting) but $B \neq 0$ (there is an external magnetic field)

$$\langle M(B, T) \rangle = N\mu \tanh(B\beta)$$

$$\frac{\langle M(B, T) \rangle}{N} = \mu \tanh(B\beta)$$

$B=0$ } Non-interacting spins
 $J=0$ } and no external magnetic field

$$\beta \rightarrow \infty \quad \langle M(B, T) \rangle = N\mu \frac{\sinh(B\beta)}{\sqrt{\sinh^2(B\beta) + 0}} = N\mu \quad (\text{independent of } B \text{ any finite } B)$$

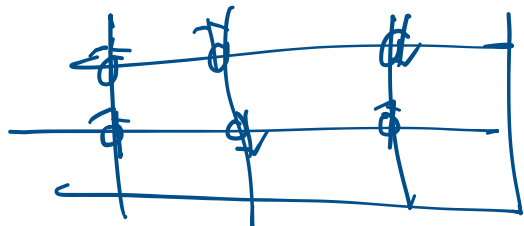
$$\sinh^2(B\beta) + 1 = \cosh^2(B\beta)$$

$J \rightarrow$ Completely aligned phase \rightarrow Never Wins in 1D
 $k_B T$ (thermal energy) \downarrow No phase transition
 \sim Disorder (maximizes entropy)

Lars Onsager 1944 ($B=0$) (Analytically)

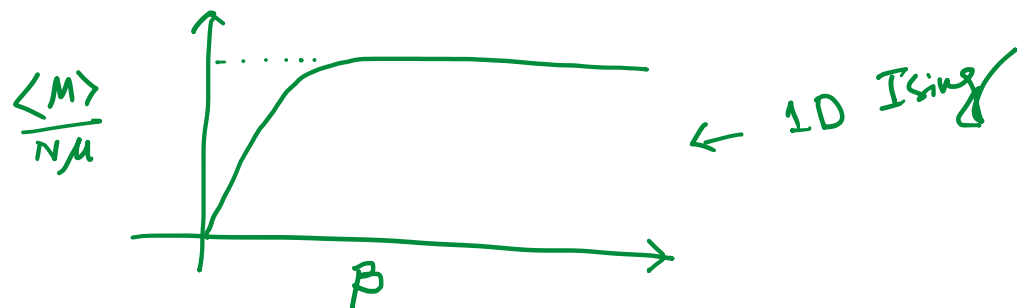
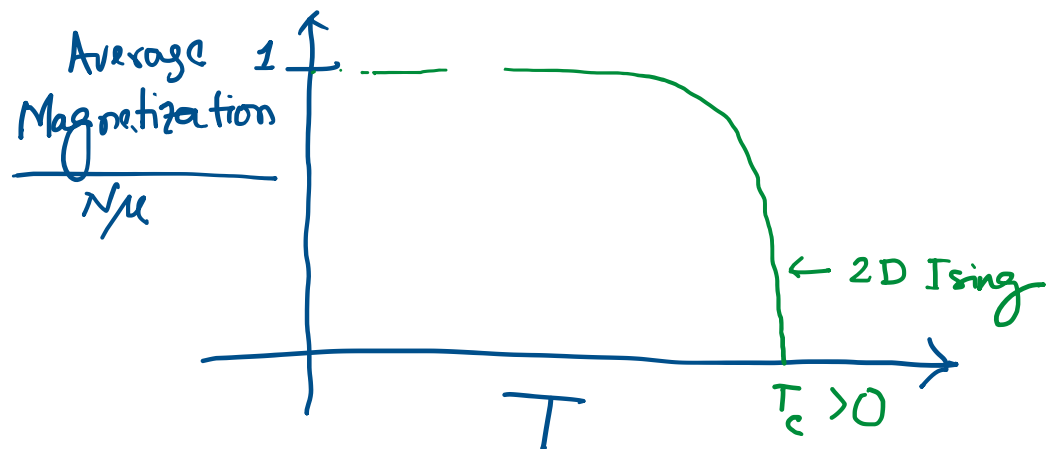
2D Ising Model

($B \neq 0$) \rightsquigarrow



\rightsquigarrow Shows a phase transition!!

\hookrightarrow Transition to a ferromagnetic phase at finite temperature.



For 1D Ising model

$T \gg 0$

$\langle M \rangle = 0$

$T \rightarrow \infty$

$\langle M \rangle \rightarrow 0$

Distribution function theory of liquids

Partition function

$$Q = \sum_j e^{-\beta E_j}$$

"Quantization"
"Energy"

QM

Purely classical system

$$H[r^N, p^N] = \underbrace{K[p^N]}_{\text{K.E.}} + \underbrace{U[r^N]}_{\text{P.E.}}$$

$r^N \equiv 3N$ position
co-ordinates

$p^N \equiv 3N$ momenta
co-ordinates

Hamiltonian
of the system

Canonical partition function for a purely classical system

$$Q = Q_{\text{classical}} = \frac{1}{N! h^{3N}} \int dr^N \int dp^N \exp[-\beta H(r^N, p^N)]$$

if we assume the particles are indistinguishable

$$\delta r^N \delta p^N \sim h^{3N}$$

$f(r^N, p^N)$ = probability distribution for observing a system at phase space point (r^N, p^N) .

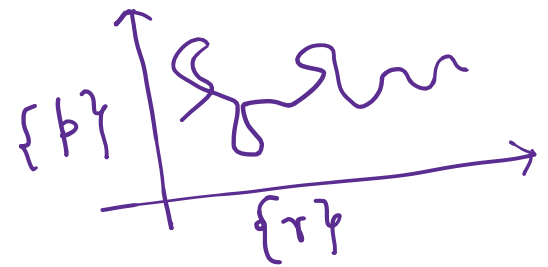
$$f(r^N, p^N) = \frac{\exp[-\beta H(r^N, p^N)]}{\int dr^N \int dp^N \exp[-\beta H(r^N, p^N)]}$$

$$= \frac{\exp[-\beta (K(p^N) + U(r^N))]}{\int dr^N \int dp^N \exp[-\beta (K(p^N) + U(r^N))]}$$

$$= \frac{\exp[-\beta K(p^N)]}{\int dp^N \exp[-\beta K(p^N)]} \times \frac{\exp[-\beta U(r^N)]}{\int dr^N \exp[-\beta U(r^N)]}$$

$\phi(p^N) \rightarrow P(r^N)$

$$f(r^N, p^N) = \phi(p^N) P(r^N)$$



$$\int dr^N \equiv \int dr_1 \int dr_2 \dots \int dr_N$$

$$\int dp^N \equiv \int dp_1 \int dp_2 \dots \int dp_N$$

$$\int dr \equiv \int dx \int dy \int dz$$

$$\int dp \equiv \int dp_x \int dp_y \int dp_z$$

$\phi(p^N) \equiv$ prob. distribution for observing
system at momentum space point p^N

$P(r^N) \equiv$ prob. distribution for observing
system at configuration space point r^N .

Reduced Configurational distribution functions:

$$P^{(2/N)}(r_1, r_2) = \int dr_3 \int dr_4 \dots \int dr_N P(r^N) \quad \searrow \quad p(r_1, r_2, \dots, r_N)$$

↓
Joint probability
distribution for
finding particle 1
at r_1 and particle
2 at r_2

↪ Specific probability
distribution (specifically requires particle 1
at r_1 and particle 2 at r_2).

generic reduced distribution functions:

$\rho^{(2/N)}(r_1, r_2) \equiv$ joint prob. distribution function
for finding a particle (anyone)
at position r_1 and any other
particle at position r_2 (in the
 N particle system)

$$\rho^{(2/N)}(r_1, r_2) = \frac{N(N-1)}{N!} P^{(2/N)}(r_1, r_2)$$

$$\rho^{(n/N)}(r_1, r_2, \dots, r_n) = \frac{N!}{(N-n)!} \int dr^{N-n} \exp[-\beta U(r^N)]$$

$$\int dr^N \exp[-\beta U(r^N)] \rightarrow Z_N \quad \left(\begin{array}{l} \text{configurational} \\ \text{integral} \end{array} \right)$$

$$Z_N \equiv V^N \quad (\text{for ideal gas})$$

$$\int dr^{N-n} \equiv \int dr_{n+1} \int dr_{n+2} \dots \int dr_N$$

$$\int dr_1 \int dr_2 \dots \int dr_N e^0$$

$$\rho^{(n/N)}(\tau_1, \tau_2, \dots, \tau_n) = \frac{N!}{(N-n)!} \frac{1}{Z_N} \int d\tau_{n+1} \int d\tau_{n+2} \dots \int d\tau_N \exp(-\beta U(\tau^N))$$

$\{\tau_1, \tau_2, \dots, \tau_N\}$

↓
"Integrate"

$$\int \dots \int \rho^{(n/N)}(\tau_1, \tau_2, \dots, \tau_n) d\tau_1 \dots d\tau_n = \frac{N!}{(N-n)!} \frac{1}{Z_N} \underbrace{\int d\tau_1 \dots \int d\tau_{n+1} \dots \int d\tau_N \exp(-\beta U(\tau_1, \tau_2, \dots, \tau_N))}_{Z_N}$$

$$\int \rho^{(n/N)}(\tau^n) d\tau^n = \frac{N!}{(N-n)!}$$

$$\int \rho^{(1/N)}(\tau_i) d\tau_i = N \quad (\text{number conservation})$$

$$\rho^{(1/N)}(r_1) = \frac{N!}{(N-1)!} \frac{1}{Z_N} \int dr_2 \int dr_3 \dots \int dr_N \exp(-\beta U(r_1, r_2 \dots r_N))$$

For homogeneous liquid

$$Z_N = \left\{ \int dr_1 \right\} \int dr_2 \dots \int dr_N e^{-\beta U(r_1, r_2 \dots r_N)}$$

$$\int dr_1 = V \int dr_2 \int dr_3 \dots \int dr_N e^{-\beta U(r^N)}$$

$$\rho^{(1/N)}(r) = \frac{N}{Z_N} \frac{Z_N}{V} = \frac{N}{V} = \rho \text{ (density)}$$

$$\begin{aligned} \rho^{(2/N)}(r_1, r_2) &= \frac{N!}{(N-2)!} \frac{1}{Z_N} \int dr_3 \int dr_4 \dots \int dr_N e^{-\beta U(r^N)} \\ &= \frac{N(N-1)}{Z_N} \int dr_3 \int dr_4 \dots \int dr_N e^{-\beta U(r^N)} \end{aligned}$$

In case of ideal gas (fluid with interparticle interaction)

$$\rho^{(2/N)}(r_1, r_2) = \frac{N(N-1)}{\underbrace{Z_N}_{V^N}} V^{N-2} = N(N-1) \frac{V^{N-2}}{V^N} = \frac{N(N-1)}{V^2} \approx \frac{N^2}{V^2} = \rho^2$$

"N is very large"

$$= \rho^{(1/N)}(r_1) \rho^{(1/N)}(r_2) \quad (U=0)$$

Let us define,

$$g(r_1, r_2) = \frac{\rho^{(2/N)}(r_1, r_2)}{\rho^2}$$

$$g(r_1, r_2) = \rho^r / \rho^2 = 1 \quad (\text{ideal gas})$$

Also,
$$h(r_1, r_2) = \frac{(\rho^{(2/N)}(r_1, r_2) - \rho^2)}{\rho^2} = g(r_1, r_2) - 1$$

$\rho \equiv$ Bulk density

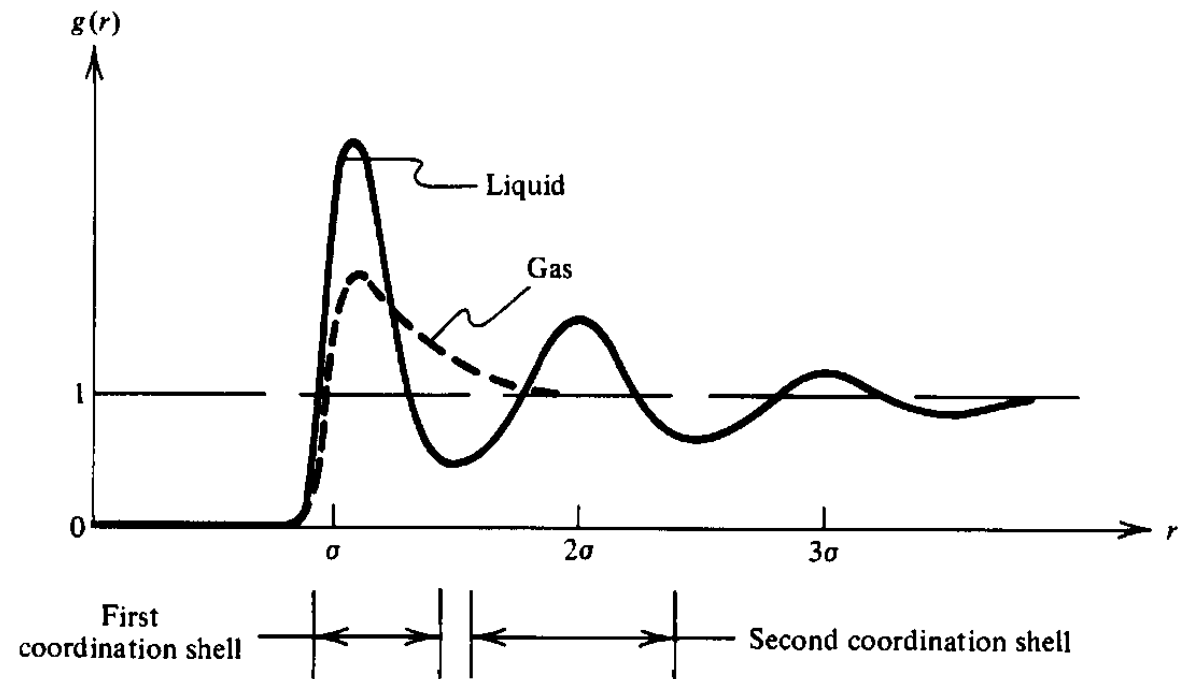
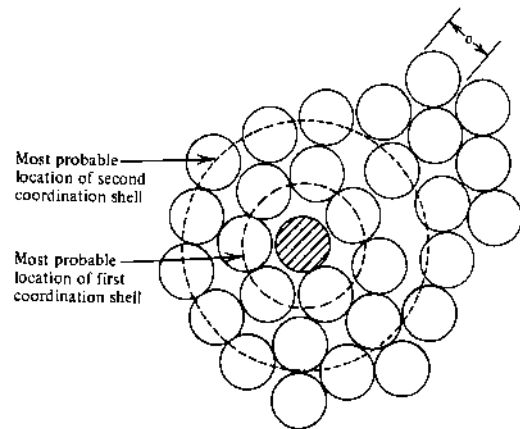
$g(r_1, r_2)$
 \hookrightarrow Pair-correlation fn / radial distribution function

$$\rho^{(2/N)}(0, r) = \rho^2 g(r)$$

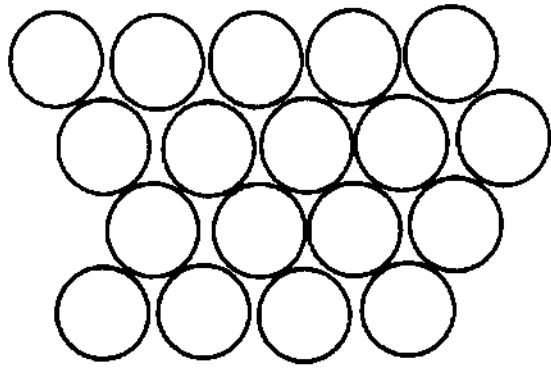
$$\frac{\rho^{(2/N)}(0, r)}{\rho} = \rho g(r) = \text{conditional prob. that a particle will be found at } r \text{ given that another is at the origin}$$

Liquid Structure $\equiv g(r)$

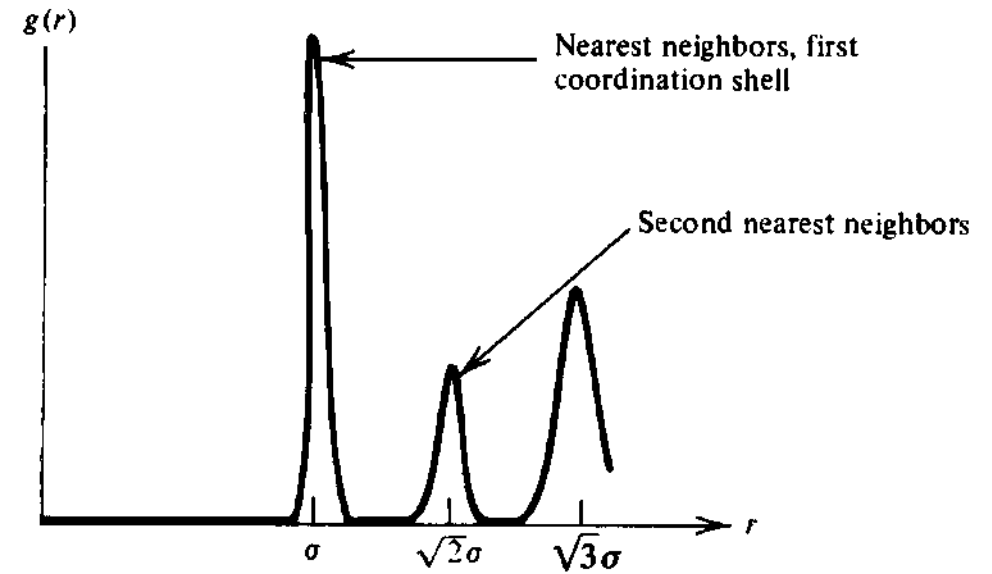
\equiv average density of particles at r given that a tagged particle is at the origin



David Chandler \rightarrow Introduction to SM



Crystalline order: solid



Highly ordered solid

$$p^{(2/N)}(r_1, r_2) = \frac{N!}{(N-2)!} \frac{1}{Z_N} \int dr_3 \int dr_4 \dots \int dr_N e^{-\beta U(r^N)}$$

Using Dirac Delta fⁿ $\rightarrow N(N-1)$

$$p^{(1/N)}(r) = \left\langle \sum_{i=1}^N \delta(r-r_i) \right\rangle$$

Ensemble average
of microscopic
particle density

Let us consider

$$\langle \delta(r-r_1) \delta(r'-r_2) \rangle = \frac{1}{Z_N} \int \delta(r-r_1) \delta(r'-r_2) e^{-\beta U(r_1, r_2, \dots, r_N)} dr^N \dots \dots \textcircled{A}$$

\downarrow
(dr₁ ... dr_N)

$$\left\langle \sum_{i=1}^N \sum_{j=1}^{N'} \delta(r-r_i) \delta(r'-r_j) \right\rangle = p^{(2/N)}(r_1, r_2)$$

\downarrow
N(N-1) x terms like \textcircled{A}

Prime means i ≠ j

$$Z_N = \int \int e^{-\beta U(r_1, r_2, \dots, r_N)} dr_1 dr_2 \dots dr_N$$

$$\begin{aligned} \langle X \rangle &= \sum_{i=1}^N X_i P_i \\ &= \frac{\int P(x) X dx}{\int P(x) dx} \end{aligned}$$

$$\left\langle \frac{1}{N} \sum_{i=1}^N \sum_{j=1}^{N'} \delta(\mathbf{r} - \mathbf{r}_j + \mathbf{r}_i) \right\rangle$$

$$= \left\langle \frac{1}{N} \sum_{i=1}^N \sum_{j=1}^{N'} \underbrace{\delta(\mathbf{r}' + \mathbf{r} - \mathbf{r}_j)}_{\substack{f(\mathbf{r}' + \mathbf{r} - \mathbf{r}_j) \\ \parallel \\ f(\mathbf{r}_i + \mathbf{r} - \mathbf{r}_j) \equiv \delta(\mathbf{r}_i + \mathbf{r} - \mathbf{r}_j)}} \underbrace{\delta(\mathbf{r}' - \mathbf{r}_i)}_{\delta(\mathbf{r}' - \mathbf{r}_i)} d\mathbf{r}' \right\rangle$$

$$= \frac{1}{N} \int \left\langle \sum_{i=1}^N \sum_{j=1}^{N'} \delta(\mathbf{r}' + \mathbf{r} - \mathbf{r}_j) \delta(\mathbf{r}' - \mathbf{r}_i) \right\rangle d\mathbf{r}'$$

$$= \frac{1}{N} \int \rho^{(2/A)}(\mathbf{r}' + \mathbf{r}, \mathbf{r}') d\mathbf{r}' = \frac{1}{N} \rho^2 \int g(\mathbf{r}, \mathbf{r}') d\mathbf{r}'$$

$$= \frac{\rho^2}{N} \left(\frac{N}{\rho} \right) \underset{\int d\mathbf{r}'}{g(\mathbf{r})} = \rho g(\mathbf{r})$$

$$\int \delta(\mathbf{r}) d\mathbf{r} = 1$$

$$\int \delta(\mathbf{r} - \mathbf{r}') d\mathbf{r}' = 1$$

$$\int f(\mathbf{r}) \delta(\mathbf{r} - \mathbf{r}') d\mathbf{r}' = f(\mathbf{r})$$

$$\int f(\mathbf{r}' + \mathbf{r} - \mathbf{r}_j) \delta(\mathbf{r}' - \mathbf{r}_i) d\mathbf{r}'$$

$$= f(\mathbf{r}_i + \mathbf{r} - \mathbf{r}_j)$$

$$\rho = N/V$$

Homogeneous and isotropic liquid

$$S(k) = \frac{1}{N} \langle \rho(-k) \rho(k) \rangle$$

↓
Structure
factor

(Using X-ray
or neutron
scattering)

$$S(k) = \left\langle \frac{1}{N} \sum_{i=1}^N \sum_{j=1}^N \exp(-ik \cdot (r_i - r_j)) \right\rangle$$

$$= \left\langle \frac{1}{N} N + \frac{1}{N} \sum_{i=1}^N \sum_{j \neq i}^N \exp(-ik \cdot (r_i - r_j)) \right\rangle = 1 + \left\langle \frac{1}{N} \sum_{i=1}^N \sum_{j \neq i}^N \exp(-ik \cdot (r_i - r_j)) \right\rangle$$

$$= 1 + \left\langle \frac{1}{N} \sum_{i=1}^N \sum_{j \neq i}^N \int \exp(-ik \cdot (r - r')) \delta(r - r_i) \delta(r' - r_j) dr dr' \right\rangle \rightarrow \rho^{(2/N)}(r - r')$$

$$\rho(k) = \int dr e^{-ik \cdot r} \rho(r)$$

↳ F.T of $\rho(r)$

"Fourier Transformation"

$$= \int dr e^{-ik \cdot r} \sum_{i=1}^N \delta(r - r_i)$$

$$\rho(k) = \sum_{i=1}^N e^{-ik \cdot r_i}$$

$$S(k) = 1 + \frac{1}{N} \iint \exp(-ik \cdot (r-r')) \underbrace{\rho^{(2/N)}(r-r')}_{\rho^2 g(r-r')} dr dr'$$

Since we have a
homogeneous liquid
(isotropic)

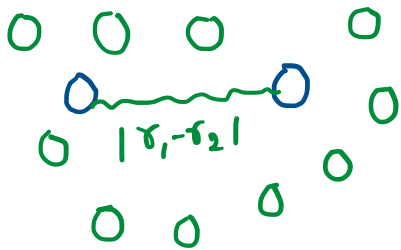
$$S(k) = 1 + \frac{1}{N} \rho^2 V \int g(r) \exp(-ik \cdot r) dr \quad \Rightarrow \quad \rho^2 g(r-r')$$

$$S(k) = 1 + \rho \int g(r) \exp(-ik \cdot r) dr$$

Structure
factor in
X-ray or
neutron scattering

~ F.T of $g(r)$

Reversible work theorem:



Force acting on particle 1

↳ Due to all other particles

$$\left\langle - \frac{dU(r^N)}{dr_1} \right\rangle_{r_1, r_2 \text{ fixed}}$$

$$- \left\langle \frac{dU(r^N)}{dr_1} \right\rangle_{r_1, r_2 \text{ fixed}} = - \frac{\int dr_3 \dots dr_N \left(\frac{dU}{dr_1} \right) e^{-\beta U(r^N)}}{\int dr_3 \dots dr_N e^{-\beta U(r^N)}}$$

$$= (k_B T) \frac{\left[\frac{d}{dr_1} \int dr_3 \dots dr_N e^{-\beta U(r^N)} \right]}{\int dr_3 \dots dr_N e^{-\beta U(r^N)}}$$

$$\begin{aligned} \frac{d}{dr_1} e^{-\beta U} &= -\beta e^{-\beta U} \left(\frac{dU}{dr_1} \right) \\ -\frac{1}{\beta} \frac{d}{dr_1} e^{-\beta U} &= e^{-\beta U} \left(\frac{dU}{dr_1} \right) \end{aligned}$$

$$-\left\langle \frac{d}{dr_1} U(r^N) \right\rangle_{r_1, r_2 \text{ fixed}} = \frac{k_B T \left[\frac{d}{dr_1} \int dr_3 \dots dr_N e^{-\beta U(r^N)} \right]}{\int dr_3 \dots dr_N e^{-\beta U(r^N)}}$$

$$= k_B T \frac{d}{dr_1} \left(\ln \int dr_3 \dots dr_N e^{-\beta U(r^N)} \right)$$

$$= k_B T \frac{d}{dr_1} \ln \left[\frac{N(N-1) \int dr_3 \dots dr_N e^{-\beta U(r^N)}}{\int dr^N e^{-\beta U(r^N)}} \right]$$

$$-\frac{dW(r)}{dr} = k_B T \frac{d}{dr_1} \ln g(r_1, r_2)$$

Potential of mean force

Reversible work done

$$W(r) = -k_B T \ln g(r)$$

Helmholtz free energy

$$r = r_1 - r_2$$

(N, V, T)

How to model pair potential

$$U(r^N) \equiv U(r_1, r_2, \dots, r_N) \leftarrow \text{N-body potential.}$$

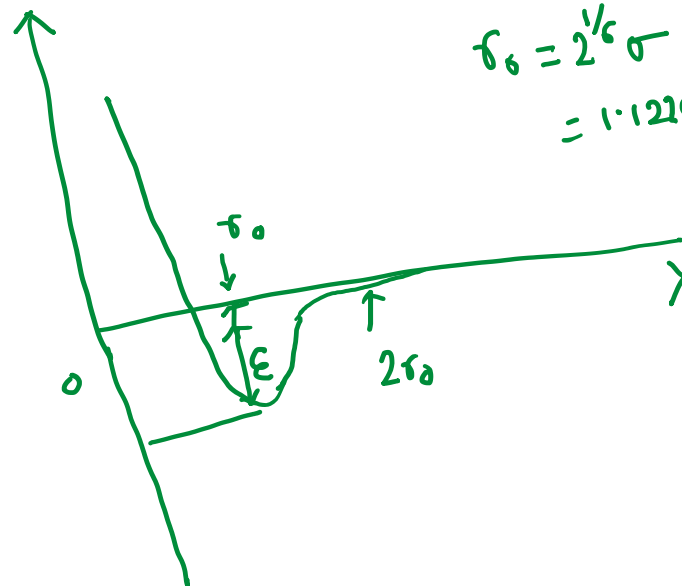
$$= \sum_{i>j=1}^N U(|r_i - r_j|) \quad \text{"Sum of pair potentials"}$$

A good example of
this pair potential

→ Lennard-Jones potential (not charged system)

$$U(r) = 4\epsilon \left[\left(\frac{\sigma}{r} \right)^{12} - \left(\frac{\sigma}{r} \right)^6 \right]$$

↖ Repulsion ↘ attraction



at $r = r_0 \quad U(r_0) = -\epsilon$

$$r_0 = 2^{1/6} \sigma = 1.1225 \sigma$$

Average internal energy

$$\begin{aligned}
 \langle E \rangle &= \langle K(\mathbf{p}^N) \rangle + \langle U(\mathbf{r}^N) \rangle \\
 &= N \langle p^2/2m \rangle + \left\langle \sum_{i,j=1}^N u(r_{ij}) \right\rangle \\
 &= N \frac{3k_B T}{2} + \dots
 \end{aligned}$$

How many distinct pairs
 $\frac{N(N-1)}{2} \leftarrow$ for not to
 have double
 counting

$$\begin{aligned}
 \left\langle \sum_{i,j=1}^N u(r_{ij}) \right\rangle &= \frac{N(N-1)}{2} \langle u(r_{12}) \rangle \\
 &= \frac{N(N-1)}{2} \frac{\int d\mathbf{r}^N u(r_{12}) e^{-\beta U(\mathbf{r}^N)}}{\int d\mathbf{r}^N e^{-\beta U(\mathbf{r}^N)}} \rightarrow \rho^{(2/N)}(r_1, r_2) \\
 &= \frac{1}{2} \frac{\int d\mathbf{r}_1 \int d\mathbf{r}_2 u(r_{12}) \underbrace{N(N-1) \int d\mathbf{r}^{N-2} e^{-\beta U(\mathbf{r}^N)}}_{Z_N}}{\int d\mathbf{r}^N e^{-\beta U(\mathbf{r}^N)}} = \frac{1}{2} \int d\mathbf{r}_1 \int d\mathbf{r}_2 \rho^{(2/N)}(r_1, r_2) u(r_{12})
 \end{aligned}$$

$$\left\langle \sum_{i,j=1}^N u(r_{ij}) \right\rangle = \frac{1}{2} \int d\mathbf{r}_1 \int d\mathbf{r}_2 \rho^{(2/N)}(\mathbf{r}_1, \mathbf{r}_2) U(r_{12})$$

$$= \frac{1}{2} \int d\mathbf{r}_1 \int d\mathbf{r}_2 \rho^2 g(r_{12}) U(r_{12}) = \frac{V}{2} \left(\frac{N}{V}\right)^2 \int d\mathbf{r}_2 g(r_{12}) U(r_{12})$$

$$\left[\int d\mathbf{r}_1 \int d\mathbf{r}_2 \equiv \int \downarrow \int d\mathbf{r}_{12} \right]$$

$$\left\langle \sum_{i,j=1}^N u(r_{ij}) \right\rangle = \frac{N}{2} \int \rho g(r) u(r) dr$$

There are $4\pi r^2 \rho g(r) dr$ neighbors in a shell of radius r and thickness dr and the energy of interaction betn the central particles and these neighbors is $U(r)$.

$$\frac{\langle E \rangle}{N} = \frac{3}{2} k_B T + \rho/2 \int dr g(r) u(r)$$

Note that $\frac{\Delta E}{N} = \frac{\partial}{\partial \beta} (\beta \Delta A / N)$

$$-\beta \Delta A = \ln(Q / Q_{\text{ideal}})$$

$$\frac{\Delta E}{N} = \rho / 2 \int d\mathbf{r} g(r) u(r)$$

low density case $g(r) \approx e^{-\beta u(r)}$

$$W(r) = u(r) + \Delta W(r) \leftarrow \begin{array}{l} \text{due to} \\ \text{surrounding} \\ \text{particle} \end{array}$$

↓
pair potential

$$W(r) = u(r) = -k_B T \ln g(r)$$

$$\Delta E = E - \bar{E}_{\text{ideal}}$$

$$= \frac{3}{2} N k_B T + \frac{N}{2} \rho \int g(r) u(r) d\mathbf{r} - \frac{3}{2} N k_B T$$

$$= \frac{N}{2} \rho \int g(r) u(r) d\mathbf{r}$$

$$= \rho / 2 \int d\mathbf{r} e^{-\beta u(r)} u(r) [1 + \underbrace{O(\rho)}_{\substack{\text{very} \\ \text{small} \\ \approx 0}}]$$

(~ 0 at low density)

$$\frac{\Delta E}{N} = \rho/2 \int d\mathbf{r} e^{-\beta u(\mathbf{r})} u(\mathbf{r})$$

$$\frac{\partial}{\partial \beta} (\beta \Delta A / N) = \rho/2 \int d\mathbf{r} e^{-\beta u(\mathbf{r})} u(\mathbf{r})$$

Integrating over β

$$\beta \frac{\Delta A}{N} = \rho/2 \int d\mathbf{r} u(\mathbf{r}) \int_0^\beta d\beta' e^{-\beta' u(\mathbf{r})} = \rho/2 \int d\mathbf{r} u(\mathbf{r}) \left. \frac{e^{-\beta' u(\mathbf{r})}}{-u(\mathbf{r})} \right|_0^\beta$$

$$-\beta \frac{\Delta A}{N} = \rho/2 \int d\mathbf{r} f(\mathbf{r}) + O(\rho^2)$$

$$f(\mathbf{r}) = 1 - e^{-\beta u(\mathbf{r})}$$

$$p = k_B T \left(\frac{\partial \ln Q}{\partial V} \right)_{N, T} \quad (\text{recall})$$

$$= - \left(\frac{\partial A}{\partial V} \right)_{N, T}$$

$$A = -k_B T \ln Q$$

$$p_{\text{ideal}} = \rho k_B T$$

$$V = N/\rho$$

$$\frac{\partial}{\partial V} = -\rho^2/N \frac{\partial}{\partial \rho}$$

$$\rho = N/V$$

$$\begin{aligned}
 P^2 \frac{\partial}{\partial P} \left(\frac{\beta \Delta A}{N} \right) &= P^2 \frac{\partial}{\partial P} \left(\frac{\beta}{N} (A - A_{\text{ideal}}) \right) \\
 &= P^2 \frac{\beta}{N} \frac{\partial}{\partial P} (A - A_{\text{ideal}}) \\
 &= P^2 \frac{\beta}{N} \left(\left(\frac{\partial A}{\partial P} \right) - \left(\frac{\partial A_{\text{ideal}}}{\partial P} \right) \right) \\
 &= P^2 \frac{\beta}{N} \left(\frac{Np}{P^2} - \frac{N}{P^2} P/\beta \right)
 \end{aligned}$$

$$\Rightarrow P^2 \frac{\partial}{\partial P} \left(\frac{\beta \Delta A}{N} \right) = \beta p - P \quad \boxed{\beta p = P + P^2 B_2(T)}$$

$$\beta p = P + P^2 \frac{\partial}{\partial P} \left(\frac{\beta \Delta A}{N} \right)$$

$$\beta p = P + P^2 \frac{\partial}{\partial P} \left\{ \left(-\frac{P}{2} \right) \int dr f(r) \right\} = P + P^2 B_2(T)$$

$$B_2(T) = -\frac{1}{2} \int f(r) dr$$

→ Second virial coefficient

$$\frac{\partial}{\partial P} A_{\text{ideal}}$$

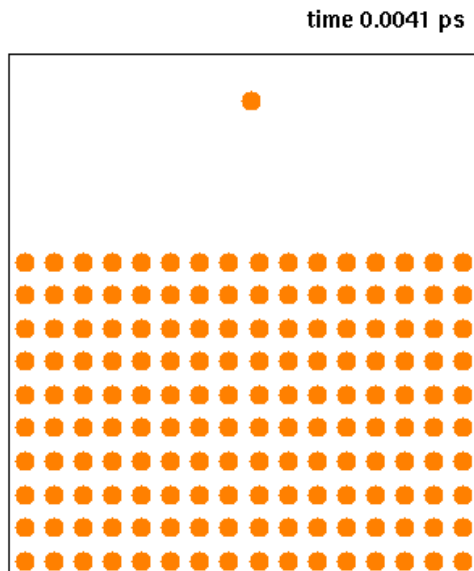
$$\begin{aligned}
 p_{\text{ideal}} &= - \left(\frac{\partial A_{\text{ideal}}}{\partial V} \right)_{N,T} \\
 &= + P^r / N \left(\frac{\partial A_{\text{ideal}}}{\partial P} \right)_{N,T}
 \end{aligned}$$

$$\left(\frac{\partial A}{\partial P} \right)_{N,T} = N / P^2 p$$

$$\left(\frac{\partial A_{\text{ideal}}}{\partial P} \right)_{N,T} = N / P^2 p_{\text{ideal}}$$

$$f(r) = \frac{e^{-\beta U(r)}}{1}$$

Molecular dynamics simulation: Theory, Algorithm, and Applications



“If, in some cataclysm, all of scientific knowledge were to be destroyed, and only one sentence passed on to the next generations of creatures, what statement would contain the most information in the fewest words? I believe it is the *atomic hypothesis* (or the *atomic fact*, or whatever you wish to call it) that *all things are made of atoms - little particles that move around in perpetual motion, attracting each other when they are a little distance apart, but repelling upon being squeezed into one another*. In that one sentence, you will see, there is an *enormous* amount of information about the world, if just a little imagination and thinking are applied.”

Richard Feynman (1961)

Source: Wikipedia

- ❑ A method to capture the physical movements of atoms and molecules dynamically for a significant period of time
- ❑ Positions of a set of interacting particles over time are calculated by solving Newton's equation of motion
- ❑ The forces between the particles and their potential energies are calculated using molecular mechanics force field

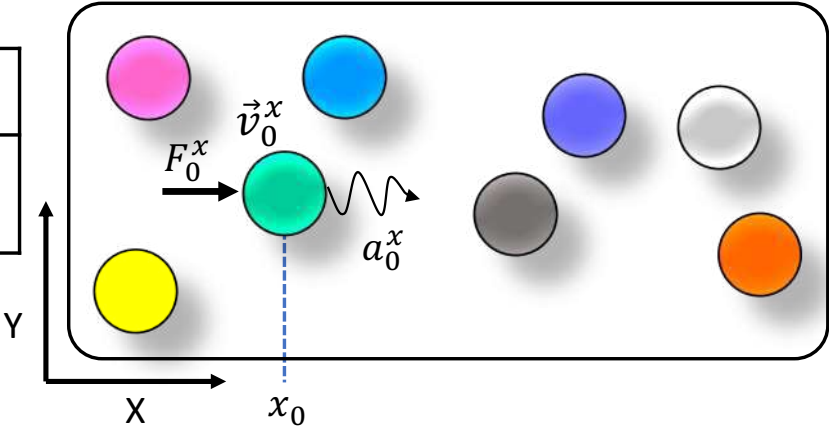
Classical Mechanics

$$F_i = m_i a_i = m_i \frac{d^2 r_i}{dt^2} = - \frac{dV}{dr_i} \quad V = \text{Potential energy of the system}$$

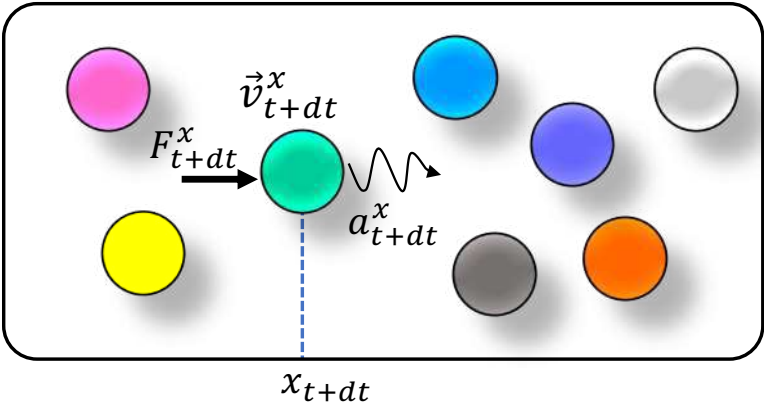
Time	Position (x_0)	Velocity (v_0^x)	Acceleration (a_0^x)
$t = 0$	x_0	v_0^x	$-\frac{1}{m_i} \frac{dV_0}{dx_i}$

Maxwell Boltzmann
distribution

Molecular mechanics
force field

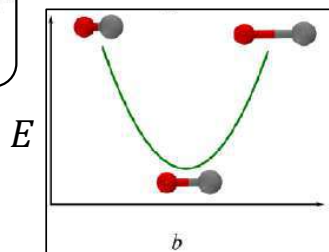
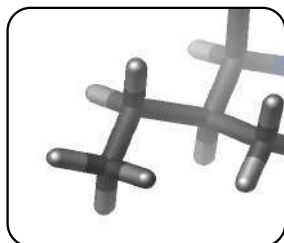


Time	Position (x_{t+dt})	Velocity (v_{t+dt}^x)	Acceleration (a_{t+dt}^x)
$t = t + dt$	$x_0 + v_0^x dt + \frac{1}{2} a_0^x dt^2$	$v_0^x + a_0^x dt$	$-\frac{1}{m_i} \frac{dV_{t+dt}}{dx_i}$



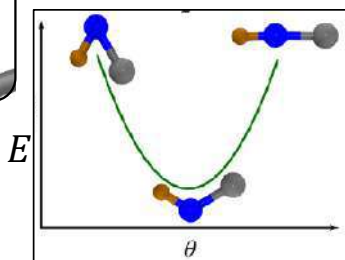
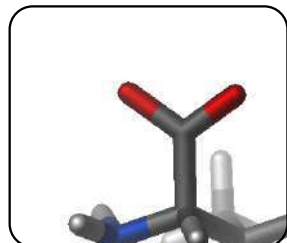
Force field

$$V = E_{bonded} + E_{non-bonded} \quad E_{bonded} = E_{bond-stretching} + E_{angle-bending} + E_{dihedral-rotation}$$



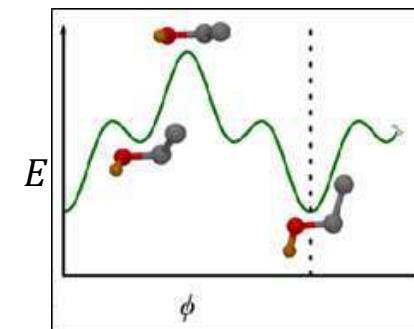
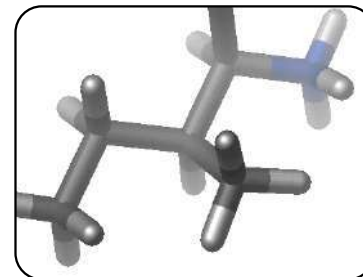
$$E_{bond-stretching} = \sum_{bonds} k_b (b - b_0)^2$$

b = Bond length at any instance
 b_0 = Equilibrium bond length
 k_b = Force constant



$$E_{angle-bending} = \sum_{angles} k_\theta (\theta - \theta_0)^2$$

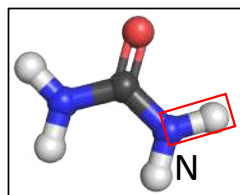
θ = Bond angle at any instance
 θ_0 = Equilibrium bond angle
 k_θ = Force constant



$$E_{dihedral-rotation} = \sum_{dihedral\ angles} k_\varphi (1 - \cos(n\varphi - \delta))$$

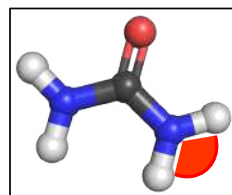
φ = Bond angle at any instance
 n = Dihedral multiplicity (number of local minima)
 k_φ = Dihedral force constant
 δ = Equilibrium dihedral angle

Example: Force field parameters of urea (Charmm36 force field)



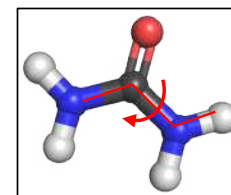
N---H

$b_0 = 1 \text{ \AA}$
 $k_b = 480 \text{ kcal/mol}$



H--N--H

$\theta_0 = 120^\circ$
 $k_\theta = 23 \text{ kcal/mol}$



H--N--C--N

$n = 2, k_\varphi = 1.5 \frac{\text{kcal}}{\text{mol}}$
 $\delta = 180^\circ$

Force field

$$E_{non-bonded} = E_{electrostatic} + E_{van\ der\ Waals}$$

$$E_{electrostatic} = \sum_{i,j} \frac{1}{4\pi\epsilon_0} \frac{q_i q_j}{r_{ij}}$$

$$E_{van\ der\ Waals} = 4\epsilon_{ij} \left[\left(\frac{\sigma_{ij}}{r_{ij}} \right)^{12} - \left(\frac{\sigma_{ij}}{r_{ij}} \right)^6 \right] = \epsilon_{ij} \left[\left(\frac{R_{minij}}{r_{ij}} \right)^{12} - 2 \left(\frac{R_{minij}}{r_{ij}} \right)^6 \right]$$

For n particles, the number of pair interaction terms = $n(n-1)/2$

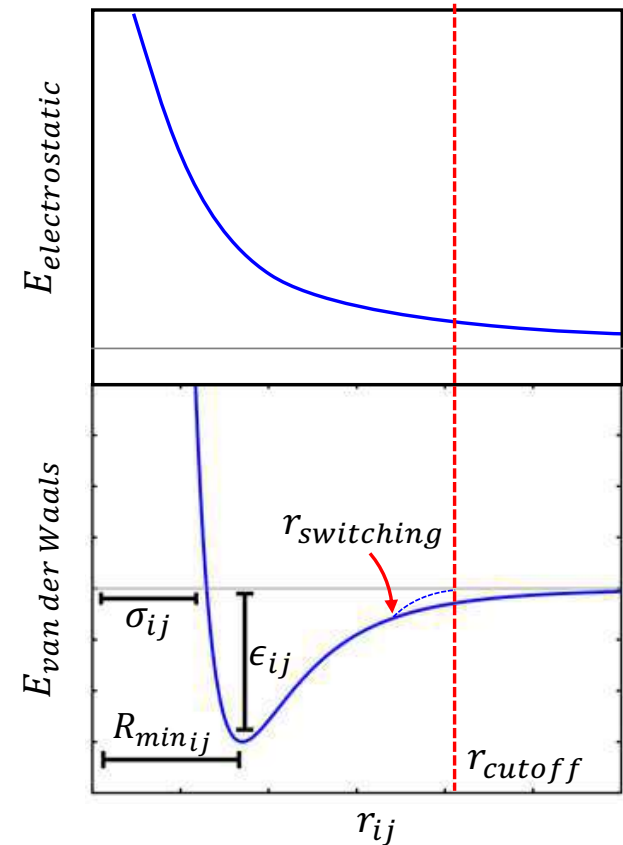
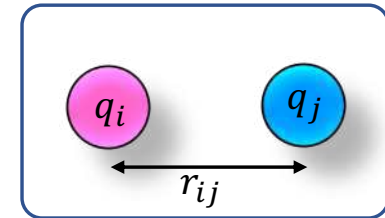
Truncation of interacting partner and cutoff distance

r_{cutoff} Distance at which the short-ranged interactions are turned off

Between $r_{switching}$ and r_{cutoff} the van der Waals interaction is switched to zero

Long-ranged electrostatic interactions are calculated using Particle Mesh Ewald (PME) summation method

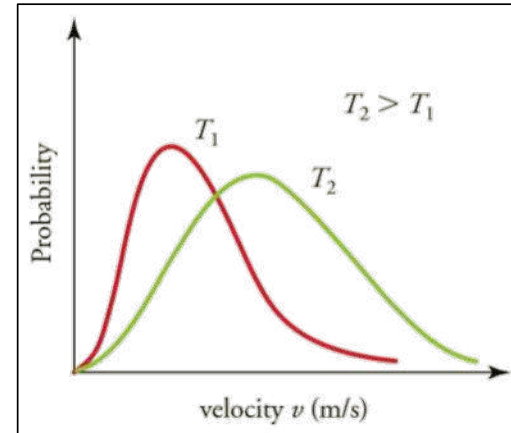
Reduces the computational time significantly



Velocity assignment

- ❑ Velocity is assigned randomly from Maxwell-Boltzmann distribution at a given temperature

$$p(v_i^x) = \left(\frac{m_i}{2\pi k_B T} \right)^{1/2} \exp \left[-\frac{1}{2} \frac{m_i (v_i^x)^2}{k_B T} \right]$$



- ❑ The *overall momentum is also ensured to be zero* $\sum_{i=1}^N m_i v_i = 0$

Integration algorithms

❑ Verlet Algorithm

$$r(t + \delta t) = r(t) + \frac{dr(t)}{dt} \delta t + \frac{1}{2!} \frac{d^2 r(t)}{dt^2} \delta t^2 + \frac{1}{3!} \frac{d^3 r(t)}{dt^3} \delta t^3 + \dots$$

$$r(t - \delta t) = r(t) - \frac{dr(t)}{dt} \delta t + \frac{1}{2!} \frac{d^2 r(t)}{dt^2} \delta t^2 - \frac{1}{3!} \frac{d^3 r(t)}{dt^3} \delta t^3 + \dots$$

Combining,

$$r(t + \delta t) + r(t - \delta t) = 2r(t) + \frac{d^2 r(t)}{dt^2} \delta t^2$$

$$r(t + \delta t) = 2r(t) - r(t - \delta t) + a(t)\delta t^2$$

Faster since velocity need not to be calculated

Lower precision

Integration algorithms

❑ Leap-frog Algorithm

$$v\left(t + \frac{1}{2}\delta t\right) = \frac{r(t + \delta t) - r(t)}{\delta t} \Rightarrow r(t + \delta t) = r(t) + v\left(t + \frac{1}{2}\delta t\right)\delta t$$

$$v\left(t + \frac{1}{2}\delta t\right) = v\left(t - \frac{1}{2}\delta t\right) + a(t)\delta t$$

Velocities at time $t + \frac{1}{2}\delta t$ is used to calculate position at time t .

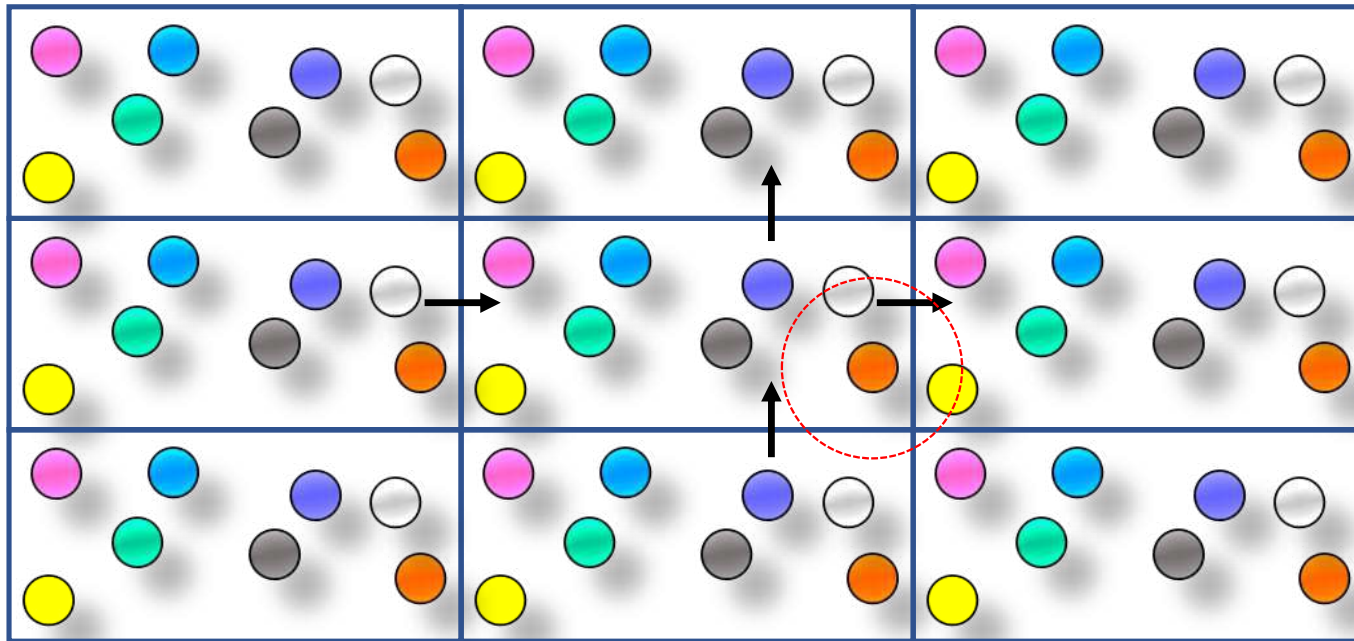
Velocity leaps over position and position leaps over velocity

velocities are calculated explicitly

Position and velocity are not calculated at the same time

Periodic boundary condition

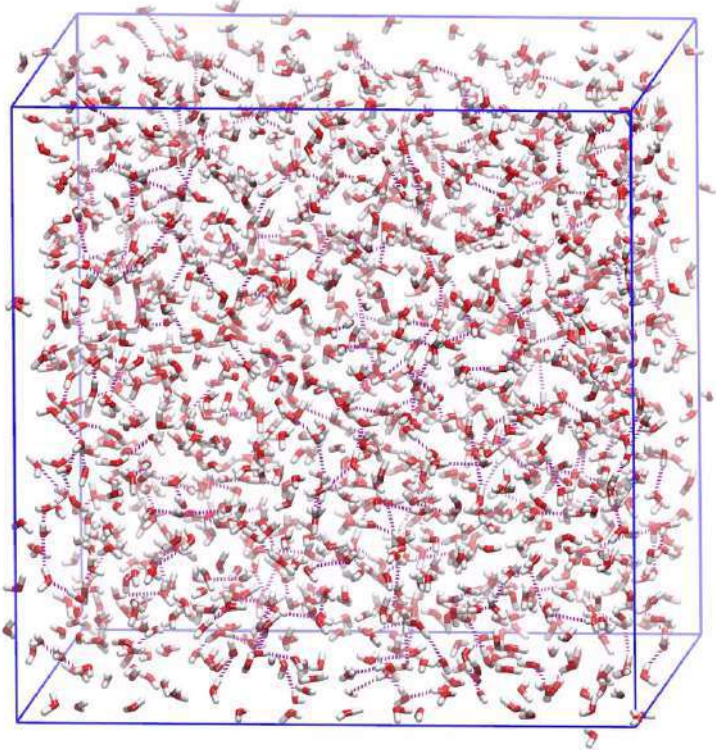
Periodic boundary conditions enable a simulation to be performed using a relatively small number of particles in such a way that the particles experience forces as if they are in a larger volume



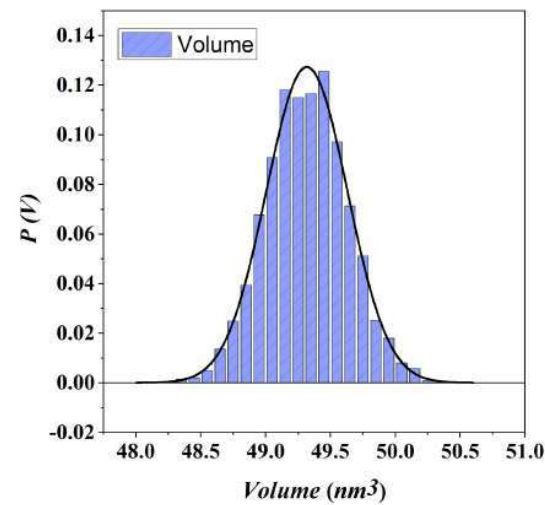
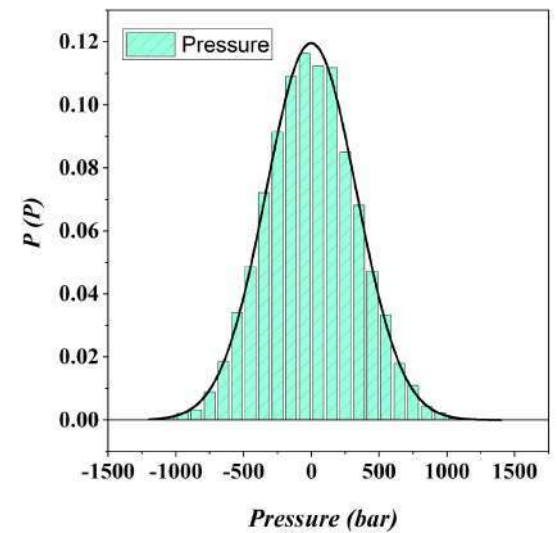
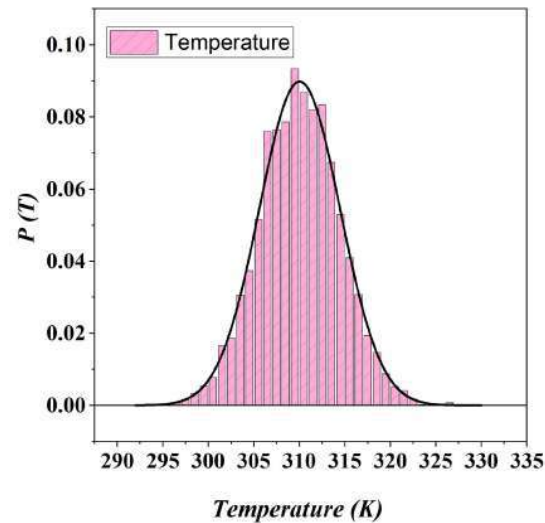
Ergodic hypothesis: Simulation of a water-box

Initial conditions:

- 6136 water molecules
- Cubic box of edge length of 3.67 nm (Volume $\sim 49 \text{ nm}^3$)
- Pressure: 1.01325 bar
- Temperature: 310 K
- Simulation time: 1ns
- Water model: TIP3P



Statistics of temperature, pressure and volume

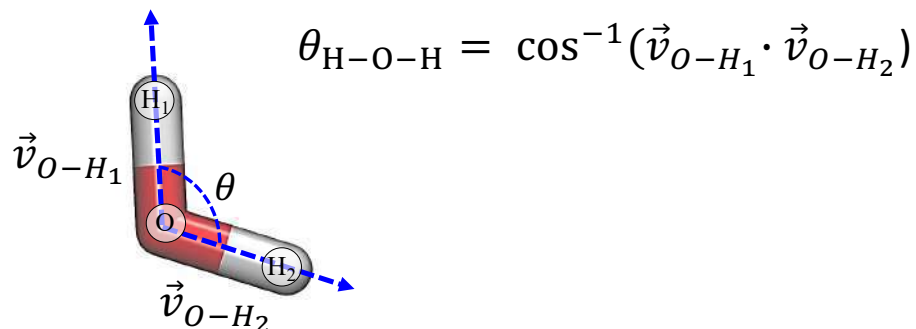


$$\langle T \rangle = 310.032 \text{ K}$$

$$\langle P \rangle = 1.13 \text{ bar}$$

$$\langle V \rangle = 49.32 \text{ nm}^3$$

Ergodic hypothesis: Distributions of H-O-H angle in water



Time →

Ensemble ↓

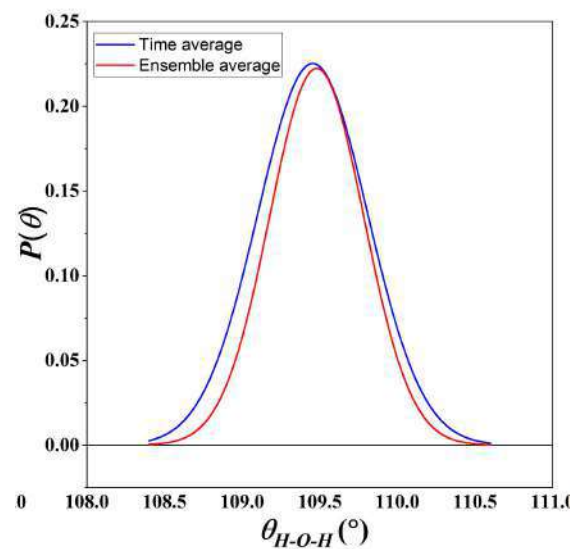
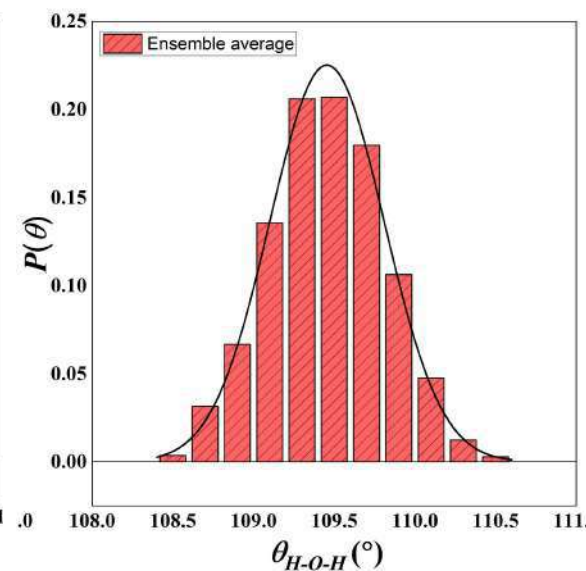
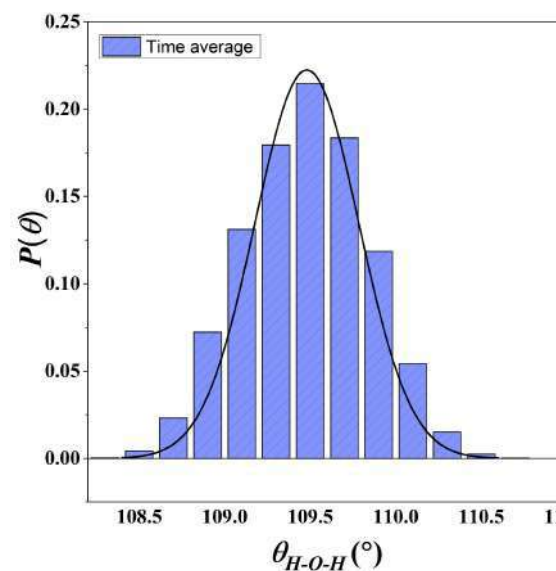
	t_1	t_2	t_3	t_4	t_n
w_1	θ	θ	θ	θ	θ
w_2	θ	θ	θ	θ	θ
w_3	θ	θ	θ	θ	θ
⋮	⋮	⋮	⋮	⋮	⋮	⋮
w_N	θ	θ	θ	θ	θ

Ensemble average

$$\langle \theta \rangle_N = \lim_{N \rightarrow \infty} \frac{1}{N} \sum_{i=1}^N \theta(i)$$

Time average

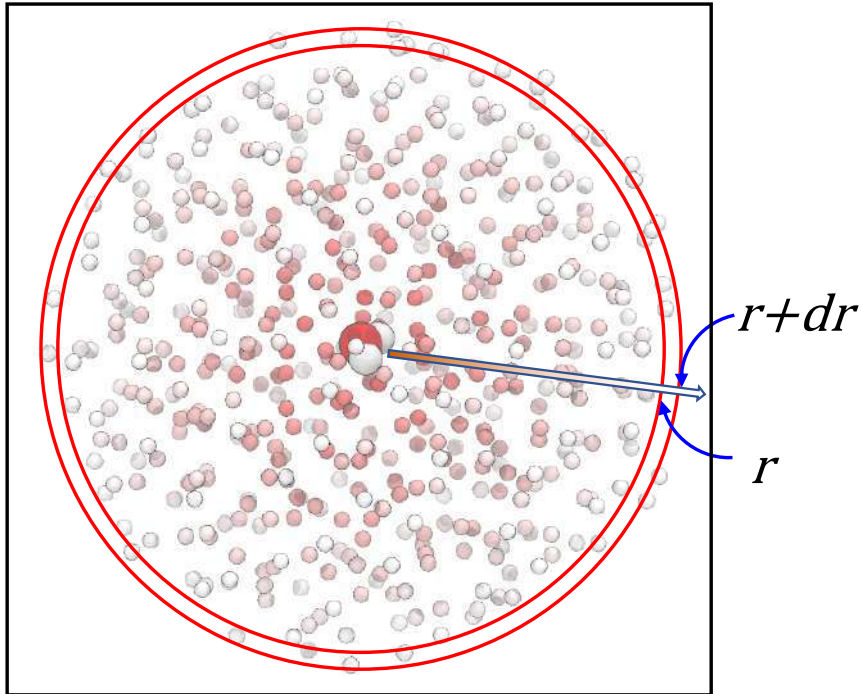
$$\langle \theta \rangle_t = \lim_{t \rightarrow \infty} \frac{1}{t_n} \int_0^{t_n} \theta(t) dt$$



$$\langle \theta \rangle_t = 109.47^\circ$$

$$\langle \theta \rangle_N = 109.45^\circ$$

Calculation of radial distribution function ($g(r)$):



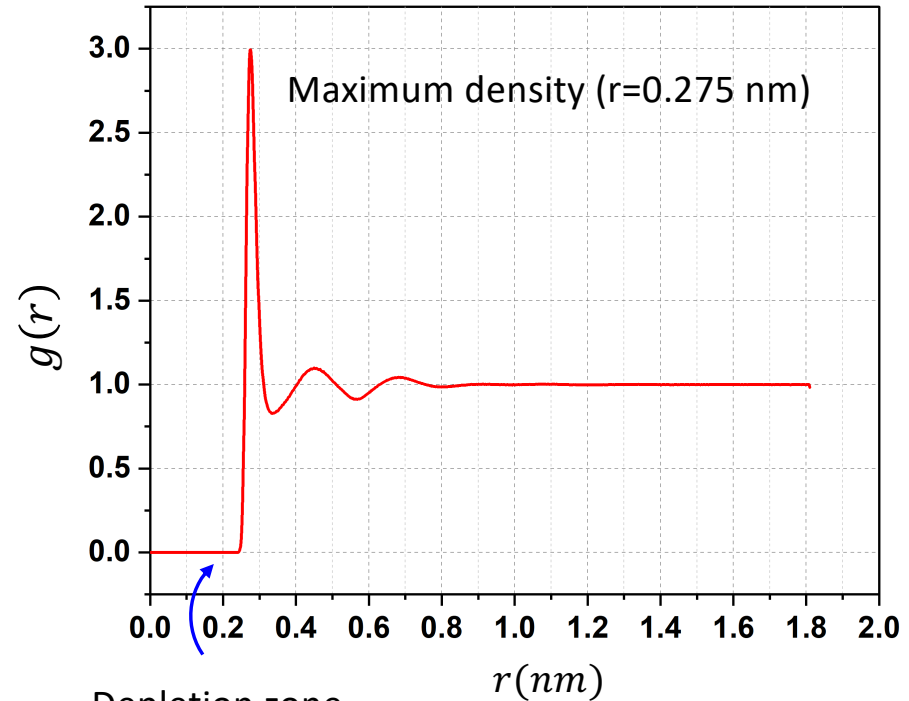
$$4\pi r^2 g(r) \rho_{r_{Bulk}} dr = dN_r$$

$$4\pi \rho_{r_{Bulk}} \int g(r) r^2 dr = \int dN_r = N_r$$

Coordination number

$$g(r) = \frac{\rho_r}{\rho_{r_{Bulk}}} \quad \rho_r = \frac{dN_r}{4\pi r^2 dr} = \frac{N_{r+dr} - N_r}{4\pi r^2 dr}$$

N_r = No. of water oxygen atoms in the sphere of radius r surrounding a central water oxygen atom



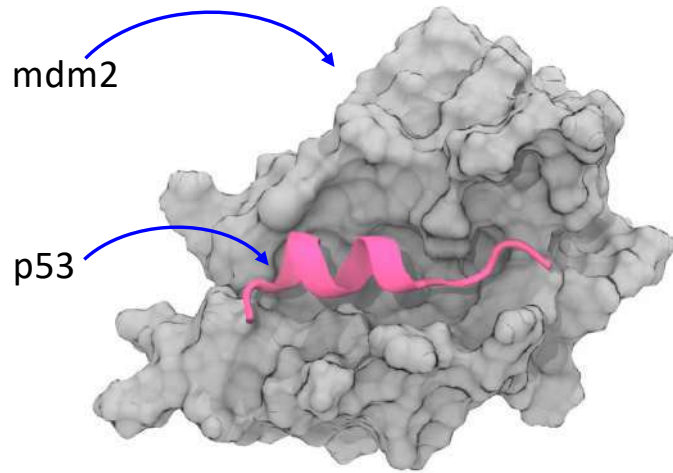
Depletion zone

$[> 2r_{O-H} (r_{O-H} = 0.095 \text{ nm})]$

Simulation of protein-peptide complex

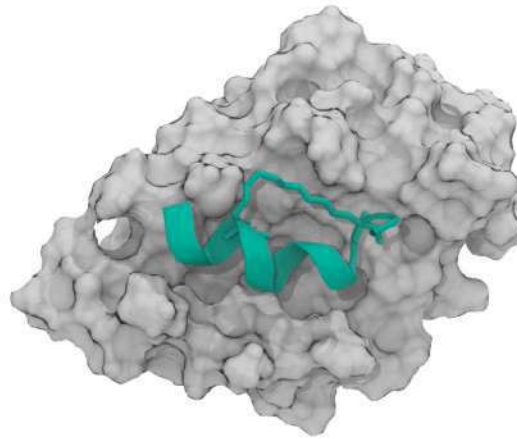
- ❑ p53 regulates the DNA repairing or apoptosis (removal of diseased cell) in response to different stresses
- ❑ mdm2 is overexpressed in cancerous cells to bind to p53 and inhibit its function to ensure tumor progression
- ❑ Cross-stitched peptides are designed to replace p53 and bind to mdm2
- ❑ For that, these must have higher binding affinity with mdm2 than the natural binding partner.

mdm2 + p53 (natural binding partner)



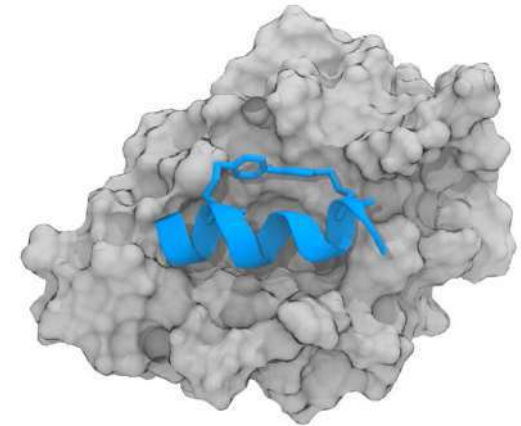
Unit: kcal/mol	ΔH	-62.34
	$T\Delta S$	- 53.88
	ΔG	-8.46 ± 0.69

mdm2 + designed peptide inhibitor

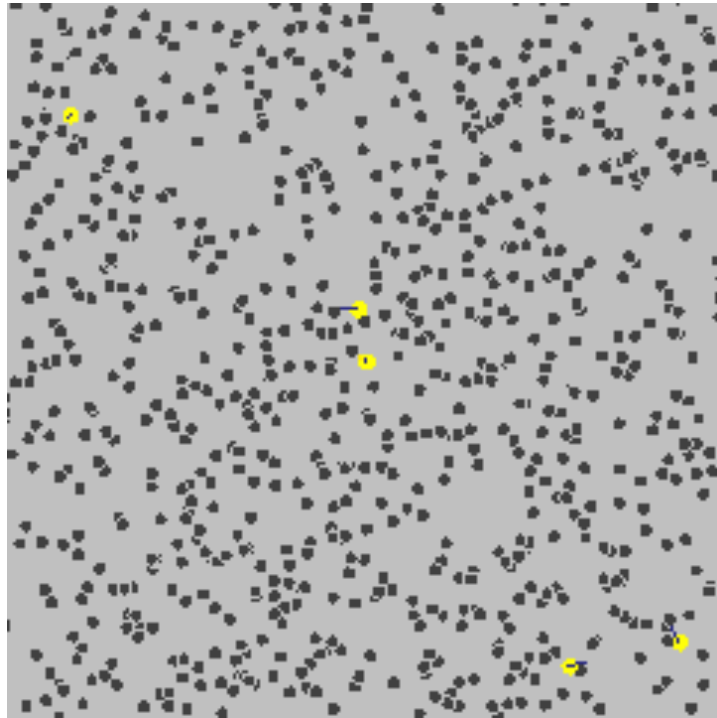
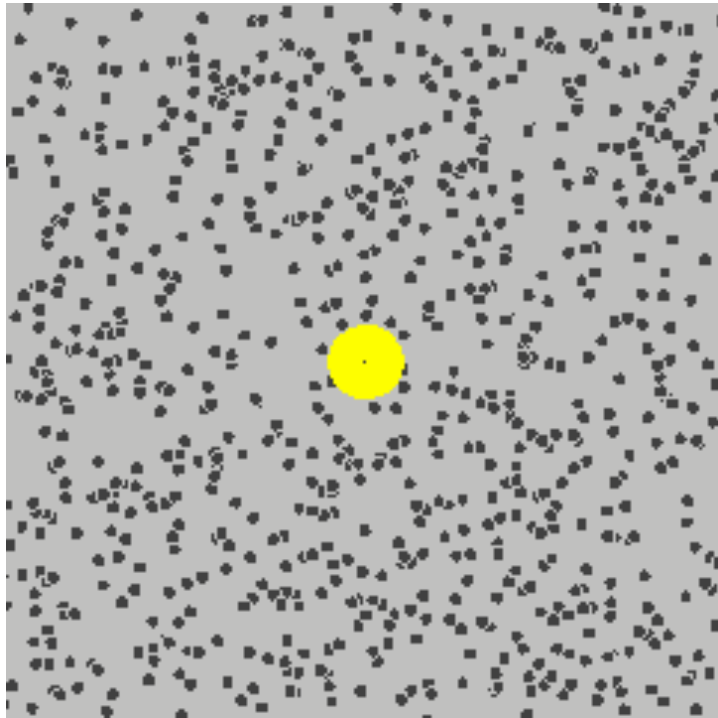


ΔH	-56.53
$T\Delta S$	- 26.41
ΔG	-30.12 ± 2.97

mdm2 + designed peptide inhibitor



ΔH	-54.65
$T\Delta S$	- 15.95
ΔG	-38.70 ± 2.46



Einstein
1905

$n \rightarrow$ number of particles

$dn \rightarrow$ No. of particles

Which experience a displacement
between Δ and $\Delta + d\Delta$ in
time interval $\tau \rightarrow$

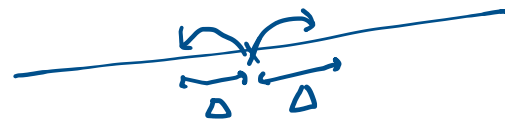
is small compared
to observation time
but larger than
any correlation time

$x(t)$
 $x(t+\tau)$ } they are
uncorrelated

$$\frac{dn}{n} = \phi(\Delta) d\Delta$$

\leadsto probability of jump

For simplicity take 1D motion



$$(*) \int_{-\infty}^{+\infty} \phi(\Delta) d\Delta = 1$$

$$(*) \phi(-\Delta) = \phi(\Delta)$$

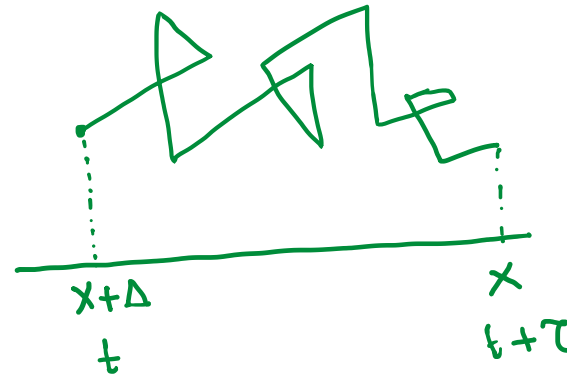
"Even"
function

$$n = f(x, t)$$

→ No. of particles
per unit vol^m at
time "t"

$$f(x, t+\tau) = \int_{-\infty}^{+\infty} f(x+\Delta, t) \phi(\Delta) d\Delta$$

↓
Taylor expansion



$$f(x, t) + \tau \frac{\partial f(x, t)}{\partial t} = \int_{-\infty}^{+\infty} \left\{ f(x, t) + \Delta \frac{\partial f}{\partial x} + \frac{\Delta^2}{2!} \frac{\partial^2 f}{\partial x^2} + \dots \right\} \phi(\Delta) d\Delta$$

$$= \int_{-\infty}^{+\infty} f(x, t) \phi(\Delta) d\Delta + \int_{-\infty}^{+\infty} \left(\frac{\partial f}{\partial x} \right) \Delta \phi(\Delta) d\Delta + \int_{-\infty}^{+\infty} \left(\frac{\partial^2 f}{\partial x^2} \right) \frac{\Delta^2}{2!} \phi(\Delta) d\Delta$$

$$\cancel{f(x, t)} + \tau \frac{\partial f}{\partial t} = \cancel{f(x, t)} + \int_{-\infty}^{+\infty} \frac{\Delta^2}{2!} \phi(\Delta) d\Delta \left(\frac{\partial^2 f}{\partial x^2} \right)$$

$$\frac{\partial f(x,t)}{\partial t} = \int_{-\infty}^{\infty} \frac{\Delta^2}{2! 2} \phi(\Delta) d\Delta \frac{\partial^2 f}{\partial x^2}$$

D (Diffusion coefficient)

$$\frac{\partial f(x,t)}{\partial t} = D \frac{\partial^2 f}{\partial x^2} \leftarrow \text{"Einstein's Diffusion equation"}$$

$$i\hbar \frac{\partial \psi}{\partial t} = H\psi = \left[-\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} \right] \psi \Rightarrow$$

→ Schrödinger eqⁿ for a free particle

"Initial value problem"

$$f(x,0) = n\delta(x)$$

→ Dirac Delta function"

$$\frac{\partial \psi}{\partial t} = +i\frac{\hbar}{2m} \frac{\partial^2 \psi}{\partial x^2}$$

"Boundary value problem"

→ Solving it using separation in variables technique

$$f(x, 0) = n \delta(x)$$

$$\delta(x) = 0 \quad x \neq 0$$

$$= \infty \quad x = 0$$

$$f(x_0) = \int_{-\infty}^{+\infty} f(x) \delta(x - x_0) dx$$

Fourier representation of Delta function

$$\delta(x) = \frac{1}{2\pi} \int_{-\infty}^{+\infty} e^{ikx} dk$$

$$f(x, t) = \frac{1}{2\pi} \int_{-\infty}^{+\infty} \tilde{f}(k, t) e^{ikx} dk$$

representing this as the Fourier t-transform of $\tilde{f}(k, t)$

$$\frac{\partial f(x, t)}{\partial t} = \frac{1}{2\pi} \int_{-\infty}^{+\infty} \frac{\partial \tilde{f}(k, t)}{\partial t} e^{ikx} dk \quad \dots \dots \textcircled{A}$$

$$\frac{\partial f}{\partial t} = 0 \quad \frac{\partial^r f}{\partial x^r}$$

$$D \frac{\partial^2 f}{\partial x^2} = \frac{1}{2\pi} D \int_{-\infty}^{+\infty} \tilde{f}(k, t) (ik)^2 e^{ikx} dk = \frac{D}{2\pi} \int_{-\infty}^{+\infty} \tilde{f}(k, t) (-k^2) e^{ikx} dk \quad \dots \textcircled{B}$$

$$\frac{\partial \tilde{f}(k, t)}{\partial t} = -D k^2 \tilde{f}(k, t)$$

$$\frac{\partial \tilde{f}(k,t)}{\partial t} = -Dk^2 \tilde{f}(k,t)$$

$$\tilde{f}(k,t) = \tilde{f}(k,0) e^{-Dk^2 t}$$

$$\tilde{f}(k,t) = n e^{-Dk^2 t}$$

$$f(x,t) = \frac{1}{2\pi} \int_{-\infty}^{+\infty} \tilde{f}(k,t) e^{ikx} dk$$

$$f(x,t) = \frac{1}{2\pi} \int_{-\infty}^{+\infty} n e^{-Dk^2 t} e^{ikx} dk = \frac{n}{2\pi} \int_{-\infty}^{+\infty} e^{-Dk^2 t} e^{ikx} dk$$

$$f(x,0) = n \delta(x) \quad (\text{initial condition})$$

$$f(x,0) = \frac{1}{2\pi} \int_{-\infty}^{+\infty} e^{ikx} \tilde{f}(k,0) dk$$

$$f(x,0) = n \frac{1}{2\pi} \int_{-\infty}^{+\infty} e^{ikx} dk$$

$$\tilde{f}(k,0) = n$$

$$f(x, t) = \frac{n}{2\pi} \int_{-\infty}^{+\infty} e^{-Dk^2 t} e^{ikx} dk = \frac{n}{2\pi} \int_{-\infty}^{+\infty} e^{-\left\{ (k\sqrt{Dt})^2 - 2k\sqrt{Dt} \frac{1}{2} \frac{ix}{\sqrt{Dt}} + \frac{x^2 i^2}{4Dt} \right\}} e^{-x^2/4Dt} dk$$

$$= \frac{n}{2\pi} e^{-x^2/4Dt} \int_{-\infty}^{+\infty} e^{-\left(k\sqrt{Dt} - \frac{ix}{2\sqrt{Dt}} \right)^2} dk$$

$$= \frac{n}{2\pi} e^{-x^2/4Dt} \int_{-\infty}^{+\infty} e^{-y^2} \frac{dy}{\sqrt{Dt}}$$

$$= \frac{n}{2\pi} e^{-x^2/4Dt} \frac{1}{\sqrt{Dt}} \sqrt{\pi}$$

$$f(x, t) = \frac{n}{\sqrt{4\pi Dt}} e^{-x^2/4Dt} \rightarrow \text{"Probability density"}$$

$n=1$

$$f(x, t) = \frac{1}{\sqrt{4\pi Dt}} e^{-x^2/4Dt}$$

$$\text{Let, } y = k\sqrt{Dt} - \frac{ix}{2\sqrt{Dt}}$$

$$dy = dk\sqrt{Dt}$$

$$\left[\int_{-\infty}^{+\infty} e^{-\alpha x^2} dx = \left(\pi/\alpha \right)^{1/2} \right]$$

$$f(x, t) = \frac{1}{\sqrt{4\pi Dt}} e^{-x^2/4Dt}$$

$\langle x \rangle = 0$, all the odd moments are zero
and all the even moments are non-zero

$$\langle x^2 \rangle = \int_{-\infty}^{+\infty} x^2 \frac{1}{\sqrt{4\pi Dt}} e^{-x^2/4Dt} dx$$

$$= \frac{1}{\sqrt{4\pi Dt}} \int_{-\infty}^{+\infty} x^2 e^{-x^2/4Dt} dx$$

$$= \frac{1}{\sqrt{4\pi Dt}} \frac{1}{2 \cdot \frac{1}{4Dt}} \left(\pi / \frac{1}{4Dt} \right)^{\frac{1}{2}}$$

$$= \frac{1}{\sqrt{4\pi Dt}} 2Dt \sqrt{4\pi Dt}$$

$$\alpha \equiv \frac{1}{4Dt}$$

$$I = \int_{-\infty}^{+\infty} e^{-\alpha x^2} dx = \left(\pi / \alpha \right)^{\frac{1}{2}}$$

$$\frac{2I}{2\alpha} = - \int_{-\infty}^{+\infty} e^{-\alpha x^2} x^2 dx$$

$$= -\frac{1}{2\alpha} \left(\pi / \alpha \right)^{\frac{1}{2}}$$

$$\int_{-\infty}^{+\infty} x^2 e^{-\alpha x^2} dx = \frac{1}{2\alpha} \left(\pi / \alpha \right)^{\frac{1}{2}}$$

$$\langle x^2 \rangle = 2Dt$$

Diffusive

$$\lambda_x = \sqrt{\langle x^2 \rangle} \sim \sqrt{2Dt} t^{\frac{1}{2}}$$

Mean displacement

For Newtonian particle moving with constant velocity $\lambda_x \sim vt$

$$\lambda_x \sim t^2 \text{ (Ballistic)}$$

Mean square displacement

Diffusion & Mobility

(1-Dimension)

f = force of gravity

Viscous drag

$$= 6\pi\eta r v_0 \leftarrow \text{"Stokes law"}$$

When this external force ' f '
balances the viscous drag, then
the particle moves with the
Constant velocity v_0

If ν is the number of particles
per unit volume, then $\nu v_0 \rightarrow$ no. of
particles passing through a unit area per unit time

○ Spherical particle of radius r
is moving under the force
of gravity in a liquid.

↓
Coefficient of
viscosity " η "

$$f = 6\pi\eta r v_0$$

$$v_0 = \frac{f}{6\pi\eta r}$$

$$\Rightarrow \frac{\nu f}{6\pi\eta r}$$

$-D \frac{\partial v}{\partial x} \rightarrow$ Fick's law (first law)

$D \rightarrow$ "Diffusion coefficient"

$f \rightarrow$ force due to gravity

\hookrightarrow "Diffusion flux"

Boltzmann distribution of v (Density)

\rightarrow no. of particles per unit area per unit time

$$-D \frac{\partial v}{\partial x} = \frac{vf}{6\pi\eta r}$$

$$-\int_{v_0}^v \frac{dv'}{v'} = \frac{f}{6\pi\eta r D} \int_{x_0}^x dx'$$

$$\Rightarrow \ln(v/v_0) = -\frac{f}{6\pi\eta r D} (x - x_0)$$

$$\Rightarrow v = v_0 \exp\left[-\frac{f(x-x_0)}{6\pi\eta r D}\right] \dots \textcircled{\text{II}}$$

$$D = \frac{RT}{N_A} \cdot B$$

Comparing $\textcircled{\text{I}}$ & $\textcircled{\text{II}}$

$$\frac{N_A}{RT} = \frac{1}{6\pi\eta r D}$$

$$B = \frac{1}{6\pi\eta r}$$

\hookrightarrow Mobility

$$D = k_B T B$$

$$D = \frac{k_B T}{6\pi\eta r}$$

\downarrow
Stokes-Einstein relation

Determination of Avagadro's number:

$$D = \frac{RT}{N_A} \frac{1}{6\pi\eta r}$$

$$\lambda_x = \sqrt{2t} \left(\frac{RT}{N_A} \right)^{\frac{1}{2}} \left(\frac{1}{6\pi\eta r} \right)^{\frac{1}{2}}$$

↪ to find N_A

Experimentally

can be measured

$\eta \rightarrow$ known for liquids

$$\lambda_x = \sqrt{2Dt}$$

$$\begin{aligned} \langle x^2 \rangle &= 2Dt \\ \sqrt{\langle x^2 \rangle} &= \sqrt{2Dt} \\ &\downarrow \\ \lambda_x \end{aligned}$$

At the single particle level



$$m \frac{dv}{dt} = -6\pi\eta r v \quad \begin{array}{l} \swarrow \text{Force} \\ \text{(Newton's equation)} \end{array}$$

$$v = v_0 \exp\left(-\frac{6\pi\eta r}{m} t\right)$$

$$\ln(v/v_0) = -\frac{6\pi\eta r}{m} t$$

$$t_{1/10} \equiv v/v_0 = \frac{1}{10}$$

$$\ln 10 = \frac{6\pi\eta r}{m} t_{1/10}$$

$$\Rightarrow t_{1/10} = 3.3 \times 10^{-7} \text{ sec}$$

$$\sim 10^{-7} \text{ sec}$$

Platinum particles

$$r = 2.5 \times 10^{-6} \text{ cm}$$

$$\eta = 0.01 \text{ poise (water)}$$

$$m = 2.5 \times 10^{-15} \text{ g}$$

Within very short time the particle nearly completely loses its original velocity \rightarrow In reality it does not happen!!

$$m \frac{dv}{dt} = -6\pi\eta r v + F(t)$$

random Kicks
 \hookrightarrow random function
 of time \equiv Noise

1908

Langevin eqn

"Paul Langevin"

$$\langle F(t) \rangle = 0$$

$$\langle F(t) F(t') \rangle = 2C'\delta(t-t')$$

$$m \ddot{x} = -6\pi\eta r \dot{x} + F(t)$$

Multiplying both sides by x

$$m x \ddot{x} = -6\pi\eta r x \dot{x} + x F(t)$$

$$\Rightarrow x \ddot{x} = -\frac{6\pi\eta r}{m} x \dot{x} + \frac{1}{m} x F(t)$$

$$\dot{x} = v$$

$$\ddot{x} = \dot{v} = \frac{dv}{dt}$$

$$x \ddot{x} = -\Gamma \dot{x} + \frac{1}{m} x F(t)$$

$$\frac{1}{2} \ddot{x}^2 - (\dot{x})^{\ddot{}} = -\Gamma \dot{x} + \frac{1}{m} x F(t)$$

$$\Rightarrow \ddot{x}^2 - 2(\dot{x})^{\ddot{}} = -2\Gamma \dot{x} + \frac{2}{m} x F(t)$$

$$\Rightarrow \ddot{x}^2 - 2(\dot{x})^{\ddot{}} = -\Gamma \dot{x} + \frac{2}{m} x F(t) \Rightarrow \frac{d^2}{dt^2} \langle x^2 \rangle - 2 \left(\frac{dx}{dt} \right)^2 = \dots$$

Taking the averages on both sides

$$\frac{d^2}{dt^2} \langle x^2 \rangle - 2 \langle \dot{x}^2 \rangle = -\Gamma \frac{d}{dt} \langle x^2 \rangle + \frac{2}{m} \langle x F(t) \rangle$$

$$\downarrow$$

$$-2 \langle v^2 \rangle$$

$$\rightarrow \frac{2k_B T}{m}$$

Zero (position and the random force are uncorrelated)

$$\dot{x}^2 = \frac{d}{dt} (x^2)$$

$$= 2x\dot{x}$$

$$\ddot{x}^2 = \frac{d^2}{dt^2} (x^2)$$

$$= \frac{d}{dt} (2x\dot{x})$$

$$\ddot{x}^2 = 2(\dot{x})^{\ddot{}} + 2x\ddot{x}$$

$$\Rightarrow x \ddot{x} = \frac{1}{2} \ddot{x}^2 - (\dot{x})^{\ddot{}}$$

$$\frac{d^2}{dt^2} \langle x^2 \rangle + \Gamma \frac{d}{dt} \langle x^2 \rangle - \frac{2k_B T}{m} = 0$$

$$\ddot{y} + \Gamma \dot{y} - c = 0$$

$$\text{Let, } \frac{d}{dt} \langle x^2 \rangle = y$$

$$\frac{2k_B T}{m} = c$$

$$\text{Let, } \Gamma y - c = y' \text{ (not derivative)}$$

$$\dot{y} = -y' \Rightarrow \ddot{y} = -\dot{y}'$$

$$\dot{y}' = \Gamma \dot{y} + 0$$

$$\dot{y}' = \Gamma \dot{y}$$

$$\dot{y}' = -\Gamma y' \quad (\text{as } \dot{y} = -y')$$

$$\Rightarrow \frac{dy'}{dt} = -\Gamma y'$$

$$\Rightarrow \int_{y'(t=0)}^{y'} \frac{dy'}{y'} = -\Gamma \int_0^t dt'$$

$$\Rightarrow$$

$$\dot{x} = -\Gamma x$$

$$y' = A e^{-\Gamma t}$$

$$A = y'(t=0)$$

$$\Gamma y - c = y' = A e^{-\Gamma t}$$

$$y = \frac{d}{dt} \langle x^2 \rangle$$

$$\Rightarrow y = \frac{c}{\Gamma} + \frac{A}{\Gamma} e^{-\Gamma t}$$

$$\frac{d}{dt} \langle x^2 \rangle = \frac{c}{\Gamma} + \frac{A}{\Gamma} e^{-\Gamma t}$$

Integrating

$$\langle x^2 \rangle = \frac{c}{\Gamma} t + \frac{A}{\Gamma} \int_0^t e^{-\Gamma t'} dt'$$

$$\langle x^2 \rangle = \frac{c}{\Gamma} t + \frac{A}{\Gamma} \frac{e^{-\Gamma t'}}{-\Gamma} \Big|_0^t = \frac{c}{\Gamma} t + \frac{A}{\Gamma^2} (1 - e^{-\Gamma t})$$

$$\langle x^2 \rangle = \frac{c}{\Gamma} t + \frac{A}{\Gamma^2} (1 - e^{-\Gamma t})$$

$$y'(0) = A$$

$$\Gamma y(0) - c = A$$

$$y(0) = \left(\frac{d}{dt} \langle x^2 \rangle \right)_{t=0}$$

$$= 0$$

$$A = -c$$

$$\langle x^2 \rangle = \frac{c}{\Gamma} t + \frac{(1-c)}{\Gamma^2} (1 - e^{-\Gamma t})$$

$$c = \frac{2k_B T}{m}$$

$$\langle x^2 \rangle = \frac{2k_B T}{m\Gamma} t - \frac{2k_B T}{m\Gamma^2} (1 - e^{-\Gamma t})$$

→ general expression
for the mean

square displacement

of a Brownian

particle suspended

in a fluid

(no external force)

When $t \rightarrow \infty$ (long time limit)

$$\langle x^2 \rangle = \frac{2k_B T}{m\Gamma} t - \frac{2k_B T}{m\Gamma^2} (1 - 0)$$

$$\langle x^2 \rangle \simeq \frac{2k_B T}{m\Gamma} t = \frac{2k_B T}{m \frac{6\pi\eta r}{m}} t = 2 \left(\frac{k_B T}{6\pi\eta r} \right) t$$

$$\langle x^2 \rangle = 2Dt$$

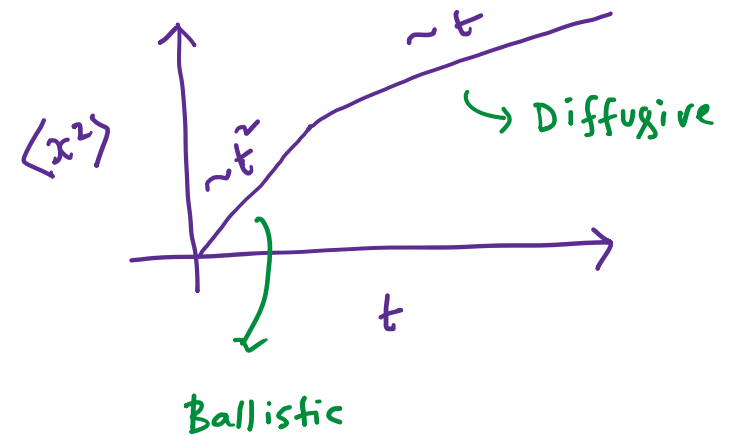
$$\langle x^2 \rangle = \frac{2k_B T}{m\Gamma} t - \frac{2k_B T}{m\Gamma^2} \left(1 - \left(1 - \Gamma t + \frac{\Gamma^2 t^2}{2!} + \dots \right) \right)$$

$t \rightarrow 0$ (short time limit)

$$\langle x^2 \rangle = \frac{2k_B T}{m \Gamma} t - \frac{2k_B T}{m \Gamma^2} \left(\cancel{\Gamma} - \cancel{\Gamma} + \Gamma t - \frac{\Gamma^2 t^2}{2} \dots \right)$$

$$= \frac{2k_B T}{\cancel{m \Gamma}} t - \frac{2k_B T}{\cancel{m \Gamma}} t + \frac{k_B T}{m} t^2 \dots$$

$$\langle x^2 \rangle = \frac{k_B T}{m} t^2 \rightarrow \text{Ballistic} \rightarrow \text{inertia driven}$$



A Langevin description without inertia (large viscosity limit).

$$6\pi\eta a \dot{x} = F(t) \rightarrow \langle x^2 \rangle \sim t$$

"no ballistic"

Relation between the random and the viscous force: The fluctuation-dissipation theorem (FDT)

Langevin equation

$$\frac{dv}{dt} = -\frac{6\pi\eta r v}{m} + \frac{F(t)}{m}$$

$$\Gamma = \frac{6\pi\eta r}{m}$$

$$\frac{dv}{dt} = -\Gamma v + \frac{F(t)}{m}$$

Multiplying both sides by v

$$v \frac{dv}{dt} = -\Gamma v^2 + \frac{1}{m} v F(t)$$

$$\downarrow \frac{d}{dt} v^2 = -2\Gamma v^2 + \frac{2}{m} v F(t)$$

$$\frac{d}{dt} v^2 = 2v \frac{dv}{dt}$$

Taking ensemble average

$$\frac{d}{dt} \langle v^2 \rangle = -2\Gamma \langle v^2 \rangle + \frac{2}{m} \underbrace{\langle v F(t) \rangle}_{\text{How to evaluate ??}}$$

Now $\int_{t-\Delta t}^t \dot{v}(t') dt' = v(t) - v(t-\Delta t) \Rightarrow v(t) = v(t-\Delta t) + \int_{t-\Delta t}^t \dot{v}(t') dt'$

Multiplying both sides by $F(t)$ and taking average

$$\langle F(t)v(t) \rangle = \underbrace{\langle F(t)v(t-\Delta t) \rangle}_{=0} + \int_{t-\Delta t}^t \langle F(t)\dot{v}(t') \rangle dt'$$

$= 0$ [as the velocity at earlier instant, $v(t-\Delta t)$ $t-\Delta t < t$, has no dependence on the future fluctuating force]

$$\langle F(t)v(t) \rangle = \int_{t-\Delta t}^t \langle F(t)\dot{v}(t') \rangle dt'$$

$$\langle F(t) v(t) \rangle = \int_{t-\Delta t}^t \langle \dot{v}(t') F(t) \rangle dt'$$

$$= \int_{t-\Delta t}^t \left\langle \left(-\Gamma v(t') + \frac{F(t')}{m} \right) F(t) \right\rangle dt'$$

$$= -\Gamma \int_{t-\Delta t}^t \langle v(t') F(t) \rangle dt' + \frac{1}{m} \int_{t-\Delta t}^t \langle F(t') F(t) \rangle dt'$$

?

Since t' is the earlier time ($< t$)

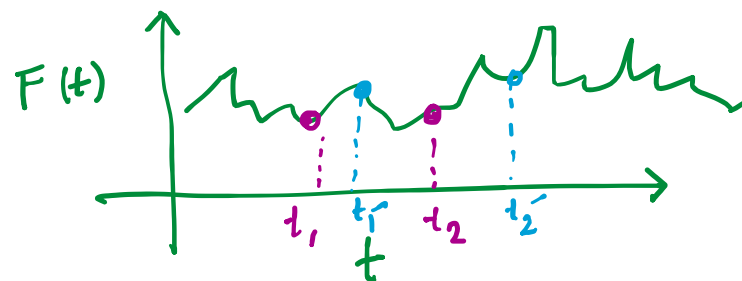
and the fluctuating force is
at a later time t has no dependence
on velocity at earlier time t' , except
at $t=t'$, for which the integral is 0.

$$\langle \psi(t) \bar{F}(t) \rangle = \frac{1}{m} \int_{t-\Delta t}^t \langle F(t) F(t') \rangle dt'$$

We assume that $\bar{F}(t)$ is stationary in time.
This means the value of the integral depends on the difference $t-t'$ but not on t and t' individually.

$$\begin{aligned} \langle \psi(t) \bar{F}(t) \rangle &= \frac{1}{2m} \int_{t-\Delta t}^{t+\Delta t} \langle F(t) F(t') \rangle dt' \\ &= \frac{1}{2m} \int_{t-\Delta t}^{t+\Delta t} \langle F(t) F(t+s) \rangle ds \end{aligned}$$

$$= \frac{1}{2m} \int_{-\infty}^{+\infty} \langle \bar{F}(0) F(s) \rangle ds$$



$$\Delta t = t_2 - t_1 = t_2' - t_1'$$

(time translation invariance)

① Since $\Delta t \gg (t' - t = s)$
We put $\Delta t \sim \infty$

② Without any loss of generality
We put $t = 0$

$$; t = t' + s$$

We had

$$\frac{d}{dt} \langle v^2 \rangle = -2\Gamma \langle v^2 \rangle + \frac{2}{m} \langle v(t) F(t) \rangle$$

$$\langle v^2 \rangle = \frac{k_B T}{m}$$

$$0 = -2\Gamma \langle v^2 \rangle + \frac{2}{m} \frac{1}{2m} \int_{-\infty}^{+\infty} \langle F(0) F(s) \rangle ds$$

$$= -2\Gamma \frac{k_B T}{m}$$

$$\Gamma = \frac{1}{2k_B T m} \int_{-\infty}^{+\infty} \langle F(0) F(s) \rangle ds$$

Dissipation

FDT

(Fluctuation Dissipation theorem)

Correlation betⁿ
the random
forces

Recall

$$\langle F \rangle \Gamma = 2\zeta L$$

Balance betⁿ Dissipation & fluctuation

$$\langle F(t)F(s) \rangle = 2C' \delta(s)$$

$$\Gamma = \frac{1}{2k_B T m} \int_{-\infty}^{+\infty} 2C' \delta(s) ds = \frac{C'}{k_B T m}$$

$$\int \delta(s) ds = 1$$

$$C' = \Gamma k_B T m = \frac{6\pi\eta r}{m} k_B T m$$

$$C' = 6\pi\eta r k_B T \rightarrow \text{Special form of FDT}$$

$$\langle F(t)F(t') \rangle = 2\gamma k_B T \delta(t-t') ; \gamma = 6\pi\eta r$$

friction coefficient

Brownian Motion in velocity space:

Recall, in the co-ordinate space

$$f(x, t+\tau) = \int_{-\infty}^{+\infty} f(x+\Delta, t) \phi(\Delta) d\Delta$$

$\Delta \rightarrow$ jump in coordinate space

$$\hookrightarrow \frac{\partial f}{\partial t} = D \frac{\partial^2 f}{\partial x^2}$$

"Einstein's Diffusion equation"

Velocity space

$$f(v, t+\tau) = \int_{-\infty}^{+\infty} f(v-\Delta, t) \underbrace{\phi(v-\Delta, \Delta)}_{\text{jump probability}} d\Delta$$

L.H.S.

$$f(v, t) + \tau \frac{\partial f}{\partial t} + \frac{\tau^2}{2} \frac{\partial^2 f}{\partial t^2} + \dots$$

R.H.S (Integrand)

$$= \left(f(v, t) - \Delta \frac{\partial f}{\partial v} + \frac{\Delta^2}{2} \frac{\partial^2 f}{\partial v^2} \right) \left(\phi(v, \Delta) - \Delta \frac{\partial \phi}{\partial v} + \frac{\Delta^2}{2} \frac{\partial^2 \phi}{\partial v^2} \dots \right)$$

$\Delta \equiv$ has the dimension of velocity

$$\int_{-\infty}^{+\infty} \phi(v, \Delta) d\Delta = 1$$

$$= f(v, t) \phi(v, \Delta) - \Delta \frac{\partial f}{\partial v} \phi(v, \Delta) + \frac{\Delta^2}{2} \frac{\partial^2 f}{\partial v^2} \phi(v, \Delta) \dots$$

$$R.H.S = f(v, t) \int_{-\infty}^{+\infty} \phi(v, \Delta) d\Delta - \frac{\partial}{\partial v} \left[f(v, t) \int_{-\infty}^{+\infty} \Delta \phi(v, \Delta) d\Delta \right]$$

$$+ \frac{\partial^2}{\partial v^2} \left[f(v, t) \int_{-\infty}^{+\infty} \frac{\Delta^2}{2} \phi(v, \Delta) d\Delta \right]$$

$$\frac{\partial f}{\partial t} = - \frac{\partial}{\partial v} \left[f(v, t) \underbrace{\int_{-\infty}^{+\infty} \Delta \phi(v, \Delta) d\Delta}_{M_1(v) \equiv \frac{\langle \Delta \rangle}{\tau}} \right] + \frac{\partial^2}{\partial v^2} \left[f(v, t) \underbrace{\int_{-\infty}^{+\infty} \frac{\Delta^2}{2} \phi(v, \Delta) d\Delta}_{M_2(v)} \right]$$

$$\frac{\partial f(v,t)}{\partial t} = - \underbrace{\frac{\partial}{\partial v} M_1(v) f(v,t)}_{\text{Drift term}} + \underbrace{\frac{\partial^2}{\partial v^2} M_2(v) f(v,t)}_{\text{Diffusion term (in velocity)}} \rightarrow \text{Fokker-Planck equation.}$$

Calculation of $M_1(v)$

$$\frac{dv}{dt} = -\Gamma v + \frac{F(t)}{m}$$

$$\Rightarrow dv = -\Gamma v dt + \frac{F(t)}{m} dt$$

Integrating betⁿ t and $t+\tau$

$$(v(t+\tau) - v(t)) = -\Gamma v \tau + \frac{1}{m} \int_t^{t+\tau} F(t') dt'$$

$$\Delta = -\Gamma v \tau + \frac{1}{m} \int_t^{t+\tau} F(t') dt'$$

Take average $\langle \Delta \rangle = -\Gamma v \tau$

$$\langle F(t) \rangle = 0$$

$$\langle \Delta \rangle = -\Gamma v \tau$$

$$\frac{\langle \Delta \rangle}{\tau} = -\Gamma v$$

↳ Drift term

Calculation of $M_2(v)$:

$$\Delta = -\Gamma v \tau + \frac{1}{m} \int_t^{t+\tau} F(t') dt'$$

$$\Delta^2 = +\Gamma^2 v^2 \tau^2 - \frac{2\Gamma v \tau}{m} \int_t^{t+\tau} F(t') dt' + G_t^2(\tau)$$

$$G_t(\tau) = \frac{1}{m} \int_t^{t+\tau} F(t') dt'$$

τ is small, hence the first term can be neglected, $\langle F(t') \rangle = 0$

$$\langle \Delta^2 \rangle = \frac{1}{m^2} \int_t^{t+\tau} dt' \int_t^{t+\tau} \langle F(t') F(t'') \rangle dt''$$

$$\begin{aligned}
 \langle \Delta^2 \rangle &= \frac{1}{m^2} \int_t^{t+\tau} dt' \int_t^{t+\tau} dt'' \langle F(t') F(t'') \rangle \\
 &= \frac{1}{m^2} 2C' \int_t^{t+\tau} dt' \int_t^{t+\tau} dt'' \delta(t' - t'')
 \end{aligned}$$

$$\langle \Delta^2 \rangle = \frac{2C'\tau}{m^2}$$

$$\frac{\langle \Delta^2 \rangle}{2\tau} = \frac{C'}{m^2} \leftarrow \text{Diffusion coefficient in the velocity space}$$

$$\frac{dv}{dt} = -\Gamma v + \frac{1}{m} F(t)$$

Integrate using integrating factor

$$v(t) = v(0) e^{-\Gamma t} + \frac{e^{-\Gamma t}}{m} \int_0^t e^{\Gamma t'} F(t') dt'$$

$$\begin{aligned} \langle (v(t) - v(0) e^{-\Gamma t})^2 \rangle &= \frac{e^{-2\Gamma t}}{m^2} \int_0^t dt' \int_0^t dt'' e^{\Gamma(t'+t'')} \langle F(t') F(t'') \rangle \\ &= \frac{e^{-2\Gamma t}}{m^2} \int_0^t dt' \int_0^t dt'' e^{\Gamma(t'+t'')} 2c' \delta(t' - t'') \\ &= \frac{2c'}{m^2} e^{-2\Gamma t} \int_0^t e^{2\Gamma t'} dt' = \frac{2c'}{m^2} e^{-2\Gamma t} \left. \frac{e^{2\Gamma t'}}{2\Gamma} \right|_0^t \\ &= \frac{2c'}{2m^2} e^{-2\Gamma t} (e^{2\Gamma t} - 1) = \frac{c'}{m^2 \Gamma} (1 - e^{-2\Gamma t}) \end{aligned}$$

$$\langle (v(t) - v(0)e^{-\Gamma t})^2 \rangle$$

$$= \frac{c'}{\Gamma m^2} (1 - e^{-2\Gamma t})$$

for $t \gg \frac{1}{\Gamma}$ (long time limit)

then $\langle v(t)^2 \rangle = \frac{c'}{\Gamma m^2}$

$$\frac{c'}{m^2} = \Gamma \langle v(t)^2 \rangle = \Gamma \frac{k_B T}{m}$$

$$\hookrightarrow \frac{\langle \Delta^2 \rangle}{2\tau} = M_2(v) = \frac{\Gamma k_B T}{m}$$

$$M_1(v) = -\Gamma v$$

Now, the Fokker-Planck equation reads

$$\frac{\partial f(v, t)}{\partial t} = \frac{\partial}{\partial v} [\Gamma v] f(v, t) + \frac{\partial^2}{\partial v^2} \left[\frac{\Gamma k_B T}{m} \right] f(v, t)$$

Brownian Motion in phase Space: Under a force field

$q \rightarrow$ position $p \rightarrow$ momentum

$$q_1 = q + (p/m)\tau ; \quad p_1 = p + K\tau$$

$$K = -\frac{\partial V}{\partial q} = V'(q)$$

\downarrow force \searrow Derivative of a potential energy

$$f(p + K\tau, q + (p/m)\tau, t + \tau) = \int_{-\infty}^{+\infty} f(p - \Delta, q, t) \phi(p - \Delta, q, \Delta) d\Delta$$

$$\text{L.H.S} = f(p, q, t) + \frac{\partial f}{\partial p} (K\tau) + \frac{\partial f}{\partial q} (p/m)\tau + \dots \frac{\partial f}{\partial t} \tau$$

$$\text{R.H.S} = \int_{-\infty}^{+\infty} \left(f(p, q, t) - \Delta \frac{\partial f}{\partial p} + \frac{\Delta^2}{2} \frac{\partial^2 f}{\partial p^2} \dots \right) \left(\phi(p, q, \Delta) \dots \right) d\Delta$$

R.H.S

$$= f(p, q, t) \underbrace{\int_{-\infty}^{+\infty} \phi(p, q, \Delta) d\Delta}_1 - \frac{\partial}{\partial p} [f] \int_{-\infty}^{+\infty} \Delta \phi(p, q, \Delta) d\Delta + \frac{\partial^2}{\partial p^2} [f] \int_{-\infty}^{+\infty} \frac{\Delta^2}{2} \phi(p, q, \Delta) d\Delta$$

$$\cancel{f(p, q, t)} + \frac{\partial f}{\partial p} \cancel{\tau} + \frac{\partial f}{\partial q} (\cancel{p/m}) \tau + \frac{\partial f}{\partial t} \tau$$

$$= \cancel{f(p, q, t)} - \dots$$

$$\frac{\partial f}{\partial p} \tau + \frac{\partial f}{\partial q} (p/m) + \frac{\partial f}{\partial t} = - \frac{\partial}{\partial p} [f] \underbrace{\left[\frac{1}{\tau} \int_{-\infty}^{+\infty} \Delta \phi(p, q, \Delta) d\Delta \right]}_{M_1(p, q)} + \frac{\partial^2}{\partial p^2} [f] \underbrace{\left[\frac{1}{\tau} \int_{-\infty}^{+\infty} \frac{\Delta^2}{2} \phi(p, q, \Delta) d\Delta \right]}_{M_2(p, q)}$$

$$m=1, \quad q=x, \quad p=v$$

$$\frac{\partial f(x, v, t)}{\partial t} = - \frac{\partial f}{\partial x} v + \frac{\partial f}{\partial v} v'(x) - \frac{\partial}{\partial v} \left[\gamma v f(x, v, t) \right] + \gamma k_B T \frac{\partial^2 f}{\partial v^2} \quad \leftarrow \text{Kramers' equation}$$

\parallel
 $= -\gamma v$

\downarrow
 $= \gamma \frac{k_B T}{m}$

$$* = -\frac{dv}{dq} = -v'(q)$$

$$\frac{\partial f(x, v, t)}{\partial t} = -\frac{\partial f}{\partial x} v + \frac{\partial f}{\partial v} v'(x) - \frac{\partial}{\partial v} [\gamma v f(x, v, t)] + \gamma k_B T \frac{\partial^2 f(x, v, t)}{\partial v^2}$$

Corresponds
to the Liouville's
eqⁿ

$$\frac{\partial f}{\partial t} = 0 \rightarrow \text{Equilibrium/Steady state}$$

$$f(x, v) = \frac{1}{Z} e^{-\frac{(\frac{1}{2}v^2 + V(x))}{k_B T}}$$

$$= \frac{1}{Z} e^{-\beta H}$$

→ one can check

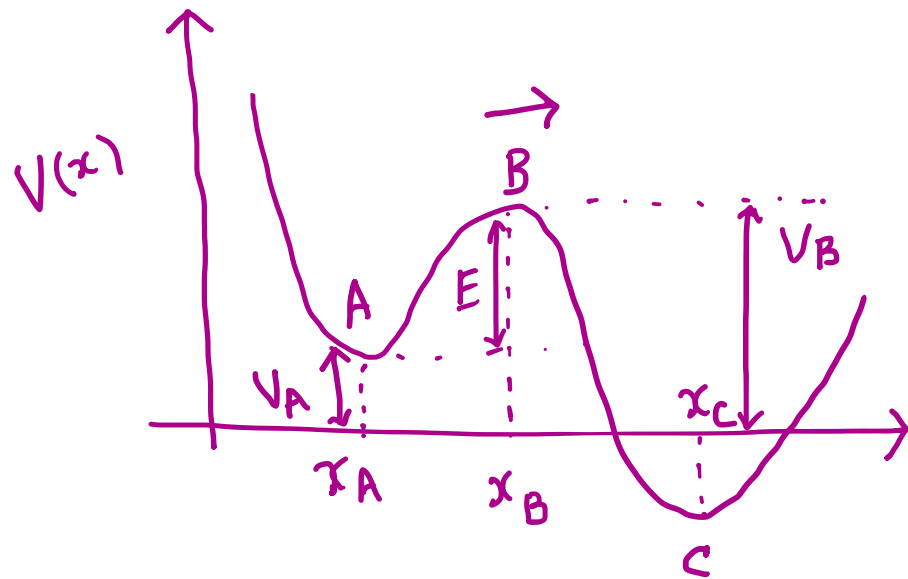
Maxwell.

(Boltzmann Distribution)

$$H = \frac{v^2}{2} + V(x)$$

Kramers' Theory of activated processes:

"Chemical Reaction"



$$E = V_B - V_A$$

↓
Activation
energy

$x_A \rightarrow$ Reactant State

$x_C \rightarrow$ Product State

$x_B \rightarrow$ Barrier top

x (Reaction Co-ordinate)

We want to calculate the Steady State rate from A to C

* Stationary Situation \rightarrow Steady state current
at the barrier top

$$f(x, v, t) \Rightarrow P(x, v, t)$$

$$\frac{\partial P}{\partial t} = 0$$

$$\left[-\frac{\partial}{\partial x} v + \frac{\partial}{\partial x} \{ V'(x) + \gamma v \} + \gamma k_B T \frac{\partial^2}{\partial v^2} \right] P(x, v) = 0$$

"inverted parabola"

At the Barrier top

$$V(x) = V(x_B) + \underbrace{\frac{\partial V}{\partial x}}_0 \bigg|_{x=x_B} (x-x_B) + \frac{1}{2} \frac{\partial^2 V}{\partial x^2} \bigg|_{x=x_B} (x-x_B)^2 + \dots$$

$$V(x) = V(x_B) - \frac{1}{2} \omega_B^2 (x-x_B)^2$$

Near the bottom well, we have equilibrium.

$$P(x, v) = \frac{1}{Z} e^{-\left(\frac{1}{2} v^2 + V(x)\right) / k_B T} \quad \text{at } x \approx x_A$$

$$V(x) = V(x_A) + \underbrace{\frac{\partial V}{\partial x}}_0 \bigg|_{x=x_A} (x-x_A) + \frac{1}{2} \frac{\partial^2 V}{\partial x^2} \bigg|_{x=x_A} (x-x_A)^2$$

$$V(x) = V(x_A) + \frac{1}{2} \omega_A^2 (x-x_A)^2$$

Population in the well A

$$n_a = \int_{-\infty}^{+\infty} dx \int_{-\infty}^{+\infty} dv P(x, v)$$

Flux j over the barrier

$$j = \int_{-\infty}^{+\infty} v P(x, v) dv$$

The steady state Kramers' rate
is given by $k_{A \rightarrow C} = j/n_a$

$$\frac{\text{flux}}{\text{population}} = \text{Steady state rate}$$

Calculation of j :

We construct

$$P(x, v) = \xi(x, v) \exp \left[-\frac{\left(\frac{1}{2} v^2 + V(x) \right)}{k_B T} \right]$$

After a little bit of algebra

$$\left[-v \frac{\partial}{\partial x} - \left\{ \omega_B^2 (x - x_B) + \gamma v \right\} \frac{\partial}{\partial v} + \gamma k_B T \frac{\partial^2}{\partial v^2} \right] \zeta(x, v) = 0$$

$$\zeta(x, v) = 1 \quad \text{inside the well } x \sim x_A$$

$$\zeta(x, v) = 0 \quad \text{beyond the barrier top B, } x > x_B$$

We use the following linear transformation

$$u = v + (x - x_B) a \quad a \text{ is a constant to be determined later}$$

$$\frac{\partial}{\partial x} = a \frac{\partial}{\partial u} ; \quad \frac{\partial}{\partial v} = \frac{\partial}{\partial u}$$

$$\left[-va \frac{\partial}{\partial u} - \left\{ \omega_B^2 (x - x_B) + \gamma v \right\} \frac{\partial}{\partial u} + \gamma k_B T \frac{\partial^2}{\partial u^2} \right] \zeta(x, u) = 0$$

$$\gamma k_B T \frac{\partial^2 \xi}{\partial u^2} - [\omega_B^2 (x - x_B) + \gamma v + a v] \frac{\partial \xi}{\partial u} = 0$$

Let $\omega_B^2 (x - x_B) + v(\gamma + a) = -\lambda u$ $\lambda = \text{constant (to be determined)}$

Since, $u = v + a(x - x_B)$

$$\omega_B^2 (x - x_B) + v(\gamma + a) = -\lambda v - \lambda a (x - x_B)$$

$$-\lambda a = \omega_B^2$$

$$a + \gamma = -\lambda$$

$$a^2 + \gamma a = -\lambda a$$

$$\underline{a^2 + \gamma a - \omega_B^2 = 0}$$

$$\left| \begin{array}{l} ax^2 + bx + c = 0 \\ x = \frac{-b \pm \sqrt{b^2 - 4ac}}{2a} \end{array} \right.$$

$$a = \frac{-\gamma \pm \sqrt{\gamma^2 + 4\omega_B^2}}{2}$$

$$= -\frac{\gamma}{2} \pm \sqrt{\left(\frac{\gamma}{2}\right)^2 + \omega_B^2}$$

$$\gamma k_B T \frac{\partial^2 \xi}{\partial u^2} - [\omega_B^2 (x - x_B) + V(a + r)] \frac{\partial \xi}{\partial u} = 0$$

$$\gamma k_B T \frac{\partial^2 \xi}{\partial u^2} + \lambda u \frac{\partial \xi}{\partial u} = 0$$

$$\text{Let } \frac{\partial \xi}{\partial u} = \gamma$$

$$\gamma k_B T \frac{\partial \gamma}{\partial u} = -\lambda u \gamma$$

$$\Rightarrow \int \frac{d\gamma}{\gamma} = -\frac{\lambda}{\gamma k_B T} \int du u$$

$$\ln \gamma = -\frac{\lambda}{2\gamma k_B T} u^2 + \ln F_2 \leftarrow \text{constant of integration}$$

$$\gamma = F_2 \exp\left(-\frac{\lambda u^2}{2\gamma k_B T}\right)$$

$$\xi(u) = F_2 \int_0^u \exp\left(-\frac{\lambda u^2}{2\gamma k_B T}\right) du + F_1$$

We look for a solⁿ that vanishes as $x \rightarrow \infty$, but the above integral remain finite for all $|u| \rightarrow \infty$

This implies $\lambda > 0$

$\lambda = -\frac{\omega_B^2}{a} \Rightarrow$ Negative root of a should be chosen

$$a = -\frac{\gamma}{2} - \sqrt{(\gamma/2)^2 + \omega_B^2} \quad (\text{negative root})$$

Now when $x \rightarrow \infty$ $u \rightarrow -\infty$

$$u = v + a(x - x_B)$$

$$x \rightarrow \infty \quad \zeta(x, v) \rightarrow 0$$

$$0 = F_1 + F_2 \int_0^{-\infty} \exp\left(-\frac{\lambda u^2}{2\gamma k_B T}\right) du$$

$$F_1 = F_2 \int_{-\infty}^0 \exp\left(-\frac{\lambda u^2}{2\gamma k_B T}\right) du = \frac{F_2}{2} \int_{-\infty}^{+\infty} \exp\left(-\frac{\lambda u^2}{2\gamma k_B T}\right) du = \frac{F_2}{2} \left(\frac{2\pi \gamma k_B T}{\lambda} \right)^{1/2}$$

$$= F_2 \left(\frac{\pi \gamma k_B T}{2\lambda} \right)^{1/2}$$

$$\zeta(u) = F_2 \left[\sqrt{\frac{\pi \gamma k_B T}{2\lambda}} + \int_0^u \exp\left(-\frac{\lambda u^2}{2\gamma k_B T}\right) du \right]$$

$$\left[\int_{-\infty}^{+\infty} e^{-\alpha x^2} = \left(\frac{\pi}{\alpha} \right)^{1/2} \right]$$

$$P(x, v) = F_2 \left[\sqrt{\frac{\pi \gamma k_B T}{2\lambda}} + \int_0^u \exp\left(-\frac{\lambda u^2}{2\gamma k_B T}\right) du \right] \times \exp\left(-\frac{(\frac{1}{2}v^2 + V(x_B) - \frac{\omega_B^2}{2}(x-x_B)^2)}{k_B T}\right)$$

then at $x \approx x_B$

$$C = \sqrt{\frac{\pi \gamma k_B T}{2\lambda}}$$

$$= F_2 \exp\left(-\frac{V(x_B)}{k_B T}\right) \left[C \exp\left(-\frac{v^2}{2k_B T}\right) + F(x, v) \exp\left(-\frac{v^2}{2k_B T}\right) \right]$$

$$F(x, v) = \int_0^v \exp\left(-\frac{\lambda u^2}{2k_B T}\right) du$$

$$j = \int_{-\infty}^{+\infty} v P(x_B, v) dv$$

$$\int_{-\infty}^{+\infty} v e^{-v^2/2k_B T} dv = 0$$

$$[\int f g = f g - \int g df]$$

$$j = F_2 \exp\left(-\frac{V(x_B)}{k_B T}\right) \left[\int_{-\infty}^{+\infty} v e^{-\frac{v^2}{2k_B T}} F(x_B, v) dv \right]$$

$$= F_2 \exp\left(-\frac{V(x_B)}{k_B T}\right) (-k_B T) \int_{-\infty}^{+\infty} \left(\frac{\partial}{\partial v} e^{-v^2/2k_B T} \right) F(x_B, v) dv$$

$$= F_2 \exp\left(-\frac{V(x_B)}{k_B T}\right) (-k_B T) \left[F(x_B, v) e^{-v^2/2k_B T} \Big|_{-\infty}^{+\infty} - \int_{-\infty}^{+\infty} e^{-v^2/2k_B T} \frac{\partial F(x_B, v)}{\partial v} dv \right]$$

$$j = F_2 \exp\left(-\frac{V(x_B)}{k_B T}\right) (k_B T) \int_{-\infty}^{+\infty} \left(\frac{\partial F}{\partial v}\right) e^{-v^2/2k_B T} dv$$

$$F(x, v) = \int_0^v \exp\left(-\frac{\lambda u^2}{2\gamma k_B T}\right) du \quad \left| \quad j = F_2 \exp\left(-\frac{V(x_B)}{k_B T}\right) (k_B T) \int_{-\infty}^{+\infty} \exp\left[-\frac{v^2}{2k_B T} \left(\frac{\lambda}{\gamma} + 1\right)\right] dv\right.$$

then $\frac{\partial F}{\partial v} = \exp\left(-\frac{\lambda v^2}{2\gamma k_B T}\right)$

$$j = F_2 \exp\left(-\frac{V(x_B)}{k_B T}\right) (k_B T) \frac{\pi^{1/2}}{\sqrt{\lambda/2k_B T \gamma + \frac{1}{2k_B T}}}$$

$$j = F_2 \exp\left(-\frac{V(x_B)}{k_B T}\right) (k_B T) \sqrt{2\pi k_B T} \left(\frac{\gamma}{\lambda + \gamma}\right)^{1/2}$$

Calculation of population n_a

$$n_a = \int_{-\infty}^{+\infty} dx \int_{-\infty}^{+\infty} dv P(x, v)$$

$$V(x) = V(x_A) + \frac{\omega_A^2}{2} (x - x_A)^2$$

$$P(x, v) = \xi(x, v) \exp \left[- \frac{(\frac{1}{2} v^2 + V(x))}{k_B T} \right]$$

$$\xi(x, v) = F_2 \left[\left(\frac{\pi \gamma k_B T}{2 \lambda} \right)^{\frac{1}{2}} + \int_0^u e^{-\frac{du^2}{2 \gamma k_B T}} du \right]$$

$$= F_2 2 \left(\frac{\pi \gamma k_B T}{2 \lambda} \right)^{\frac{1}{2}} = F_2 \left(\frac{2 \pi \gamma k_B T}{\lambda} \right)^{\frac{1}{2}}$$

$$n_a = F_2 \left(\frac{2 \pi \gamma k_B T}{\lambda} \right)^{\frac{1}{2}} \exp \left(- \frac{V(x_A)}{k_B T} \right) \int_{-\infty}^{+\infty} e^{-v^2 / 2 k_B T} dv \int_{-\infty}^{+\infty} e^{-\frac{1}{2 k_B T} \omega_A^2 (x - x_A)^2} dx$$

$$n_a = F_2 \left(\frac{2\pi\gamma k_B T}{\lambda} \right)^{1/2} \exp\left(-\frac{V(x_A)}{k_B T}\right) (2\pi k_B T)^{1/2} \left(\frac{2\pi k_B T}{\omega_A^2} \right)^{1/2}$$

$$n_a = F_2 \left(\frac{2\pi k_B T}{\omega_A} \right)^{3/2} \left(\frac{\gamma}{\lambda} \right)^{1/2} \exp\left(-\frac{V(x_A)}{k_B T}\right)$$

$$j = F_2 e^{-V(x_B)/k_B T} (k_B T 2\pi)^{3/2} \left(\frac{\gamma}{\lambda + \gamma} \right)^{1/2} \frac{1}{2\pi}$$

$$k_{A \rightarrow C} = \frac{j}{n_a} = \left(\frac{\gamma}{\lambda + \gamma} \right)^{1/2} \left(\frac{\omega_A}{2\pi} \right) \exp\left(-\frac{V(x_B) - V(x_A)}{k_B T}\right)$$

$$n_a = F_2 \left(\frac{2\pi\gamma k_B T}{\lambda} \right)^{1/2} \exp\left(-\frac{V(x_A)}{k_B T}\right) (2\pi k_B T)^{1/2} \left(\frac{2\pi k_B T}{\omega_A^2} \right)^{1/2}$$

$$n_a = F_2 \left(\frac{2\pi k_B T}{\omega_A} \right)^{3/2} \left(\frac{\gamma}{\lambda} \right)^{1/2} \exp\left(-\frac{V(x_A)}{k_B T}\right)$$

$$j = F_2 e^{-V(x_B)/k_B T} (k_B T 2\pi)^{3/2} \left(\frac{\gamma}{\lambda + \gamma} \right)^{1/2} \frac{1}{2\pi}$$

$$k_{A \rightarrow C} = \frac{j}{n_a} = \left(\frac{\gamma}{\lambda + \gamma} \right)^{1/2} \left(\frac{\omega_A}{2\pi} \right) \exp\left(-\frac{V(x_B) - V(x_A)}{k_B T}\right)$$

$$-\lambda = a + \gamma$$

↪ We should take the negative root a_-

$$V(x_B) - V(x_A) = E$$

↓
Activation energy

$$\lambda = -a_- - \gamma$$

$$= -(a_- + \gamma)$$

$$\lambda = -\left[\gamma + \left\{\frac{\gamma}{2} - \sqrt{\left(\frac{\gamma}{2}\right)^2 + \omega_B^2}\right\}\right] = -\frac{\gamma}{2} + \sqrt{\left(\frac{\gamma}{2}\right)^2 + \omega_B^2} = a_+$$

$$\frac{\lambda}{\lambda + \gamma} = \frac{a_+}{-a_-} = -\frac{a_+ a_+}{a_- a_+}$$

$$= -\frac{\left(-\frac{\gamma}{2} + \sqrt{\left(\frac{\gamma}{2}\right)^2 + \omega_B^2}\right)^2}{\left\{\left(-\frac{\gamma}{2}\right)^2 - \left(\frac{\gamma}{2}\right)^2 - \omega_B^2\right\}} = \frac{1}{\omega_B^2} \left(-\frac{\gamma}{2} + \sqrt{\left(\frac{\gamma}{2}\right)^2 + \omega_B^2}\right)^2$$

Thus the final expression
for the kramers rate is

We know for 1st order rate constant

$$k = \left(\frac{\omega_A}{2\pi\omega_B}\right) \left(-\frac{\gamma}{2} + \sqrt{\left(\frac{\gamma}{2}\right)^2 + \omega_B^2}\right) \exp\left[-E/k_B T\right]$$

$$k = A e^{-E/k_B T}$$

Arrhenius
relation

One can compare $A \rightarrow (\omega_A, \omega_B, \gamma)$.

$$k = \left(\frac{\omega_A}{2\pi\omega_B} \right) \left(-\gamma/2 + \sqrt{(\gamma/2)^2 + \omega_B^2} \right) e^{-E/k_B T}$$

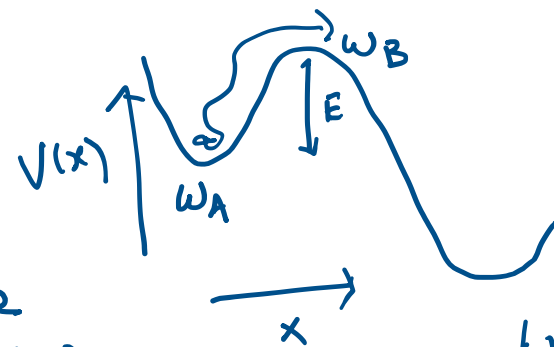
Now, $\gamma \rightarrow 0$ (low friction / small viscosity)

$$k = \left(\frac{\omega_A}{2\pi} \right) e^{-E/k_B T} = k_{TST} \quad (\text{Transition State theory rate})$$

$$k = k_{TST} \kappa \quad \leftarrow \text{Transmission coefficient}$$

$$\kappa = \frac{1}{\omega_B} \left(-\gamma/2 + \sqrt{(\gamma/2)^2 + \omega_B^2} \right)$$

varies between 0 and 1



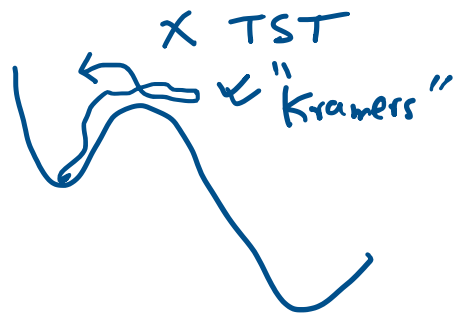
$$k_{TST} = \left(\frac{\omega_A}{2\pi} \right) e^{-E/k_B T}$$

"Upper bound to the rate"

(No dissipation)

No-recrossing

Maximum possible rate



When $\gamma \rightarrow \text{large}$

$\gamma \gg \omega_B$ (large viscosity limit)

$$k_{\gamma \rightarrow \text{large}} = \frac{\omega_A}{2\pi} e^{-E/k_B T} \frac{1}{\omega_B} \left[-\gamma/2 + \sqrt{(\gamma/2)^2 + \omega_B^2} \right]$$

$$\frac{1}{\omega_B} \left[-\gamma/2 + (\gamma/2) \sqrt{1 + \omega_B^2 \cdot \frac{4}{\gamma^2}} \right]$$

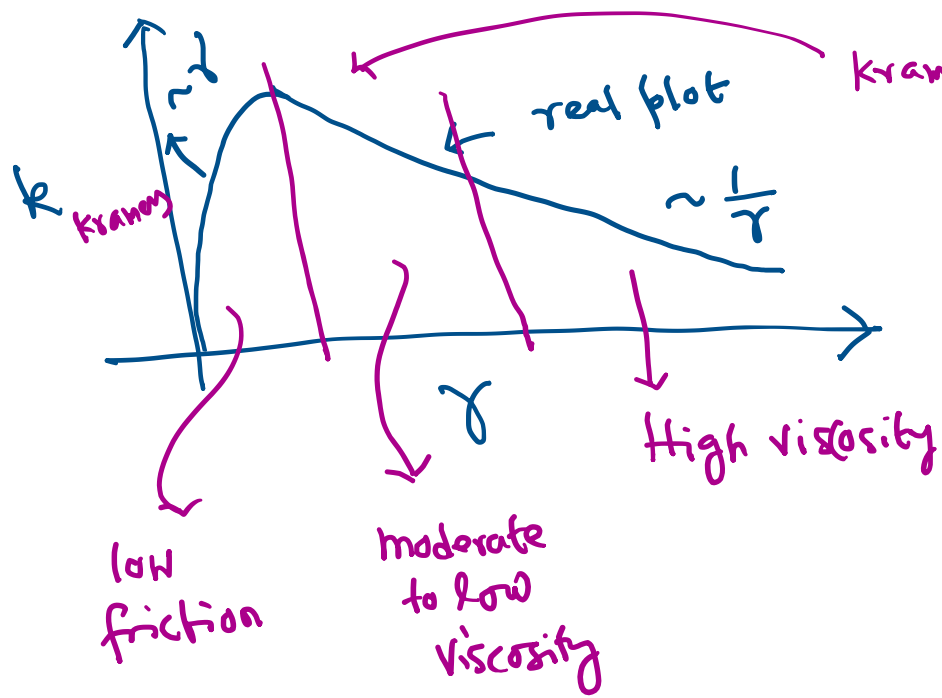
$\gamma \rightarrow \text{large}$

$$\frac{1}{\omega_B} \left[-\gamma/2 + \frac{\gamma}{2} \left(1 + \frac{1}{2} \omega_B^2 \frac{4}{\gamma^2} \right) \right]$$

$$\frac{1}{\omega_B} \left[-\gamma/2 + \gamma/2 + \frac{\gamma}{4} \omega_B^2 \frac{4}{\gamma^2} \right] = \frac{1}{\omega_B} \frac{\omega_B^2}{\gamma} = \frac{\omega_B}{\gamma}$$

$$k_{\gamma \rightarrow \text{large}} = \frac{\omega_A \omega_B}{2\pi \gamma} e^{-E/k_B T}$$

$$\rightarrow \text{rate} \sim \frac{1}{\gamma} \sim \frac{1}{\text{viscosity}}$$



Kramers turn over

→ Peter Hänggi

50 years after Kramers

RMP 1990

→ Eli Pollack → see his papers.

Quantum Kramers' problem → D.S. Ray (IACS).